Introduction to CIAO

(Chandra Interactive Analysis of Observations)

CIAO: Chandra's data analysis system
Fruscione et al. 2006, SPIE Proc. 6270, 62701V, D.R. Silvia & R.E. Doxsey, eds.

THE PEOPLE WHO DID IT!

CIAO: Chandra's data analysis system

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ACKNOWLEDGMENTS

The CIAO software package is the product of many years of work on the part of very numerous people and the authors of this paper, all current or recent members of the CXC Science Data System group (SDS) are purely reporters on behalf of a much larger group of contributors.

In particular we would like to acknowledge the valuable contributions from past members of the SDS group: David Davis, Adam Dobrzycki, Holly Jessop, Joel Kastner, Casey Law, Eric Schlegel, and Jennifer West.

None of the CIAO design and development would have been possible without the expertise and hard work of the members of the CXC Data System (DS) group and in particular Giuseppina Fabbiano, Janet DePonte Evans, Ian Evans, Mark Cresitello-Dittmar, Stephen Doe, Gregg Germain, Kenny Glotfelty, Tom Calderwood, Judy Chen, Roger Hain, Helen He, Janine Lyn, Joseph Masters, Warren McLaughlin, Robert Milaszewski, Joe Miller, Dan Nguyen, Jim Overly, Brian Refsdal, Chris Stawarz, Michele Wilson plus numerous past members of the DS group.

Support for development of CIAO is provided by the National Aeronautics and Space Administration through the *Chandra* X-ray Center, which is operated by the Smithsonian Astrophysical Observatory for and on behalf of the National Aeronautics and Space Administration under contract NAS 8-03060.

From "ahelp ciao"....

- A powerful data analysis system originally written for the needs of users of the Chandra X-ray Observatory.
- Built to handle *N-dimensional* data without concern about which particular axes are being analyzed: Chandra is the first mission with 4-dimensional data (2 spatial, time, energy) in which each dimension has many independent elements.
- Mission independent (a part from a few instrument specific tools). It is now used to analyze data at all wavelengths.

- Provides users with the ability to filter down and project the 4-D Chandra event data to manageable sizes and convenient array: all CIAO tools take a '*filtering and binning*' specification on the command line, making use of a general purpose 'regions' syntax: "`ahelp dm`" for information on the Data Model that makes all this possible.

- To keep track of how the data had been filtered and binned CIAO relies on the 'data subspace'. Tools keep track of this subspace automatically and allow users to review previous data processing: see "`ahelp subspace`".

- The CIAO design allows close interconnection of tools. For example, the output of any of source detection program can be fed into *dmextract* to create a summed spectrum which can then be fit in *Sherpa*.

- The modeling and fitting tool ***Sherpa*** is central to the CIAO system. Sherpa performs forward fitting of models to data in N-dimensions. Sherpa (and ChIPS!) includes the "**Python**" and "**S-Lang**" languages which can be used for scripting and data manipulation. See "`ahelp sherpa.`

http://cxc.harvard.edu/ciao : it all begins here!

