

Mission Planning Updates

P. Slane (Chandra Mission Planning)

Scheduling with Constraints



- 1. <u>Science Constraints</u>
 - Chandra currently accepts up to 15% constrained observations
 - Includes, e.g., roll, window, phase, un-interrupt constraints
 - Also coordinated observations.
 - Also accept up to 80 TOO observations (including follow-ups)
 - 8 very fast, 20 fast, 26 medium, 26 slow
- 2. <u>Spacecraft Constraints</u>
 - Restrictions such as not pointing toward the Sun provide constraints on when targets can be observed
 - Over time, the number of such constraints has increased considerably, leading to challenges in planning, and some impacts on observing



Thermal Issues





- Degradation of MLI results in increasing temperatures for many spacecraft components (e.g., EPHIN, shown above).
 - Monitoring, investigation, and modeling has permitted relaxation of some constraints to minimize impact to science observing
 - Major modifications to planning and scheduling observations have allowed continued highly-efficient observing
 But not without some impact

Pitch Angle Limitations





- At present, some component is impacted at every possible pitch angle.
- Maximum dwell time are significantly restricted at near-normal attitudes.
- This presents challenges for long observations, particularly in portions of the sky far from the ecliptic.
- Challenges also exist for coordinated observations with observatories that always observe at near-normal Sun attitudes (e.g., XMM)



- Scheduling is a balance of heating and cooling multiple s/c components
 - Observations often split to accommodate temperature limits



- ACIS electronics (DPA) and focal plane temperatures increase at high pitch
 - competes with restriction for S/C heating at Sun-normal attitudes
- Reducing # of CCDs in use results in lower FP and DPA temperatures
 - Max dwell ≤100 ks for 5 CCDs for pitches of 145-155; no current limits for <5 CCDs
 - Long tail-Sun dwells can yield FP temperature swings that could affect spectroscopy

Constraint Mitigation

- Observation splitting
 - Short durations for "bad pitch" targets
 - Splitting of "cold" targets for cooling
- GOs encouraged to use < 5 CCDs since Cycle 15
 - most programs meet science goals with < 5 CCDs
 - GOs can still request 5 or 6 CCDs; such observations can be more complicated to schedule
 - GOs are encouraged to specify optional CCDs that could enhance science if available
- Limit time at high ecliptic latitude
 - Maximum time at $|\beta| > 60^{\circ}$ limited to 2 Ms for XVP in Cycle 16; limits will continue in future cycles

<u>Two-week loads</u> (start_my_week → commence_thine_fortnight) Facilitates handling of components with long therm

- Facilitates handling of components with long thermal time constants (e.g., IPS tank)

Increased scrutiny on constraints (particularly hidden constraints) at peer review

- Emphasis on "preference" nature of Preferences as well





Future



- With current constraints, continued high-efficiency observing expected
 - No changes to observing requirements in Cycle 17
 - None anticipated for Cycle 18 at present, but this will be revisited in 1 yr
- Thermal constraints have, for quite some time, blurred the transition between observing Cycles; this will continue
 - Some targets get pulled forward to help with thermal issues; this necessarily delays some targets from current Cycle
 - Coupled with "borrowed" time in XVP programs, Cycle 15 will "end" (50/50 split of Cycle 15/16 targets) in early/mid-March 2015
 - Future Cycles will run on similar calendars, roughly March-to-March.
- Modeling and evaluation of spacecraft thermal conditions will continue in order to assess operational and scientific impact of evolving thermal conditions on S/C
- New mitigation strategies will continue to be developed as required
 - Discussions will proceed through Project Office and CUC