

# **Progress Report**

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On behalf of the Chandra Source Catalog Project Team

Chandra Users' Committee Meeting April 6, 2009



### **Executive Summary**

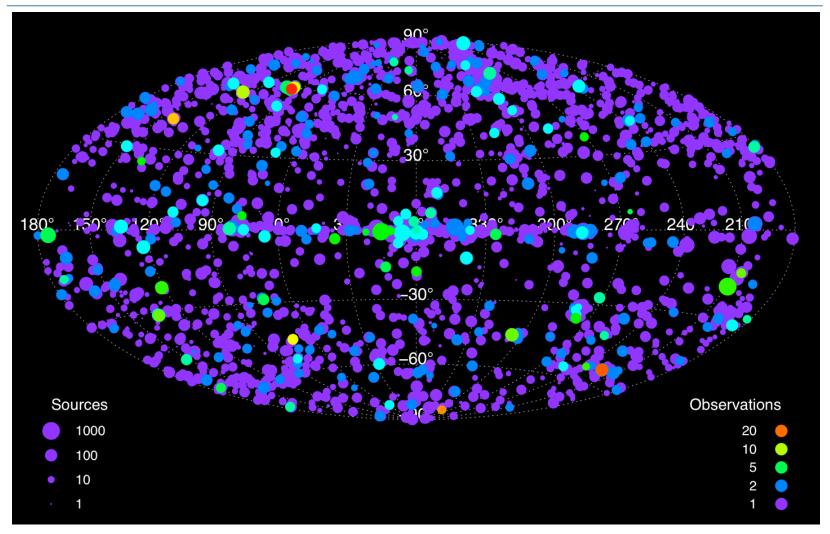
#### **Summary**

- Released version 1 of the Chandra Source Catalog on March 4, 2009
  - This release of the catalog includes
     94,676 master Chandra sources (unique sources on the sky)
     135,914 individual source observations (detected-source regions)
     3,921 observation intervals
- Release includes point and compact sources detected in a subset of ACIS imaging observations released publicly prior to 2009
  - Only sources with observed spatial extents <~30 arcseconds are included</li>
  - Observations with highly extended sources are either excluded, or a subset of the active CCDs are included if the extended source is restricted to a single chip
  - Multiple observations of the same field are not co-added prior to source detection
  - Future catalog releases will relax these restrictions
- The release includes an extensive statistical characterization of the derived source properties included in the catalog



## Chandra Source Catalog Observations





- The locations of observations included in the CSC, in Galactic coordinates
  - The size of each symbol is proportional to the logarithm of the number of sources detected in the field, while the color encodes the number of closely-located observations





### **Science Highlights Since Last CUC Meeting**

- Completed catalog production and release 1, including statistical characterization release 1 liens
- Updated public web site with latest user documentation and threads
  - http://cxc.cfa.harvard.edu/csc/
- Documents and publications delivered
  - "Chandra Source Catalog Requirements version 1.0"
  - "Statistical Characterization of the Chandra Source Catalog, Release 1"
  - Posters presented at the January AAS
    - "The Chandra Source Catalog" I. N. Evans et al.
    - "The Chandra Source Catalog: User Interface" N. R. Bonaventura et al.
    - "The Chandra Source Catalog: Source Properties and Data Products" A. H. Rots et al.
    - "The Chandra Source Catalog: Spectral Properties" S. M. Doe et al.
    - "The Chandra Source Catalog: Statistical Characterization" F. A. Primini et al.
    - "The Chandra Source Catalog: Algorithms" J. C. McDowell et al.
    - "The Chandra Source Catalog: Processing and Infrastructure" J. D. Evans et al.
    - "The Chandra Source Catalog: Automated Source Correlation" R. Hain et al.
    - "The Chandra Source Catalog: X-ray Aperture Photometry" V. L. Kashyap et al.
    - "The Chandra Source Catalog: Background Determination and Source Detection"
      - M. L. McCollough et al.
    - "The Chandra Source Catalog: Source Variability" M. Nowak et al.
  - Catalog paper (Evans et al.) and statistical characterization paper (Primini et al.) currently in advanced stages of preparation