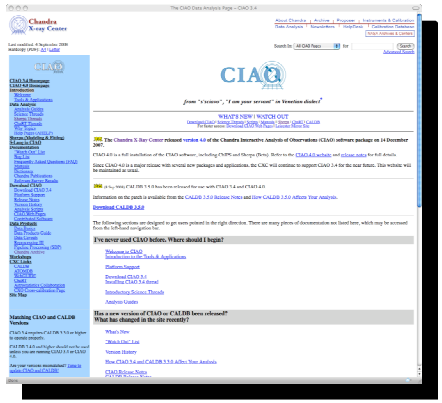
Navigate here

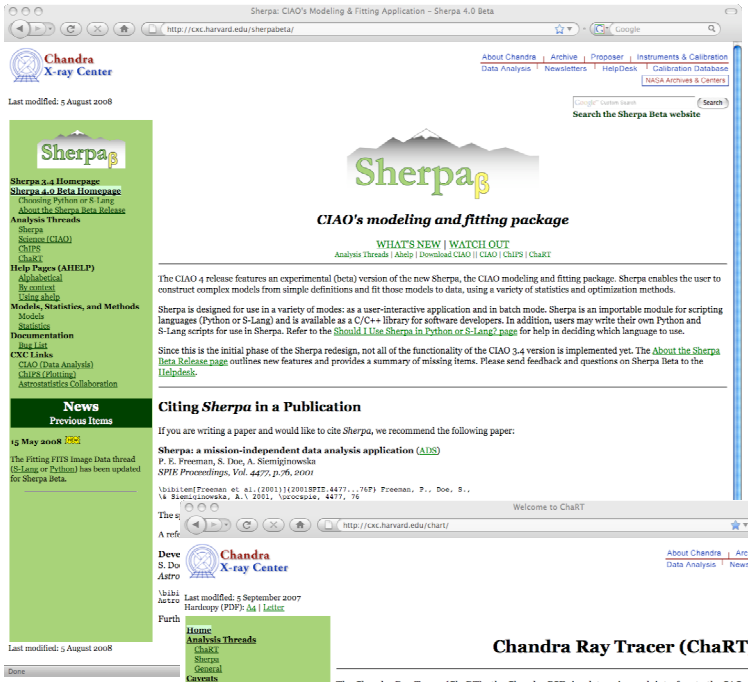
Look in here

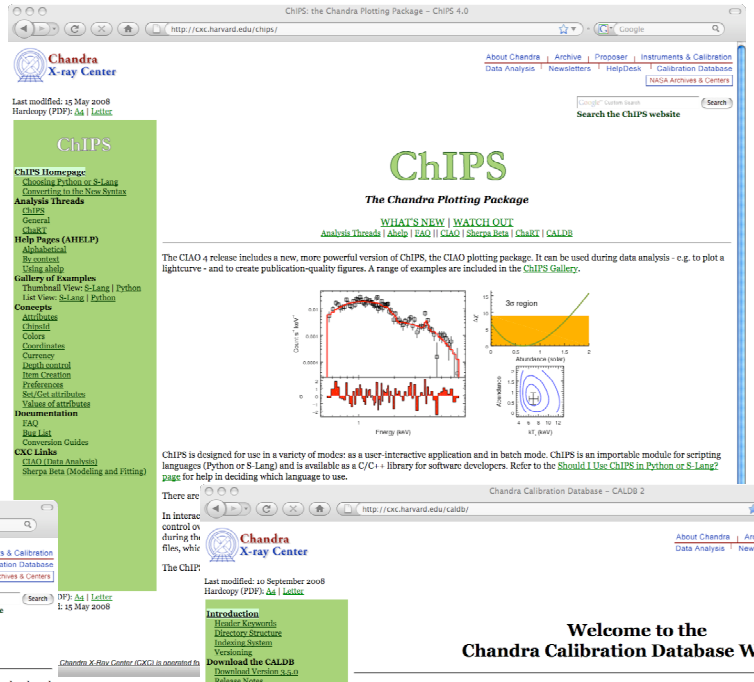
Search here

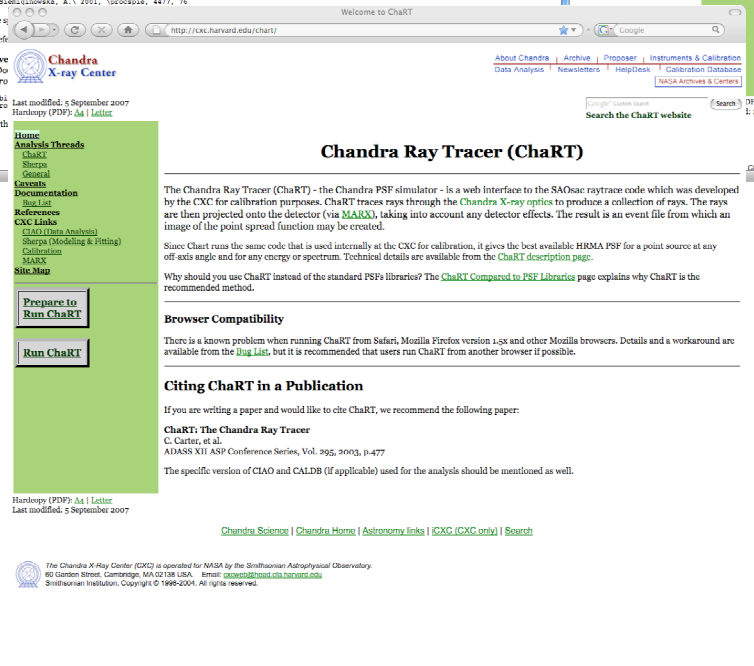
Quick links

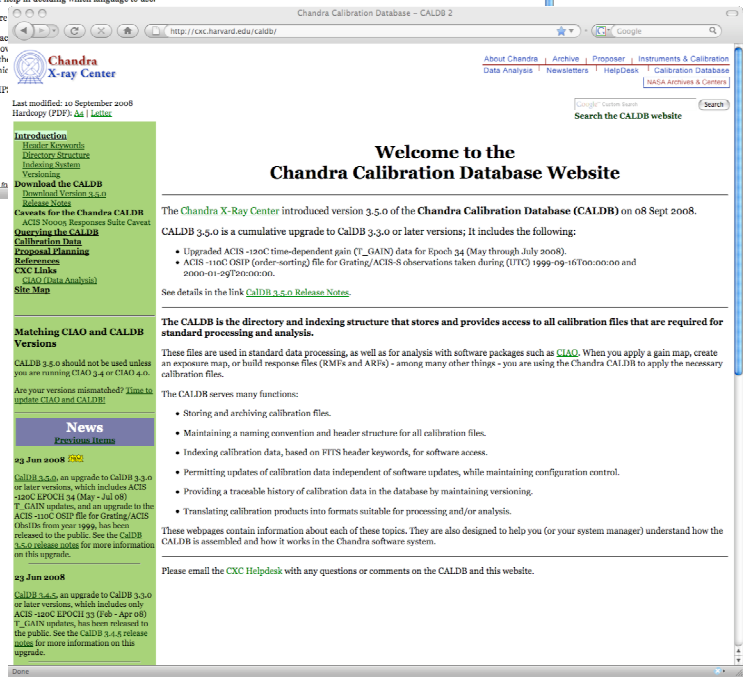
CIAO 3.4




Sherpa: CIAO's Modeling & Fitting Application – Sherpa 4.0 Beta
Chandra X-ray Center
Sherpa
CIAO's modeling and fitting package
WHAT'S NEW | WATCH OUT
Citing Sherpa in a Publication


ChIPS: the Chandra Plotting Package – ChIPS 4.0
Chandra X-ray Center
ChIPS
The Chandra Plotting Package
WHAT'S NEW | WATCH OUT


Welcome to ChaRT
Chandra X-ray Center
Chandra Ray Tracer (ChaRT)
Browser Compatibility
Citing ChaRT in a Publication


Chandra Calibration Database – CALDB 2
Chandra X-ray Center
Welcome to the
Chandra Calibration Database Website

CIAO help!

