

a long GRB and a broad lined Type Ic supernova from a single central engine

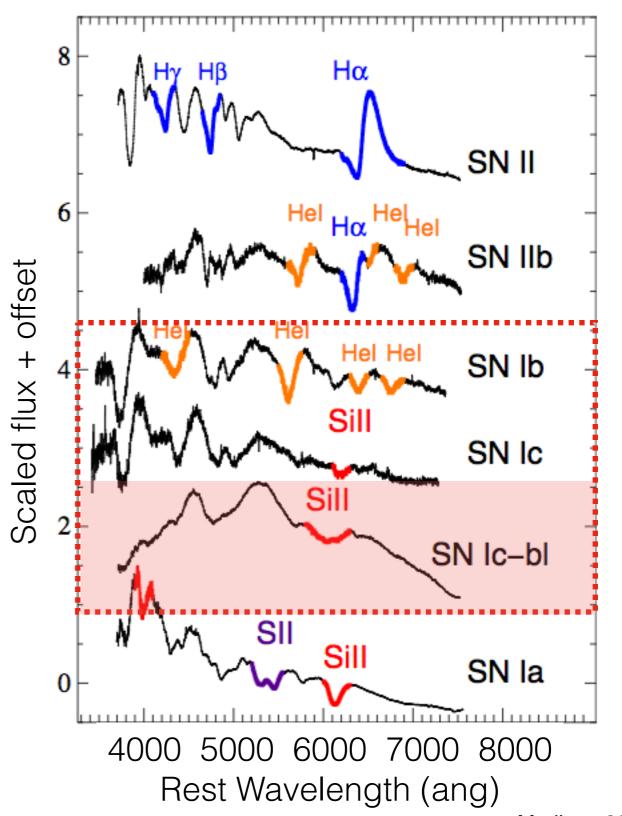
Jennifer Barnes Einstein Symposium October 12, 2017

