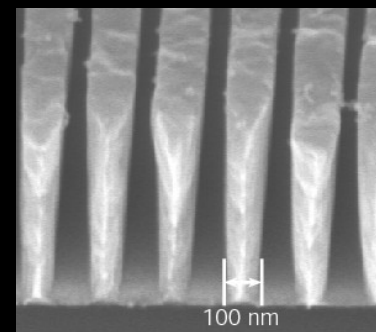


The Three Decades of the Chandra High Energy Transmission Grating (HETG)

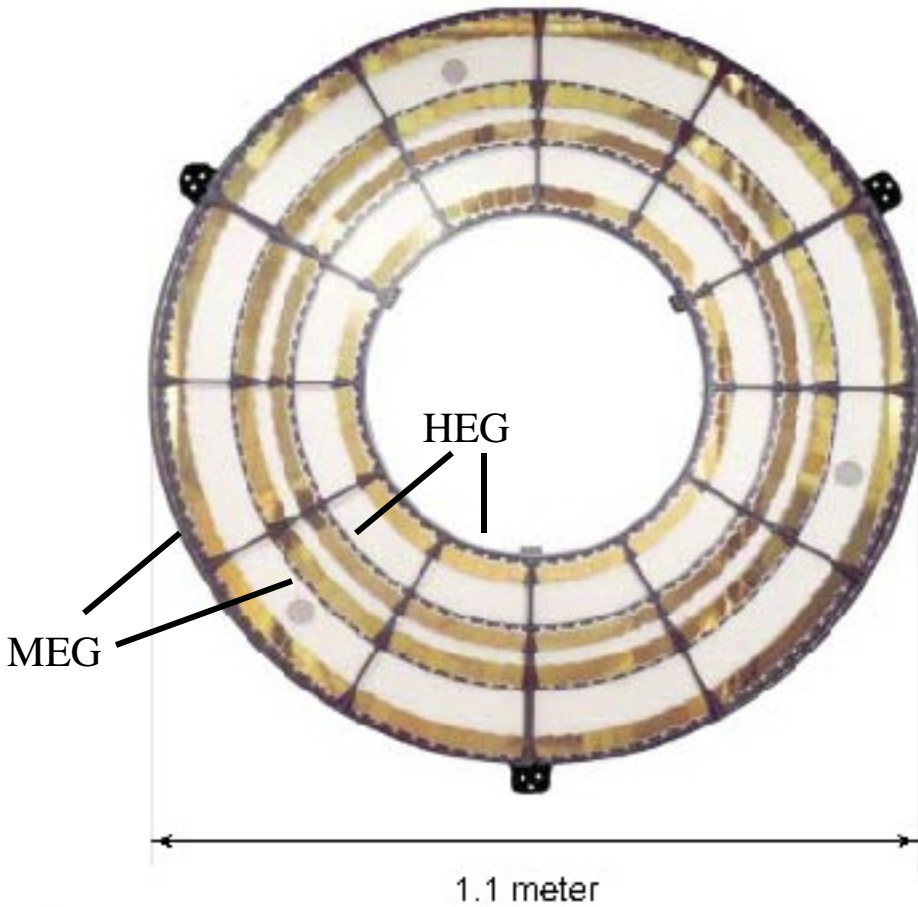


Claude R. Canizares, MIT

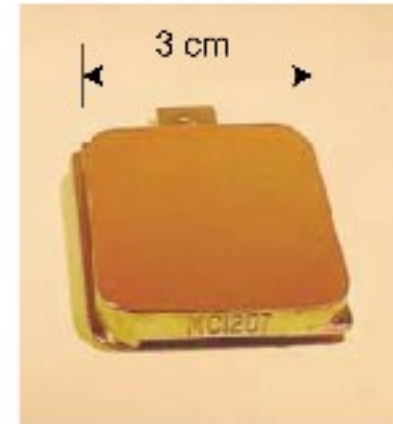
Chandra's First Decade of Discovery
September 2009

NASA Chandra X-ray Observatory High Energy Transmission Grating Spectrometer (HETGS)