AHELP

- CIAO comes with its own help system called **ahelp**.
- Every component of the CIAO system has its own help text: tools (e.g. `dmcopy, wavdetect`), applications and their components (e.g. `sherpa, fit, model`), scripting language and its functions (e.g `slang, get_fit, get_data`) and numerous concepts (e.g. `regions, coords, dmsyntax, etc.`).
- The entire collection of ahelp files currently has more than 1000 pages!

The **ahelp** commands (see `ahelp ahelp` for details) access the CIAO on-line documentation. For example:

% about <word> (e.g. `about contour`) gives a list of subjects related to the query
% ahelp <name> (e.g. `ahelp tgextract`) provides by default the ASCII version of the help file.
% sherpa/chips> ahelp("name") (e.g. `ahelp("fit")`) provides the ASCII version of the ahelp file within the sherpa or chips environments.

DATA ANALYSIS GUIDES : *start from here!*

- A roadmap through the threads.
- Arranged by instrument (ACIS, HRC) or type of analysis (e.g. extended sources)

DATA ANALYSIS THREADS (*General, Sherpa, ChIPS, Chart*)

- About 100 CIAO processing recipes designed to teach users by leading step-by-step through a procedure.
- Added or updated as needed (look for the "**new**" or "**updated**" icons!).
- Several threads and groups of threads have been or will be made more automated in the form of scripts.
- PDF version available on-line, and threads are printer friendly.
- Quick "overview" at the beginning (synopsis, purpose, when to use etc.)

IMPORTANT CIAO WEB PAGES

Read especially the

What'sNew, "Watch out", Bugs, Caveats, Dictionary, FAQ pages, Platform support, Why topics, How does CIAOX.X and CALDBY.Y affects my analysis, Manuals and Memos, !

HELPDESK

When everything else fails....

http://cxc.harvard.edu/helpdesk/

with CIAO version, platform, data, purpose etc…

See Elizabeth Galle talk!

WHAT IS CIAO?

- Collection of programs (*tools*, *applications, scripts, S-Lang and Python modules*).
- Generally run from the shell; some part can be run from GUIs.
- Source code is available to users.
- Available on several platforms; currently (as of CIAO 4.0) supported:
  - Solaris 8
  - Fedora Core 4 and 7 (also for other linux platforms)
  - Mac OS X 10.4 and 10.5 Intel
  - Mac OS X 10.4 PPC

  Has been successfully installed by users on other "unsupported"(= no regression testing done!) platforms: check the ***Platform Support*** page

The CIAO Environment

We recommend the use of an alias called "`ciao`" to start up the system: it sets up number of environment variables and path assignments.

`ciao -v` tell what version you are using (useful when reporting problems)

General Concepts

- File format
- Parameter Files
- Filters
- Regions
- Good Times Intervals (GTI)
- Scripting languages: Python and S-Lang