with: Paul Duffell, Dan Kasen, Maryam Modjaz, Yuqian Liu, Andrew MacFadyen

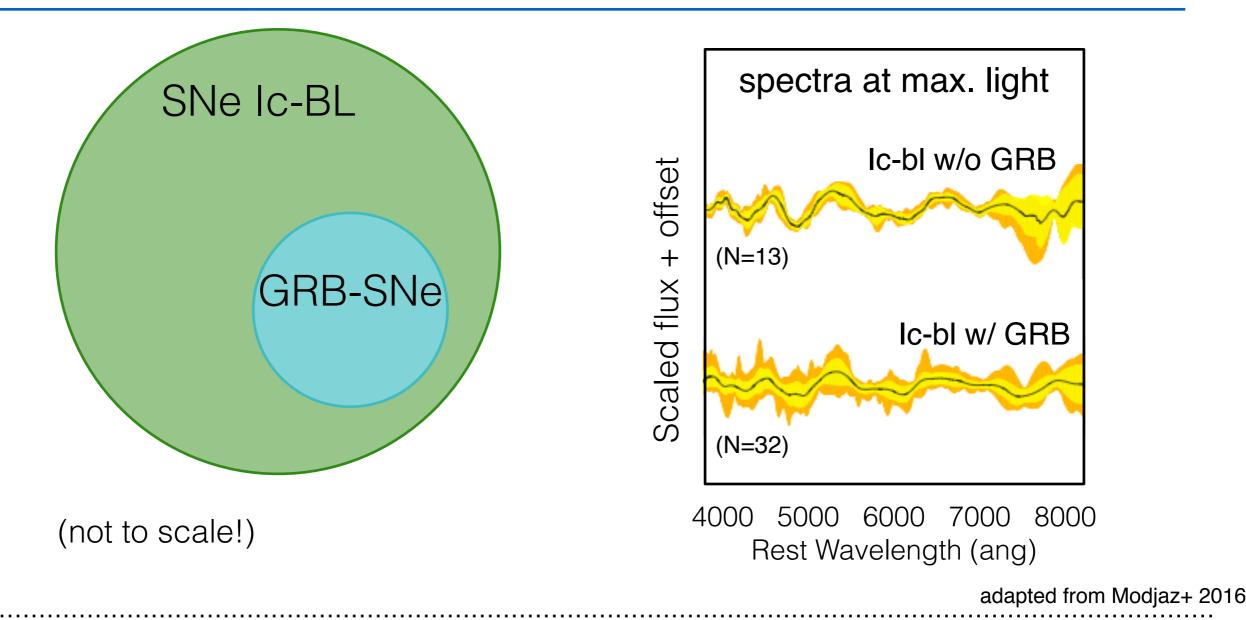
SNe Type Ibc

- Core collapse of stripped massive stars
- Spectra lack Hydrogen (lb, lc), and may lack Helium (lc)
- Subset of SNe-Ic exhibit very broad lines and strong line-blending

↓ "broad-lined Ic"



The Ic-BL/GRB connection



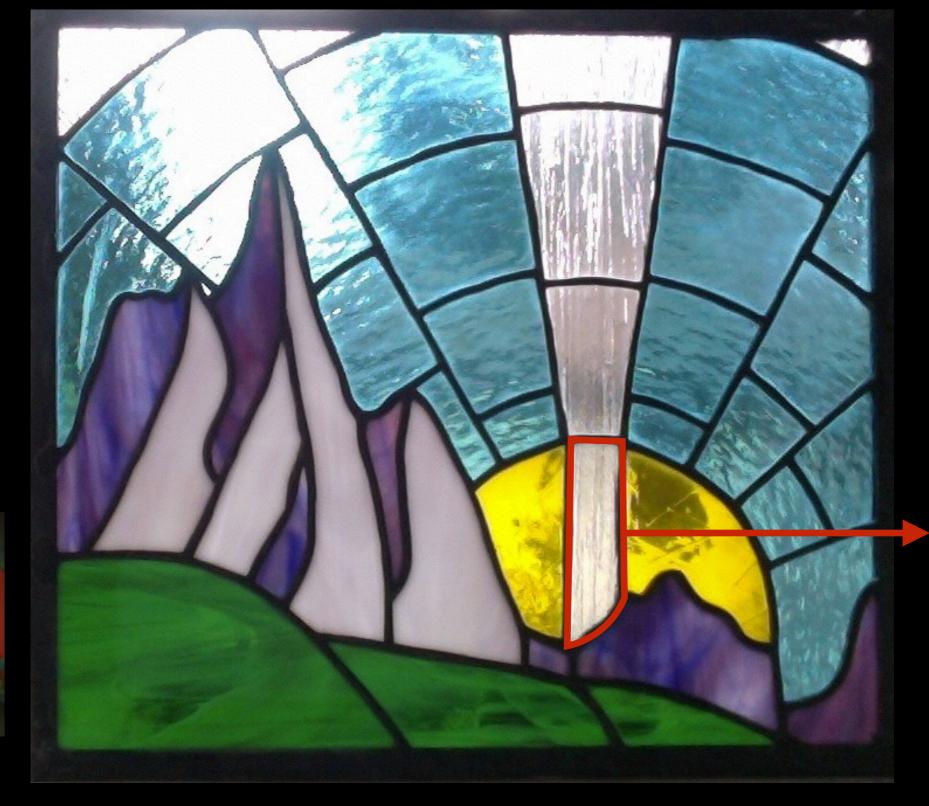
- What is the source of the extreme kinetic energies (10⁵² erg) implied by these broad lines?
- What is the nature of the GRB/Ic-BL connection?