## Software Highlights

### **Software Highlights Since Last CUC Meeting**

- CAT 3.0 (Production system) build
  - Minor patches to maintain CSC operations

CAT 3.0.3	Nov 17	ACIS 1-chip filtering / QA update / few bug-fixes
CAT 3.0.4	Dec 04	Master pipeline bug-fix (no science changes)
CAT 3.0.5	Dec 16	Archive server upgrade; added 2008 leap second
CAT 3.0.6	Dec 22	Master pipeline/Repro bug-fix
CAT 3.0.7	Jan 26	Merge Review GUI bug-fix
CAT 3.0.8	Feb 09	Master Pipeline Repro support through Merge Review GUI
CAT 3.0.9	Feb 19	Aper_90 per obi source bug-fix
CAT 3.0.10	Feb 25	Release 1 migration tasks

- CAT 3.1 (Release system) build
  - CSCview bug-fixes and enhancements
- Software team ran catalog pipeline processing operations for release 1 production and statistical characterization simulation runs



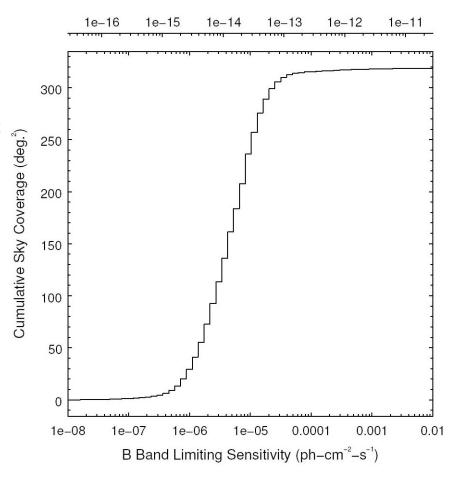
## Catalog Statistical Characterization

### **Catalog Statistical Characterization**

- Statistical characterization is a scientifically essential component of the CSC
  - Characterization determines statistically how well the science algorithms used in catalog construction actually performed
  - Statistical characterization provides the user with the information to judge whether a particular line of enquiry is feasible using the catalog release data
- Statistical characterization of catalog source properties is accomplished primarily by processing simulated datasets through the catalog pipelines
- Characterization is an ongoing, evolving process
  - The results provided with the catalog release give a good overview of the general catalog properties
  - Some specific questions identified during production or with existing characterization runs have still to be addressed
    - E.g.: source detection efficiency in the vicinity of the bright, crowded cores of galaxies
  - Feedback from users will identify areas where existing characterization should be refined
- Statistical characterization results are published on the catalog web site
  - http://cxc.cfa.harvard.edu/csc/char.html

#### **Sky Coverage**

- The sky coverage represents the total area in the CSC sensitive to point sources greater than a given flux, as a function of flux
  - We estimate sky coverage by assigning all non-zero limiting sensitivity map values to all-sky pixels, keeping only the most sensitive value in each all-sky pixel
- Approximate sky coverage
  - $\sim 300 \text{ deg}^2 \text{ at } 2 \times 10^{-5} \text{ photons cm}^{-2} \text{ s}^{-1}$  $(\sim 5 \times 10^{-14} \text{ ergs cm}^{-2} \text{ s}^{-1})$
  - $\sim$ 70 deg<sup>2</sup> at 2×10<sup>-6</sup> photons cm<sup>-2</sup> s<sup>-1</sup> ( $\sim 5\times10^{-15}$  ergs cm<sup>-2</sup> s<sup>-1</sup>)
  - $\sim$ 6 deg<sup>2</sup> at 4×10<sup>-7</sup> photons cm<sup>-2</sup> s<sup>-1</sup> ( $\sim$  1×10<sup>-15</sup> ergs cm<sup>-2</sup> s<sup>-1</sup>)



