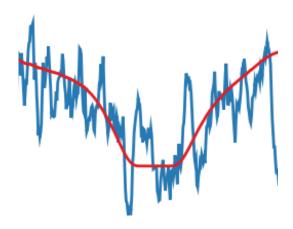
Teaching Computers to Detect DLAs

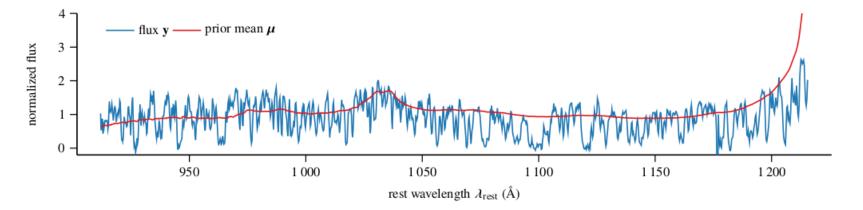
Roman Garnett (U Washington), Simeon Bird, Shirley Ho. 1605.04460, 1610.01165

Strong absorbers in quasar spectra

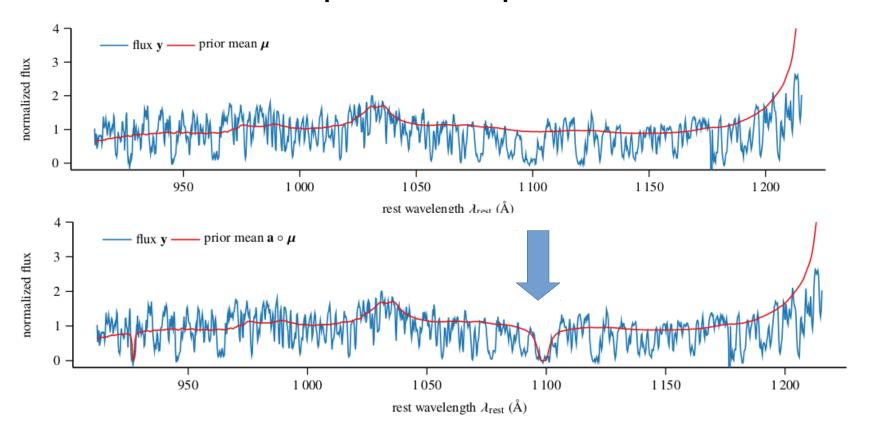
Neutral hydrogen from small galaxies



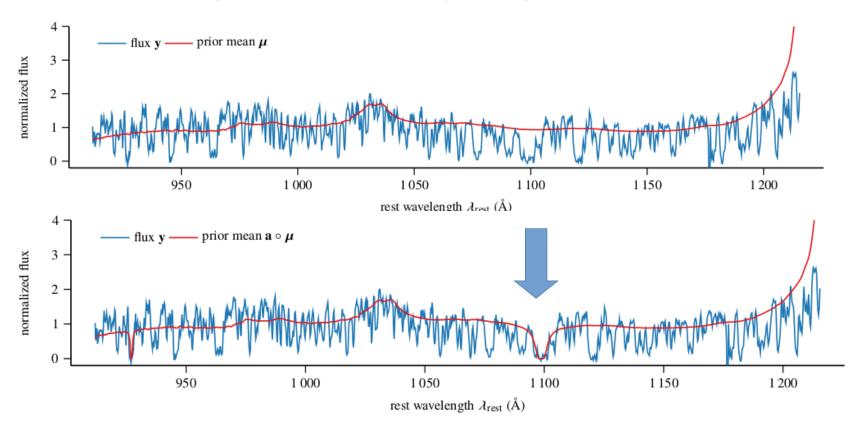
Currently done by **visual inspection** of spectra Look for wide dips in the spectrum below:



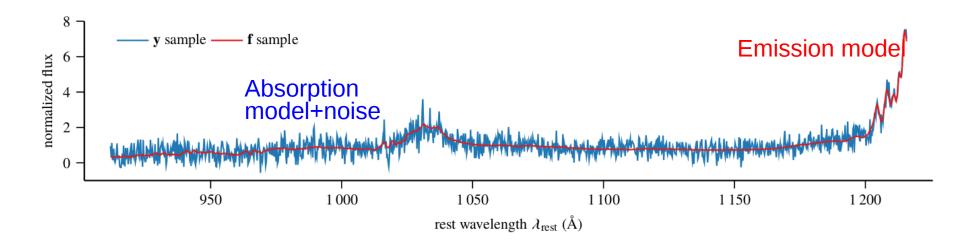
Currently done by **visual inspection** of spectra Look for wide dips in the spectrum below:



Future surveys will have > 10⁷ spectra Visual inspection clearly **impractical**



Learn Bayesian model for quasar without DLA Learn and test with DR9 visual catalog