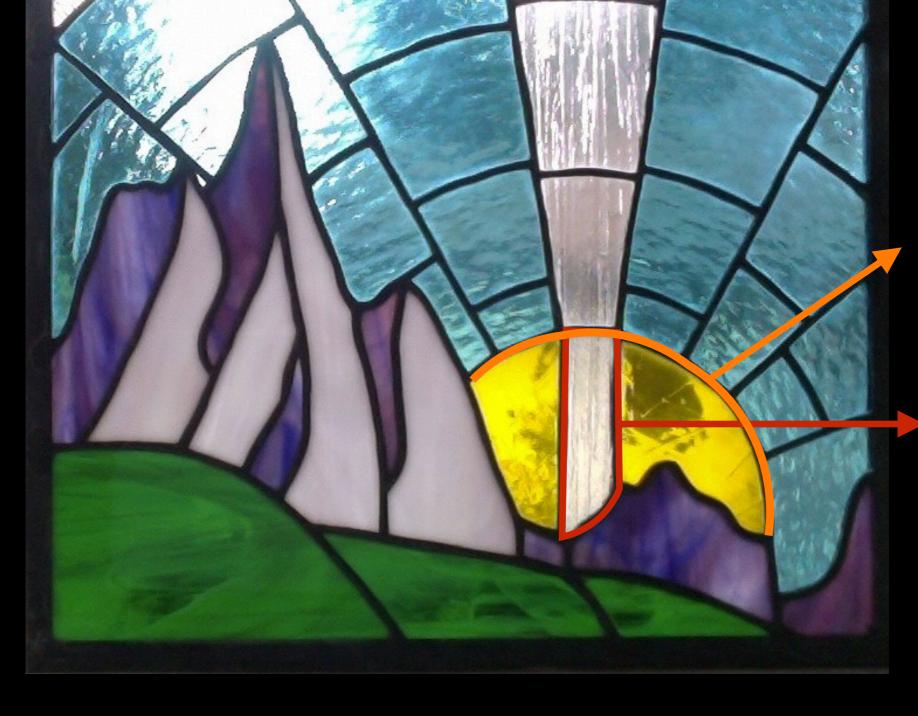




(Paul Duffell)



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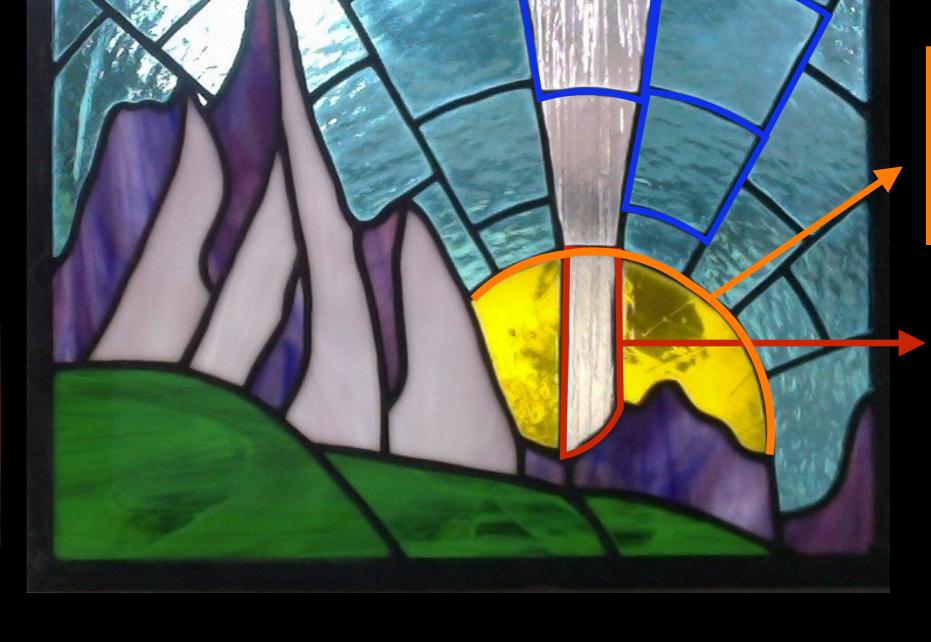


Wolf-Rayet Star



(Paul Duffell)

JET (m.m. rel. hydro)



Wolf-Rayet Star

(Paul Duffell)