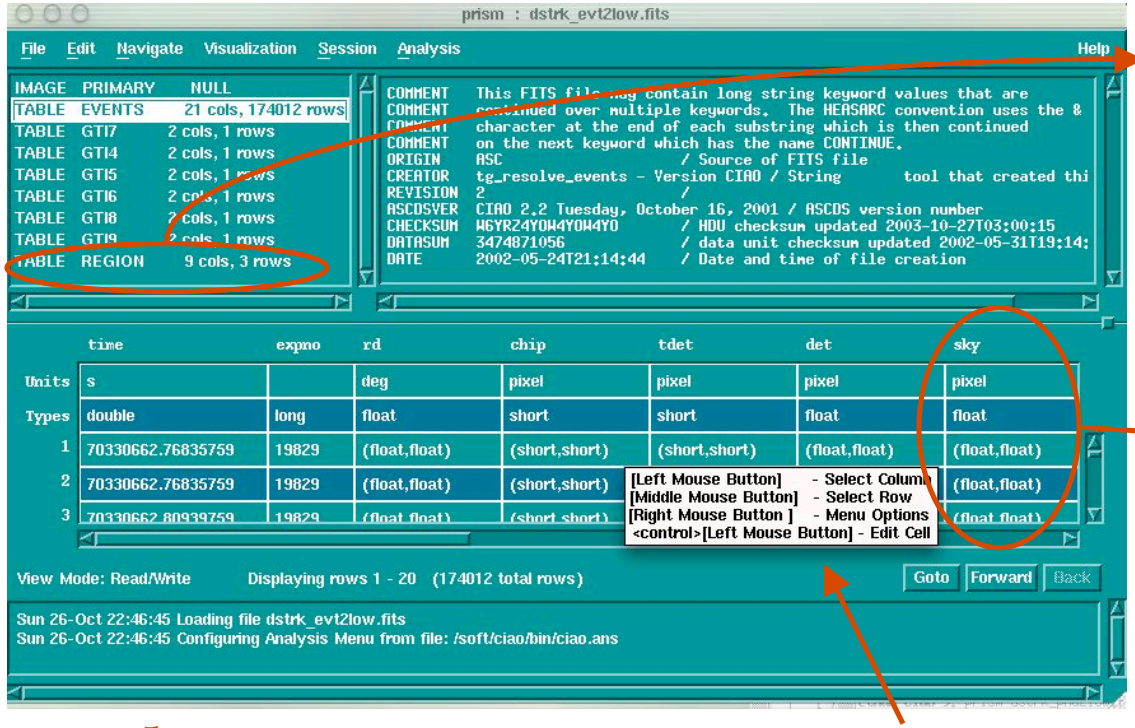
File Format

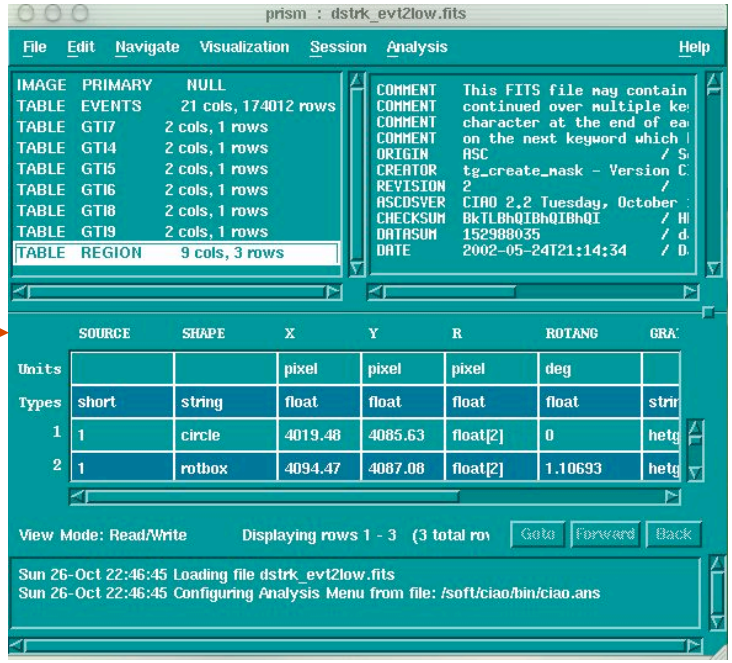
- Chandra data is stored in FITS format. As of CIAO4.0 ASCII (text) files can now be handled by many tools and applications through the new software library known as the "ASCII kernel" (`ahelp dmascii`)
- When CIAO operates on data it stores processing state/information along with data (keywords, subspace).
- A single file can contain multiple "datasets" (e.g. data, GTI, weight map, regions) stored in "blocks".
- Blocks can contain image or table data. Table columns can be vectors.
- dmlist (a command line tool) or prism (a GUI) are available to view file contents.

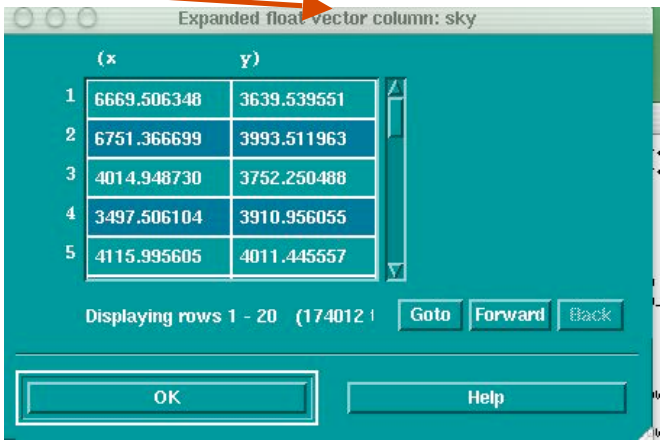
Prism: file viewer, editor, tool launcher