Total ACIS "b" band sky coverage is ~320 deg<sup>2</sup>

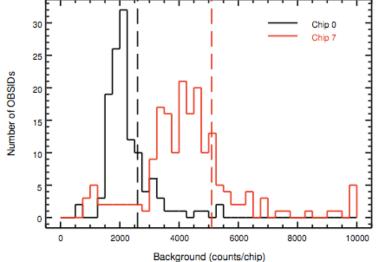


## Catalog Properties: Field Background

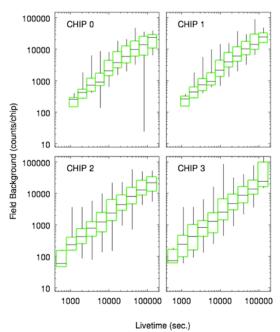
#### CXC

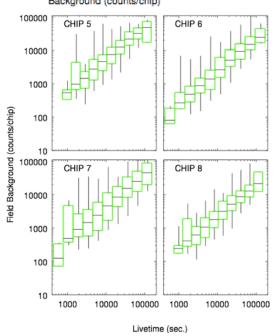
#### Field Background

- Top panel: distribution of chip 0 and chip 7 ACIS "b" band background estimated from CSC 10 ks event lists with sources removed
  - Dashed vertical lines indicate the values cited in version 11 of the POG
  - Event screening performed in the CSC pipeline processing is more aggressive than that done in standard data processing, so the non-X-ray background is typically reduced



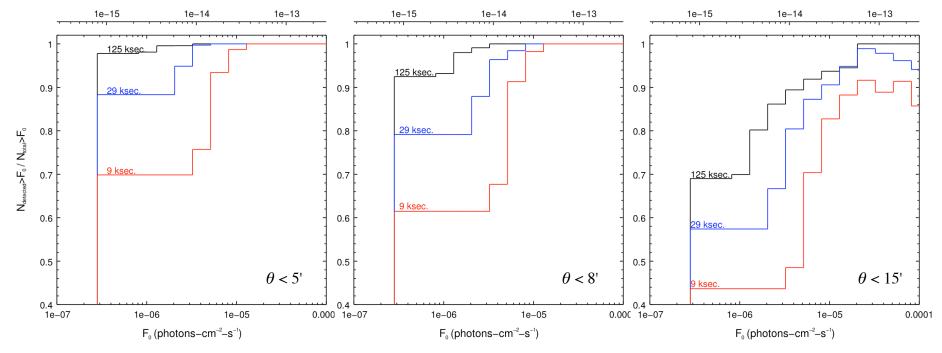
- Bottom panel: field background per chip vs. livetime
  - Median background counts per bin are indicated by horizontal lines
  - Boxes include 95% of the measurements in each bin, and vertical lines indicate extreme values





### **Source Detection Efficiency**

- Estimated from fraction of simulated sources of a given flux actually detected
  - Simulated absorbed ( $N_{\rm H} = 3 \times 10^{20} \, {\rm cm}^{-2}$ ) power-law ( $E^{-1.7}$ ) and blackbody ( $kT = 1.0 \, {\rm keV}$ ) spectra, with a power-law N > S distribution with index 1.5



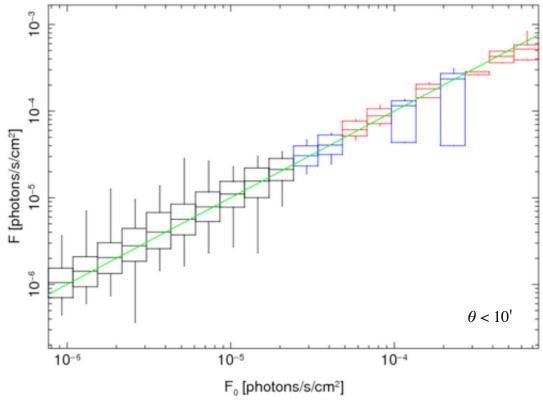
- High flux end suffers from low number statistics
- Reduced efficiencies for the  $\theta$  < 15' curves are real and indicate reduced sensitivity at large values of  $\theta$ because of the increased aperture background