Sedona Light curves, spectra

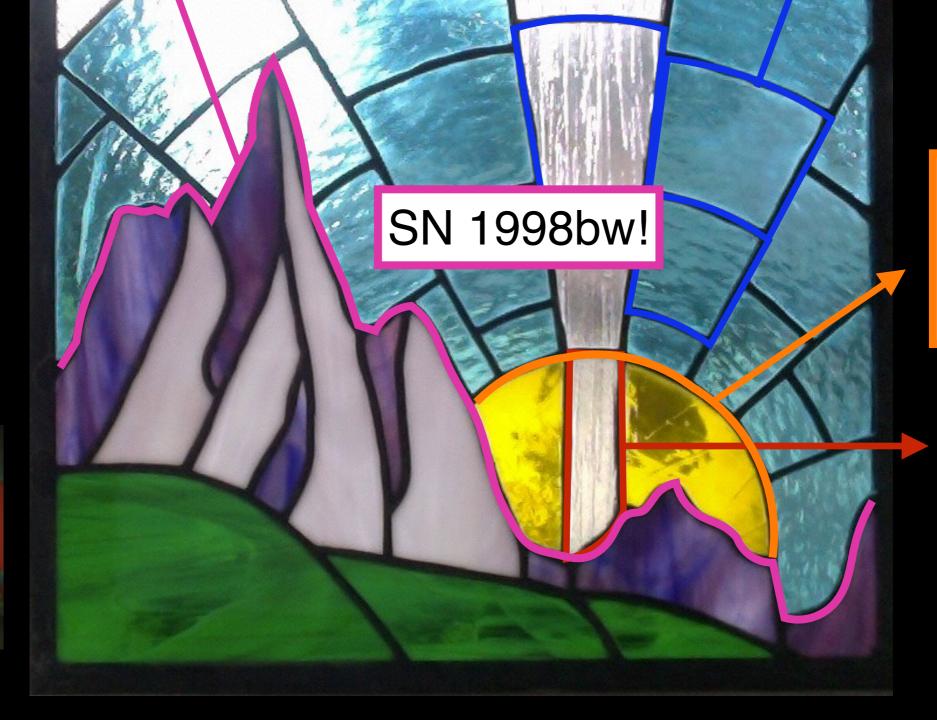
JET (m.m. rel. hydro)



Wolf-Rayet Star

(Paul Duffell)

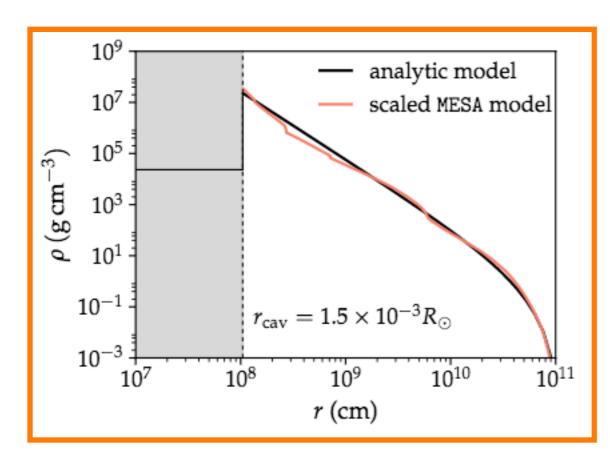
Sedona Light curves, spectra JET (m.m. rel. hydro)