blocks

header



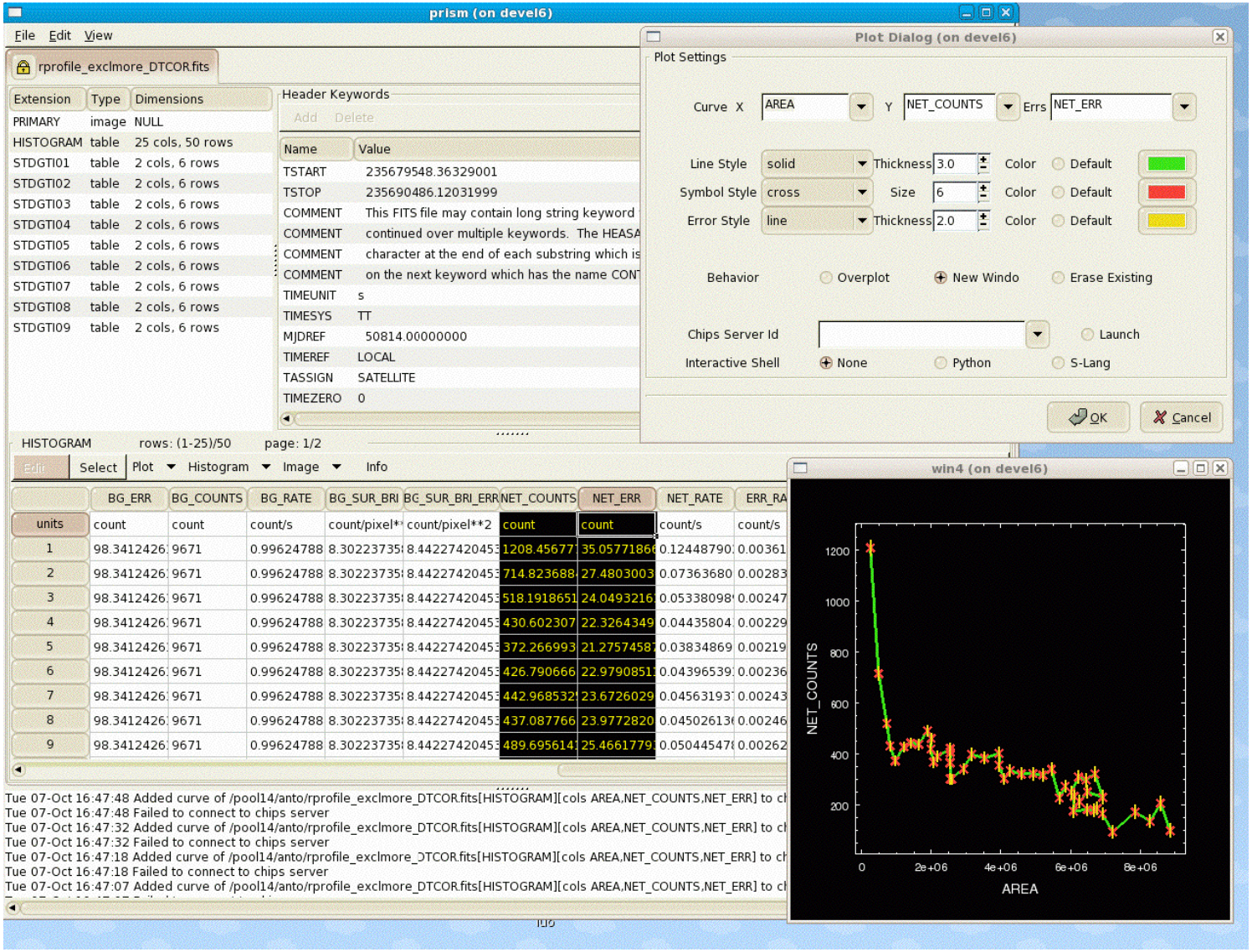




data

quick help

Prism: a preview of what is coming in CIAO 4.1!



Parameter Files (`ahelp parameter`)

- Processing parameters for CIAO tools can be set on the command line or, as with IRAF and FTOOLS, using parameter files.

- Parameters files are
  - ✓ stored in *$HOME/cxcds_param4/*
  - ✓ are called *<tool>.par* (e.g. dmcopy.par)
  - ✓ are ASCII files.

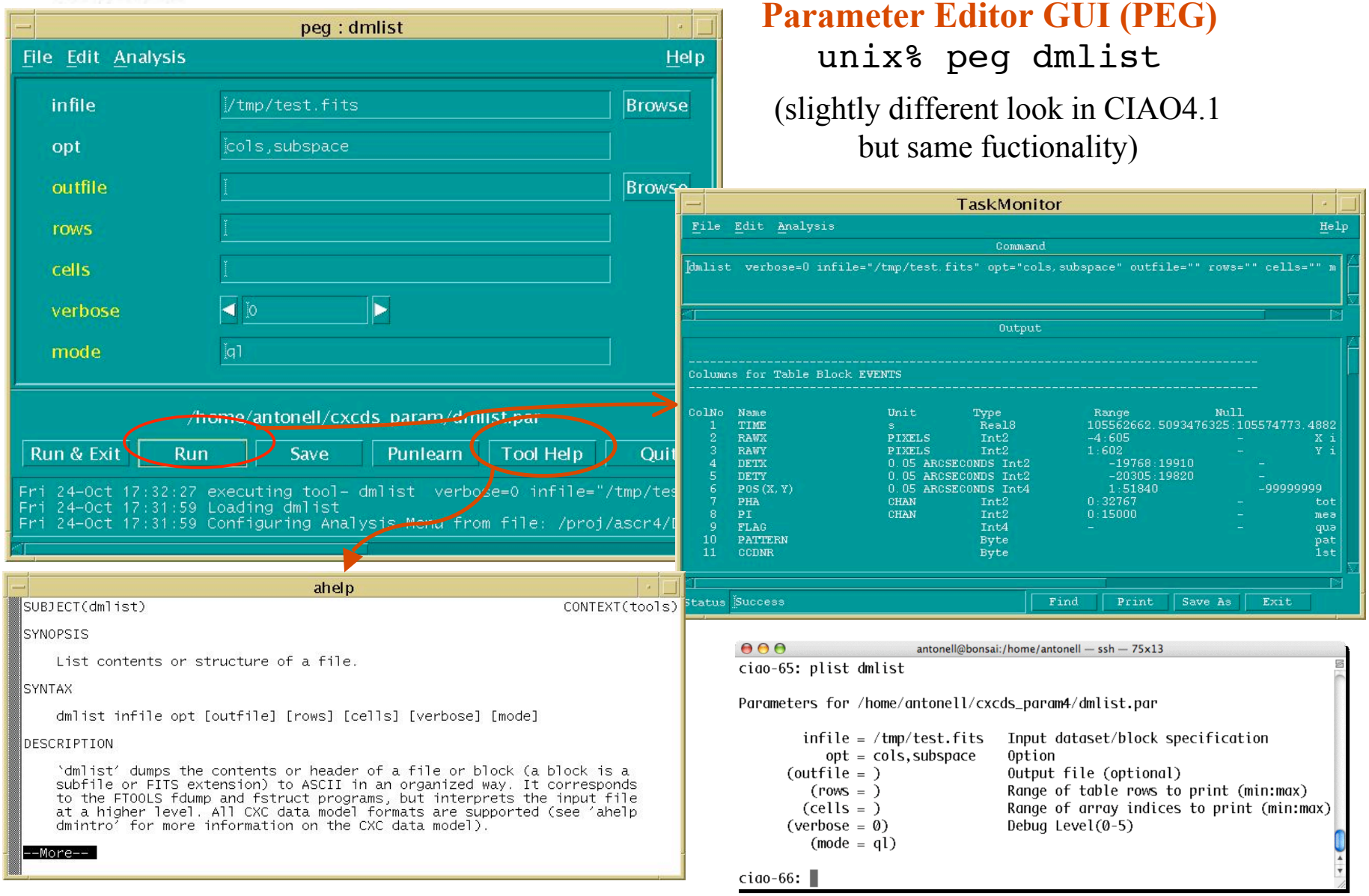
CAVEAT: Always delete or rename the *cxcds_param* when upgrading operating system or CIAO version