## Catalog Properties: Flux Accuracy

### Flux Accuracy

- Estimated from comparison of the input and measured fluxes of the simulated source
- Results indicate good agreement for sources within 10' of the aim point
  - For sources beyond 10', there appears to systematic overestimate of a factor of  $\sim$ 2 for sources fainter than  $\sim$ 3 x 10<sup>-6</sup> photons cm<sup>-2</sup> s<sup>-1</sup>



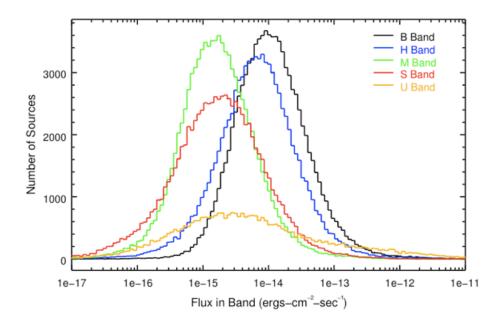
- Comparison of input and measured ACIS 'b' band fluxes for sources with power-law spectra
  - Bins in red contain fewer than 100 measurements; bins in blue contain 100–400 measurements; bins in black contain more than 400 measurements



## Catalog Properties: Flux Distribution

#### Flux Distribution

• Although CSC fluxes range from below  $\sim 10^{-18}$  erg cm<sup>-2</sup> sec<sup>-1</sup> (for the deepest exposures) to  $\sim 10^{-10}$  erg cm<sup>-2</sup> sec<sup>-1</sup>, most CSC sources have fluxes of  $\sim 10^{-15}$ – $10^{-13}$  erg cm<sup>-2</sup> sec<sup>-1</sup> (ACIS "b" band, or 0.5–7.0 keV)



• Distribution of CSC fluxes in the broad (black), hard (blue), medium (green), soft (red), and ultra-soft (orange) bands, obtained from the catalog master source table flux\_aper column



## Catalog User Interface Updates

#### **CSCview**

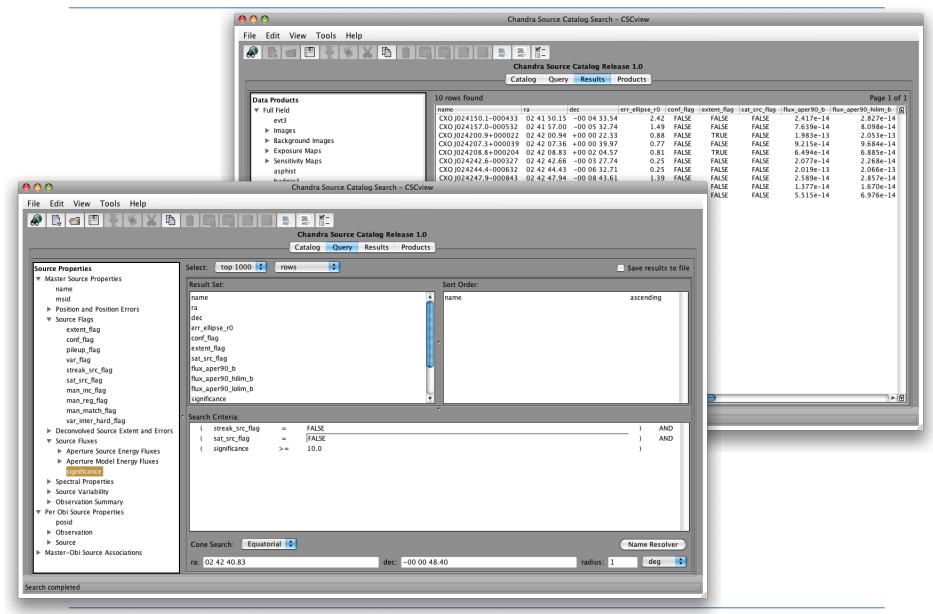
- CSCview is the primary user interface to the catalog
  - Available on the web via the catalog user website
  - Java applet runs in the user's web browser (requires Java version 1.5 or later)
- Numerous enhancements since last CUC meeting
  - Improved human interface looks more like a typical application
    - Standardized layout of menus, buttons, forms; standardized icons etc.
  - Supports ADQL VO query language in addition to the form interface
  - Provides access to several "pre-canned" result sets
    - These will be made more visible in a future release
  - Much improved interface for accessing and retrieving file-based data products
  - Improved boolean expression functionality for query form search criteria
  - Improved cone search functionality, including name resolution
  - Provides better extensibility "under the hood"