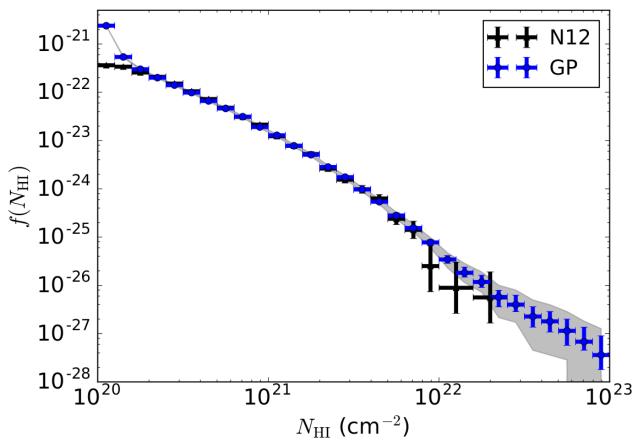


Model for DLA added using Voigt profile

Detecting DLAs

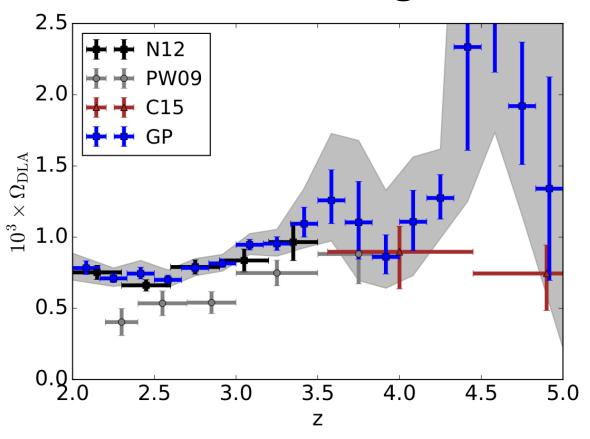
- This gives us posterior probability for DLA
- We can build an accurate model for DLA population even with poor detections
- Use all data, even with SNR < 1

Results: CDDF



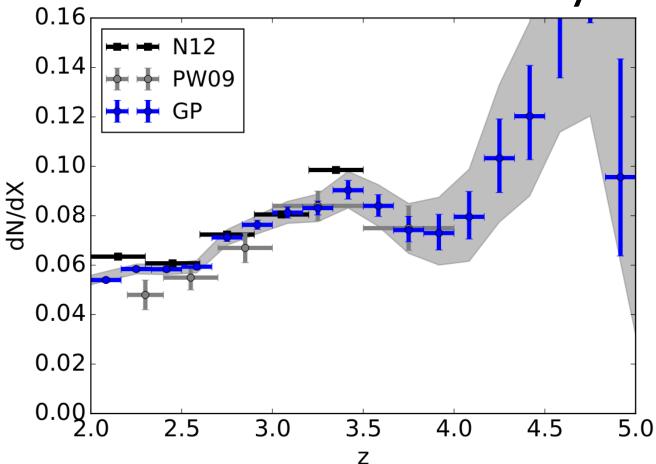
- Reproduce earlier datasets
- Small error bars!

Results: OmegaDLA



- Total gas mass in DLAs.
- Good agreement and z>4

Results: Line Density



- Number density of DLAs increases
- Error bars!

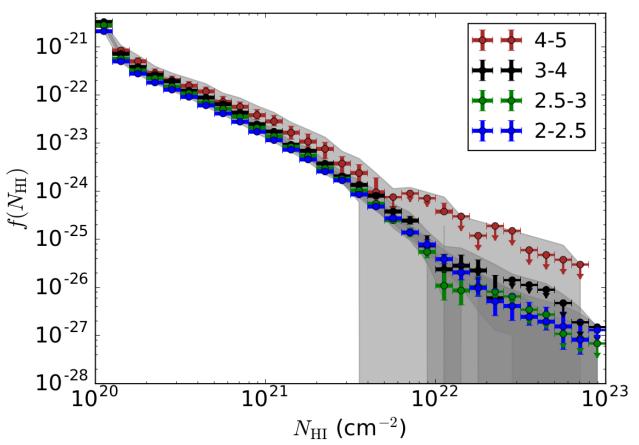
Conclusion

Automated detection of DLAs in spectra

We get a posterior probability

Precise measurement of HI

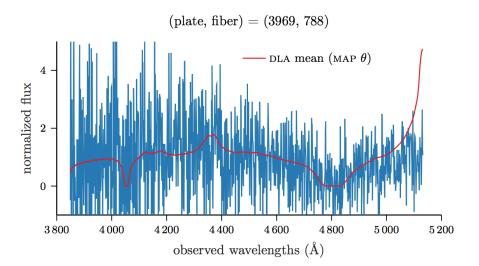
Results: CDDF



- Not much evolution for z<4
- Maybe something at z=4-5

Is this a DLA?

Raw



Is this a DLA?

Raw

(plate, fiber) = (3969, 788) $\frac{4}{3800} = \frac{1}{4000} = \frac{1}{400$

Smoothed