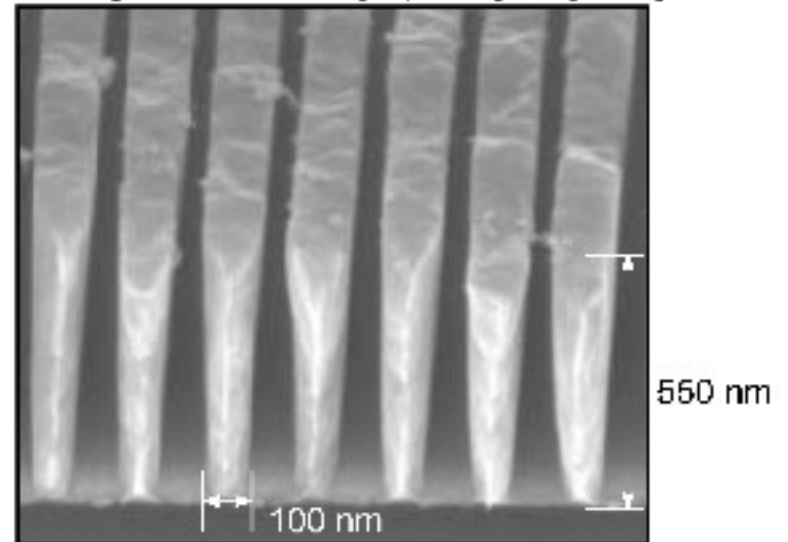
HETGS instrument.



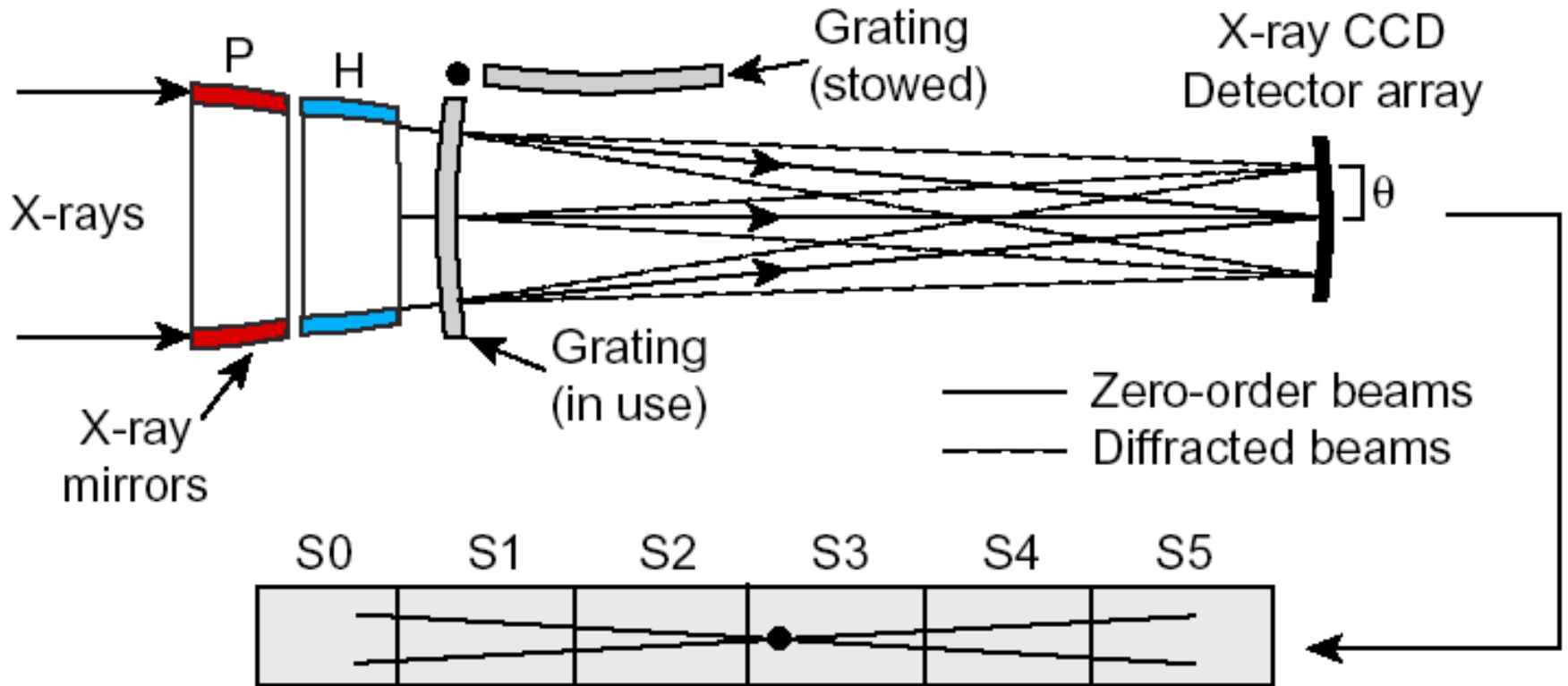
Invar grating frame.



Scanning electron micrograph of gold grating.