### **cURL/wget Interface**

Updated to accept ADQL queries



### **CSCview GUI**





### **Future Directions**

#### **Short Term Plans**

- Catalog Releases
  - Release 1.1 (Fall 2009) extends Release 1 to include public HRC-I imaging observations, and newly public ACIS observations, but otherwise retains the same limitations as Rel. 1
- Public Interfaces
  - CSCview GUI
    - Additional output file formats for query results (VOTable, Vizier TSV)
    - Support for cross-matching with user supplied catalogs
    - SAMP interface
  - Web Services
    - VO cone search service
    - Catalog limiting sensitivity and footprint services
    - CSC SDSS (DR7) cross-match
  - Additional Interfaces
    - Display catalog sources directly in **ds9** version 5.6
    - Google Sky Interface
      - » Catalog visualization for both scientific and education/public outreach users
    - External Interfaces
      - » CSC file dumps currently works with **TOPCAT** ("TST" format)
      - » In discussions with Vizier to host catalog tables
- CIAO tools
  - A number of the catalog-related tools will be made available to users in CIAO 4.1.2/4.2
    - Details discussed in SDS CIAO presentation



### **Future Directions**

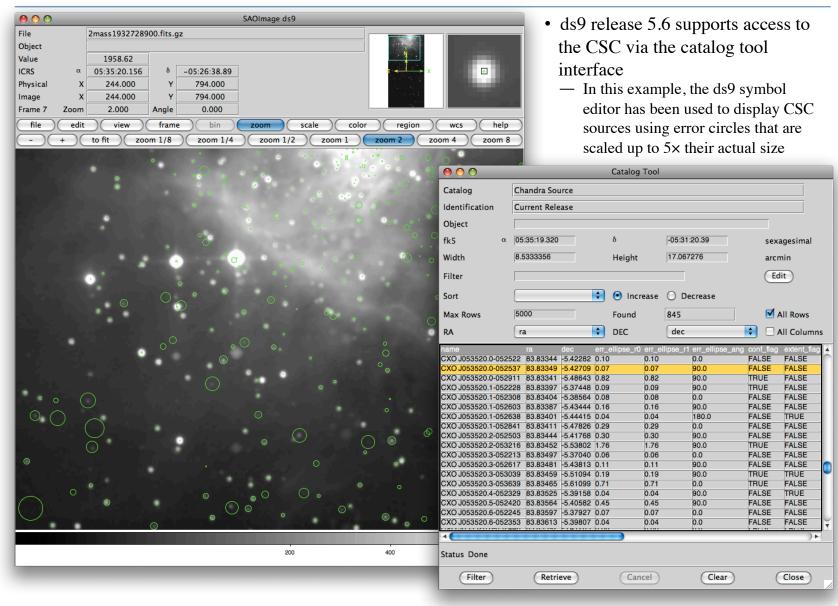
### **Longer Term Plans**

- Catalog releases
  - Catalog Release 2 (late 2010)
    - Co-add multiple observations of the same field that use the same instrument prior to source detection to achieve fainter limiting sensitivities
    - Improve background modeling, particularly in areas of extended emission
    - Improve aperture photometry in crowded fields
    - Improve extended source handling
  - Future Releases
    - Simultaneous source detection across overlapping observations with different detectors and pointings (and thus very different local PSFs)
    - Detection and classification of very extended sources
- Detailed plans for Release 2 and later are not fully established
  - Items identified above for Release 2 will have significant impact on the depth and quality of the catalog
  - Plan is to seek community feedback for guiding further development
    - Want input from a broad audience that includes the general multi-wavelength community and well as experienced Chandra users
    - Catalog needs several month soak period in the community for useful feedback from a wider audience
    - Consider establishing community working groups with focus on specific areas of complexity (such as robust extended source detection) if there is sufficient interest