- A "Parameter Editor GUI" (peg) and a number of routines (e.g. plist, pset, punlearn) are provided to read and write to these files. A S-Lang interface to the CXC parameter library is also available (see `ahelp paramio`).

Parameter Editor GUI (PEG)

```
unix% peg dmlist
```

(slightly different look in CIAO4.1
but same fuctionality)



Filters, Regions, and GTIs

- filtering (removal of unwanted events) is an essential part of X-ray analysis - e.g. to remove periods of high background or poor aspect solution, exclude uninteresting sources from an image etc.
- the DataModel (DM) provides great filtering flexibility: e.g. dmstat "evt2.fits[EVENTS][energy>300][cols -grade]" (see `ahelp filtering`, `ahelp dmimgfiltering`)
- GTIs (Good Time Intervals) are used to define what times periods of the observation can be used (i.e. contain valid data). They are generally stored as a block in the event list (see `ahelp chandra times`)
- Regions are used to define the source and background areas of an image. They are text files that can be created manually or within ds9, and are used as a filter (e.g. "[sky=region(source.reg)]"). (see `ahelp dmregions`)
- Subspace records the filters applied to a file; dmlist can read this history using opt=subspace (see `ahelp subspace`)

Scripting Languages

- Two scripting (or interpreted = no compilation is necessary) languages are supported in CIAO via Sherpa and ChIPS: Python (as of CIAO4.0) and S-Lang
- Sherpa and ChIPS are infact importable modules for Python and S-Lang.
- More info in Douglas Burke talk tomorrow

CIAO overview

Data manipulation: copy, filter, extraction, stats, etc.

Data preparation (or Chandra-specific instrument tools): update calibration, correct for instrumental effects, find & extract grating data, create aspect histograms

Response tools: exposure map, PSF, RMF and ARF

Source Detection: celldetect, wavdetect, vtpdetect

Timing & Background tools: lighcurve, power specrtrum, barycenter correction

Convolutions, Transforms, & Smoothing: csmooth, aconvolve, acrosscorr, apowerspectrum

Plotting: ChIPS (*)

Modeling/Fitting: Sherpa (*)

S-Lang & Python: ChIPS and Sherpa modules; slsh/ipthon shells

GUIs: DS9, prism,peg

The "Data Model" and the Data Manipulation Tools

- The CXC analysis and processing software is built on a common versatile interface library called the CXC Data Model (or just DM).
- The DM provides users with a powerful built-in data filtering and binning capability.
- The name "Data Model" reflects the fact that the interface can be used on data files of different format (all described by a single abstract description - the same "model") in a transparent way.
- The latest addition to the DM in CIAO4.0 is the "ASCII kernel" which gives the ability to operate on ASCII (text) file the same way as on FITS files (eg for filtering, plotting etc.)

- An important characteristic of the DM is that ANY program that asks for a data file name as input accepts a *"virtual file"* string which causes the program to see a filtered version of the file in question.
- The "virtual file" syntax is also commonly used to create on disk a filtered version of the input file.
- Another important characteristic of the DM is that all columns of event lists are treated "equally": for example binning is allowed not only in spatial coordinates but also in e.g. time, or energy coordinate, giving the ability of creating multidimensional images in space-energy, or space-time, etc.