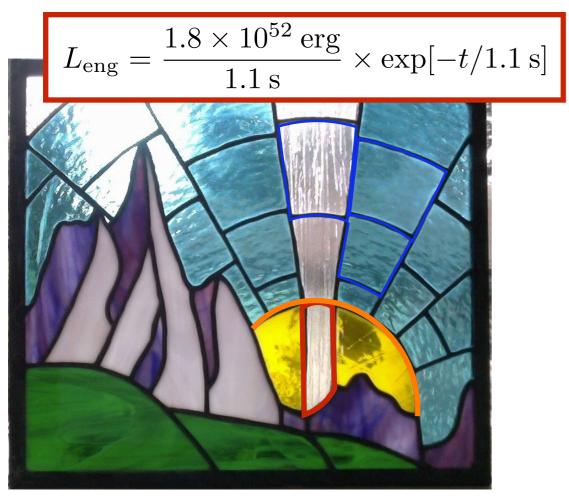


Wolf-Rayet Star

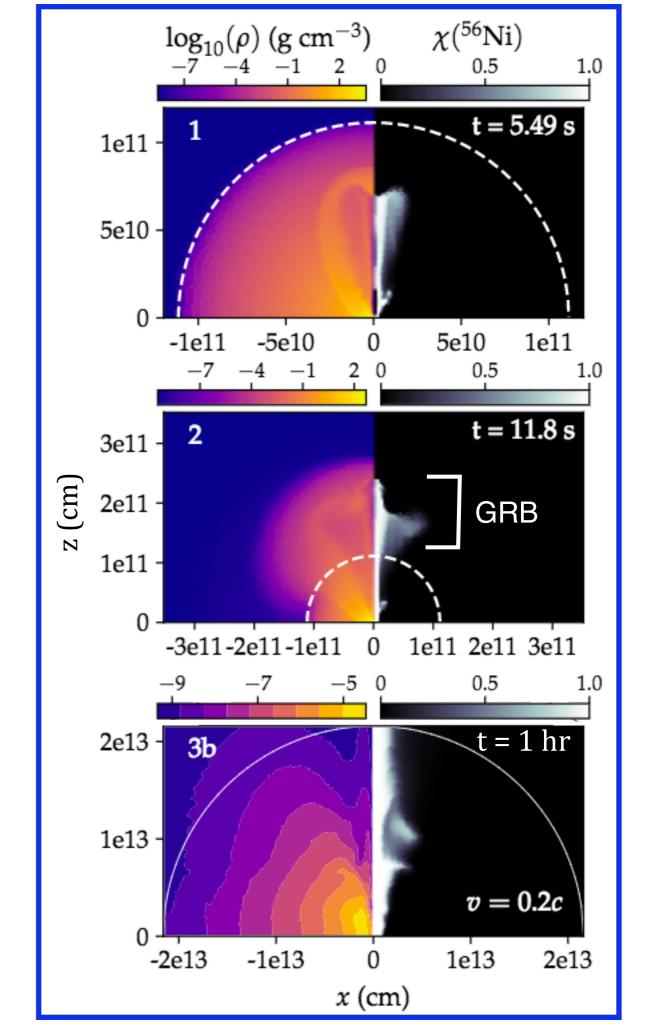


(Paul Duffell)

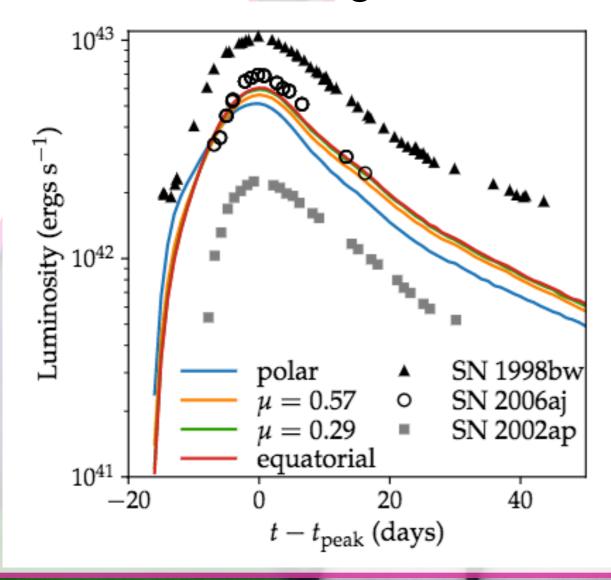




adapted from Barnes+17



Bolometric light curves



Spectrum at peak