### ds9 Release 5.6 Interface

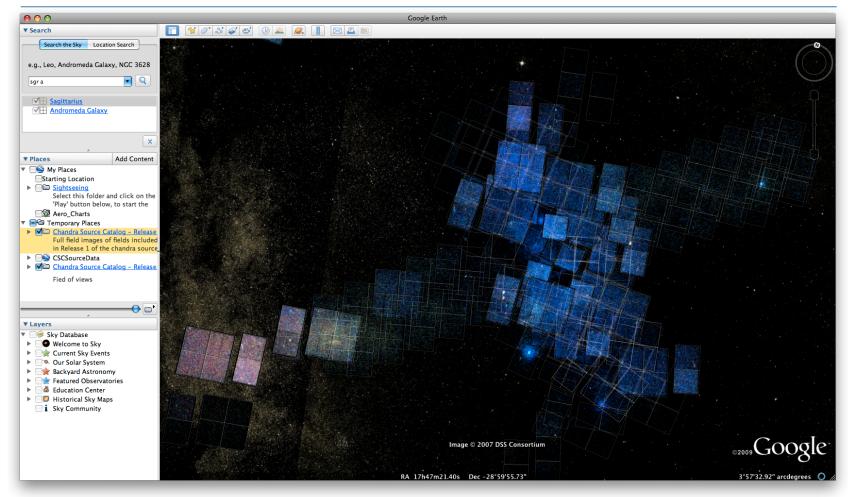






## Google Sky Interface



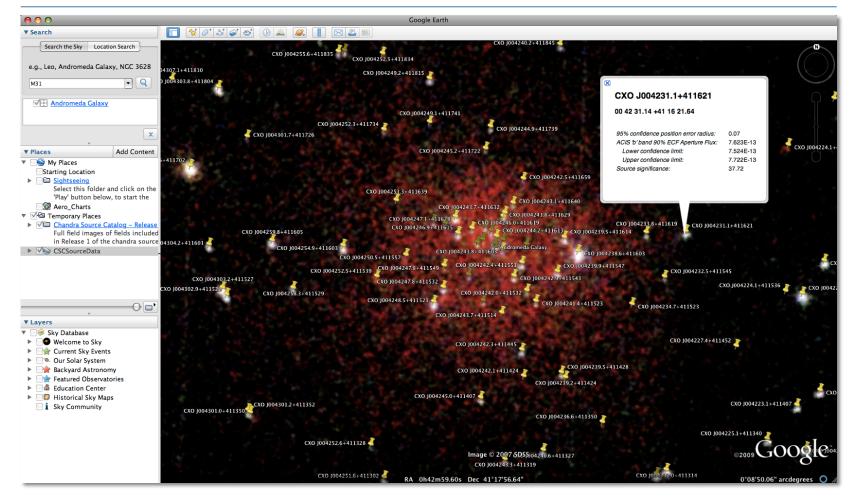


- Currently testing an experimental Google Sky interface for visualizing the CSC contents
  - Platform independent interface provides quick visualization of catalog contents for scientists
  - Simple to build and easily extensible; could be readily adapted for education and public outreach uses
  - Panel displays the catalog coverage in the vicinity of the Galactic Center, with observation FoVs displayed (optional)
    - Fields are blue because the observations were obtained using ACIS with an E > 1.0 keV energy filter



## Google Sky Interface (cont.)





- Summary source data can be added to the displayed catalog observation images
  - In this case, push pins locate and name CSC master sources in the core of M31, with associated simple source summary pop-ups, but graphical source regions and more sophisticated associated data can easily be added