Chandra HETG Schematic



Key features needed for an HETG design:

- HIGH SPECTRAL RESOLUTION ($R \sim 1000$):
 - ~ 5000 lpmm ($p = 0.2 \mu\text{m}$, bar thickness = $0.1 \mu\text{m}$)
 - fabrication of hundreds of identical grating elements to tolerances of ~ 100 ppm
- HIGH EFFICIENCY: over 1.5 decades of energy (0.4 - 8 keV)
 - ~ 0.5 - $1.0 \mu\text{m}$ tall ($>5:1$ aspect ratio grating bars)
- HIGH RELIABILITY: gratings rugged enough to withstand launch and space environment

Order of magnitude beyond previous gratings

HETG Timeline: the first decade

1979 Beginning of transmission grating development for AXAF (CRC, Mark Schattenburg with H. I. Smith)

1983 AXAF RFP issued (launch date 1991-2)

- Proposal selected for Phase B study

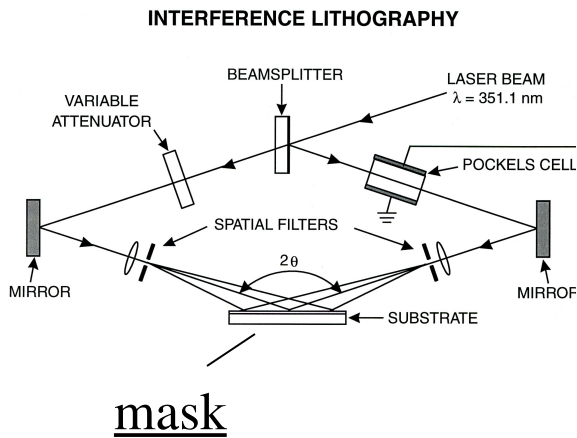
(1986 Challenger disaster)

1988 AXAF approved for “phased new start”

1989 HETG Accepted for AXAF

(launch date 1995-6)

Proposed Fabrication Method: X-ray Lithography

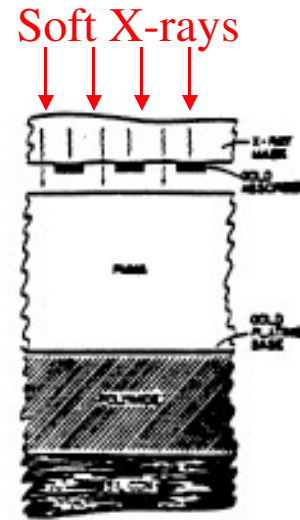


Thin mask of correct period

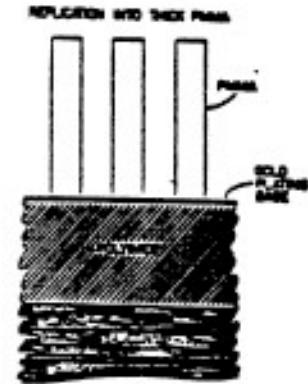
Photoresist

Polyimide support

Replicate a Master (thin grating mask fabricated via UV lithography) into many thick, phased gratings with the same period using X-ray lithography