Data Manipulation Tools

The four DM "core" tools are:

dmlist: list contents or structure of a file

dmcopy: filter and bin tables and images

dmextract: make a histogram table file (e.g. PHA file, lightcurve file) from a table column. Generate count histogram on supplied regions for a spatial table or image file.

dmgti: create custom Good Time Intervals (GTIs) from a constraint expression

30+ data manipulation tools are included in CIAO4.0

DATA MODEL SYNTAX (`ahelp dmsyntax`)

- All CIAO tools use the DM library and therefore accept as input "virtual files" described using the DM syntax.
- In the DM context a "virtual file" in represented by a filename followed by a series of optional qualifiers in square brackets [ ]:

"filename[block][filter][columns/binning][options][rename]"

where:

block - is the "section" of the file to use

filter- is the filter to be applied

columns/binning - specifies either the columns from a table to be included in an output table or the binning. When binning the data to generate an n- dimensional image, the range and binsize (min:max:bin) must be specified.

options - a sequence describing special options for the DM library

rename - specifies a name for the new block

Note that:

- the order of the qualifiers generally matters, however...
- not all qualifiers need to be present always

Simple examples of "virtual files":

- Select the first three columns of the EVENTS block by number:
  ```
  acisf01843N001_evt2.fits[EVENTS][time=84245787:84247000][cols #1,#2,#3]
  ```
  or by name:
  ```
  acisf01843N001_evt2.fits[EVENTS][grade=0,2,3][cols time,ccd_id,node_id]
  ```
  after filtering in time or grade

- Bin an events file to create a PI spectrum for a specified region (input of dmextract):
  ```
  acisf01843N001_evt2.fits[EVENTS][sky=region(mysrc.reg)][bin pi=1:1024:1]
  ```
  or an image (input of dmcopy):
  ```
  acisf01843N001_evt2.fits[EVENTS][pha<100][bin x=320:480:4,y=320:480:4]
  ```

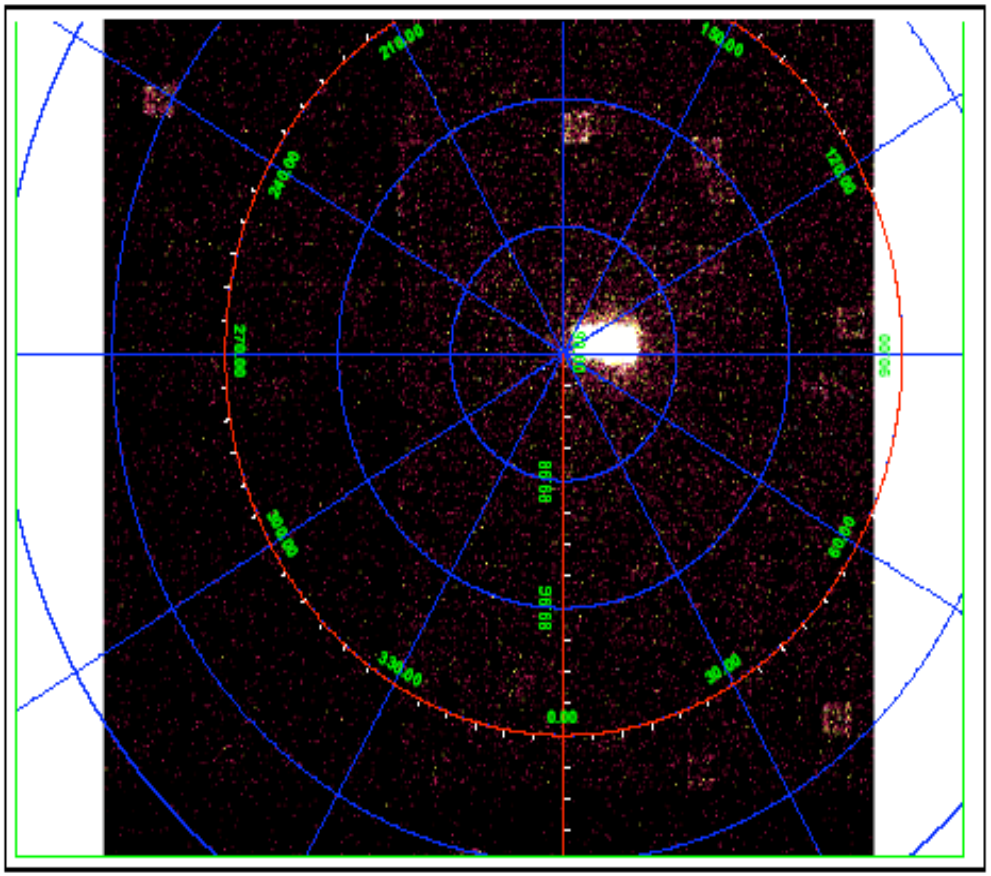
In the examples above:

block: [EVENTS]
filter: [time=84245787:84247000]
[grade=0,2,3]
[sky=region(mysource.reg)]
[pha<1000]
columns/binning:
[cols time,ccd_id,node_id]
[cols #1,#2,#3]
[bin pi=1:1024:1]
[bin x=320:480:4,y=320:480:4]

DM Examples 1: Detector Image

Imaging on multiple coordinate systems: first, let's look at a region in detector coordinates, filtered on energy and time.

dmcopy "merge3e.fits[energy=500:2000, time=:63940080, 63940180:][bin detx=3500:4500:2, dety=3500:4500:2]" det.img



DM Examples 2: Sky Image

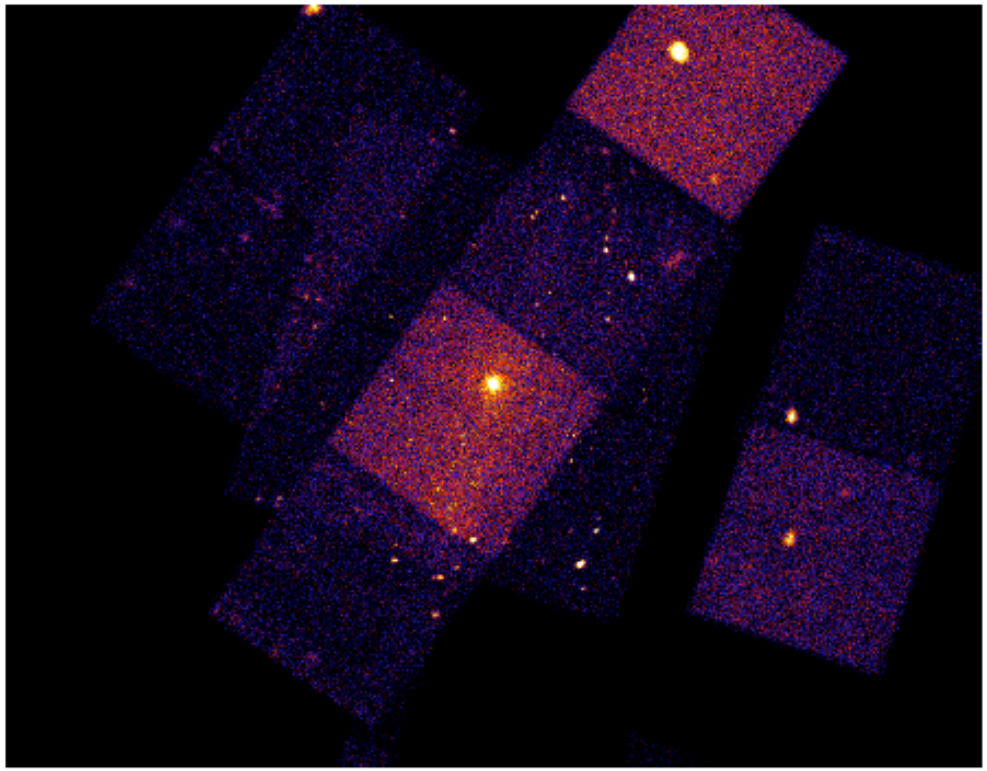
Now look at the same photons but in sky coordinates

dmcopy "merge3e.fits[energy=500:2000, time=:63940080,63940180:,detx=3500:4500,dety=3500:4500][bin x=3200:4800:2, y=3200:4800:2]" sky.img

DM Examples 3: Merged sky image

The whole field was created by merging three separate observations.

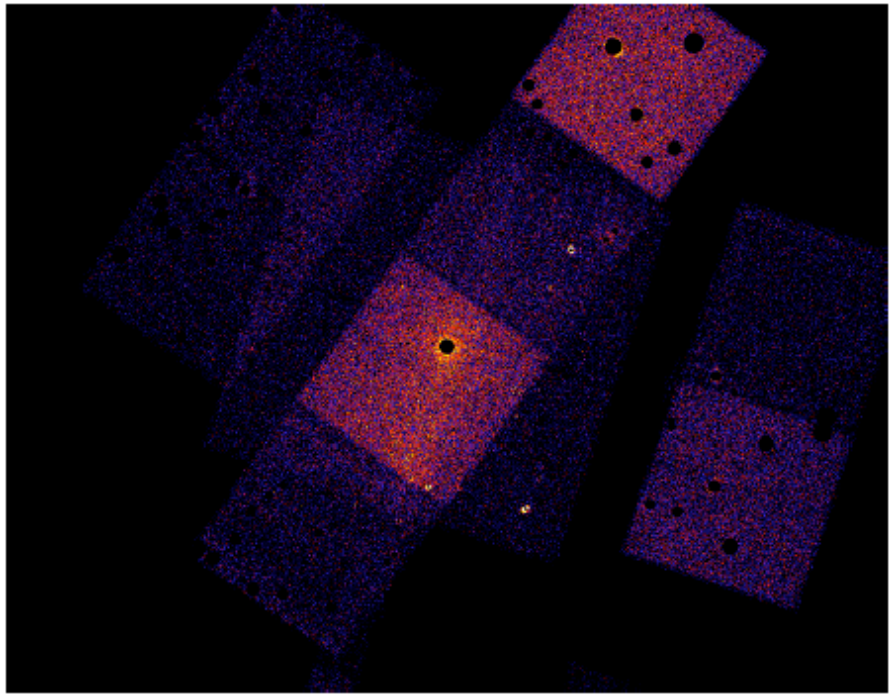
dmmerge “786.fits, 787.fits, 1730.fits” outfile=merge3e.fits



DM Examples 4: Removing sources

We can generate a background image by removing sources found by the automatics source detection program.

dmcopy “merge3e.fits[exclude sky=region (gg.reg)]” exclude.fits



DM Examples 5: Infrared spectroscopy data

ISO data: LWS LSAN file. This is a very simple file by wavelength and flux for the different detectors and scans are mixed together. We can use the DM tools to isolate a single scan and dump wavelength versus flux for it.

dmcopy "lasan59901083.fits[lsancnt=4][cols lsanwav, lsanflx]"
subset.fits

dmlist "lsan59901083.fits[lsanscnt=4][cols lsanwav, lsanflx]"
data,raw outfile=lis.asc