STEP 1 Step 1



STEP 2 Step 2



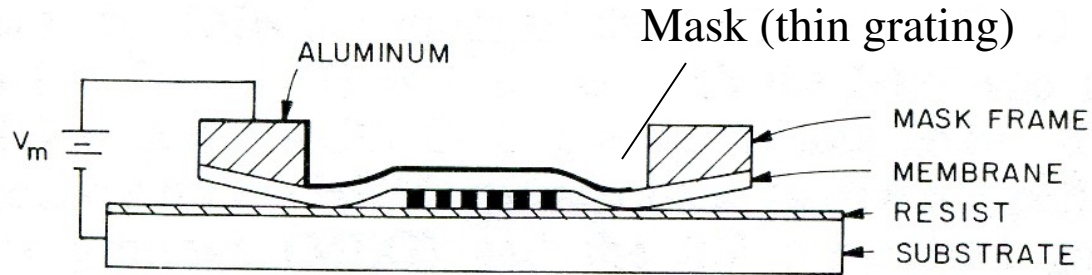
STEP 3 Step 3



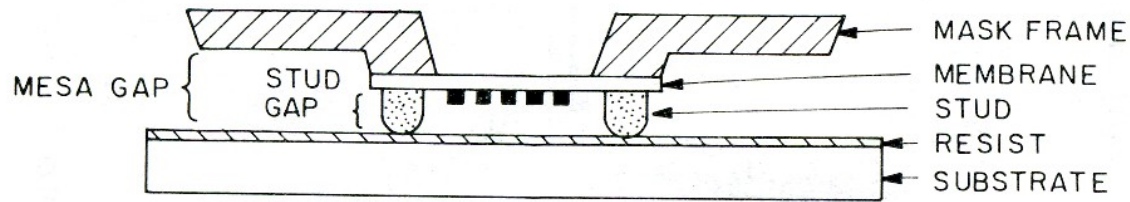
Finished grating

STEP 4 Step 4

Invention of Micro-gap X-ray Nanolithography

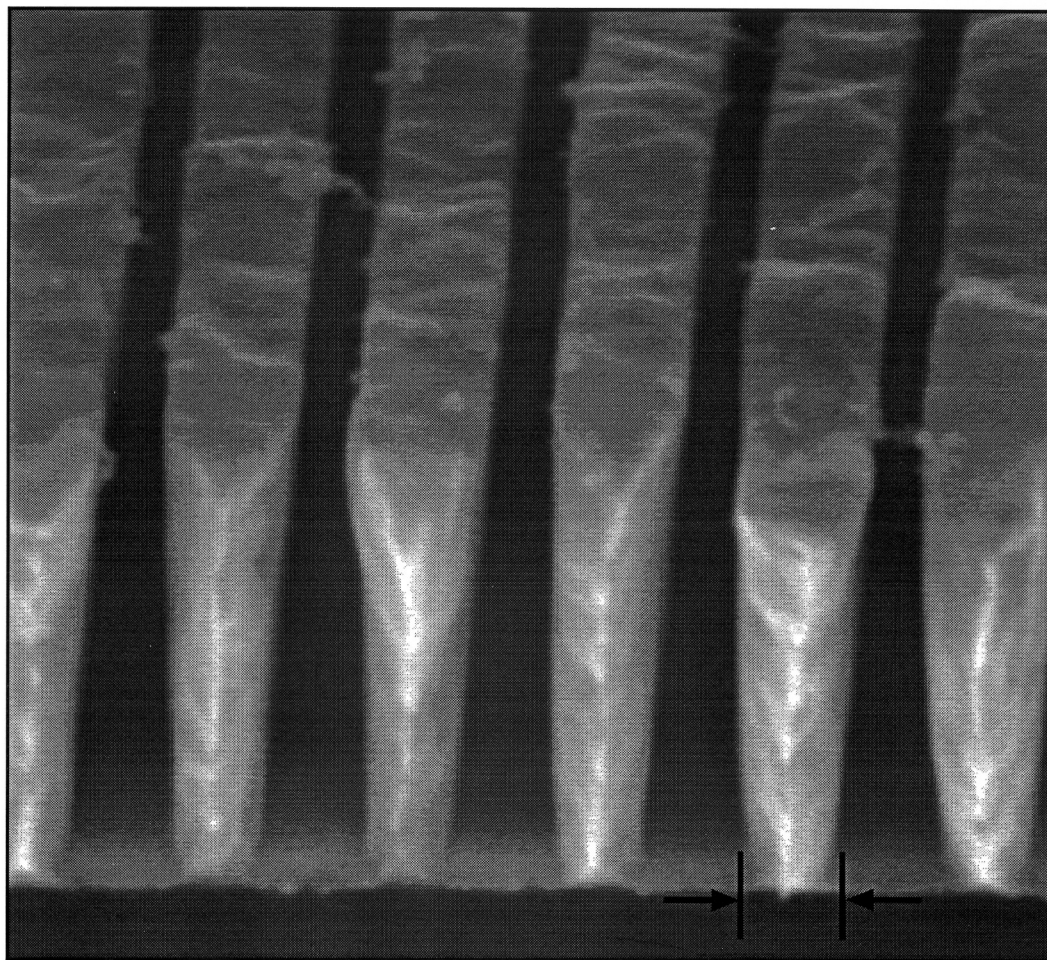


a) CONTACT X-RAY NANOLITHOGRAPHY



b) MICROGAP X-RAY NANOLITHOGRAPHY

AXAF Gold Transmission Grating



0.1 μm

X-ray Lithography Station

Soft X-ray (Cu L line)

Exposure time ~24-36 hrs
per grating!

We (and industry) needed
a high intensity X-ray
machine

Only supplier was
Hampshire Instruments



HETG Timeline: the second decade

- 1989 HETG Accepted for AXAF
- 3 AXAF Restructured to AXAF-I and AXAF-S;
- 4 HETG Systems Requirements Review (SRR)

HETG Timeline: the second decade

- 1989 HETG Accepted for AXAF
- 3 AXAF Restructured to AXAF-I and AXAF-S;
- 4 HETG Systems Requirements Review (SRR)

- 2 Hampshire Instruments ceases operations;
X-ray lithography no longer viable

**HOW COULD WE POSSIBLY BUILD
~700 NEAR-PERFECT GRATINGS ??**

Key breakthrough by Schattenburg:

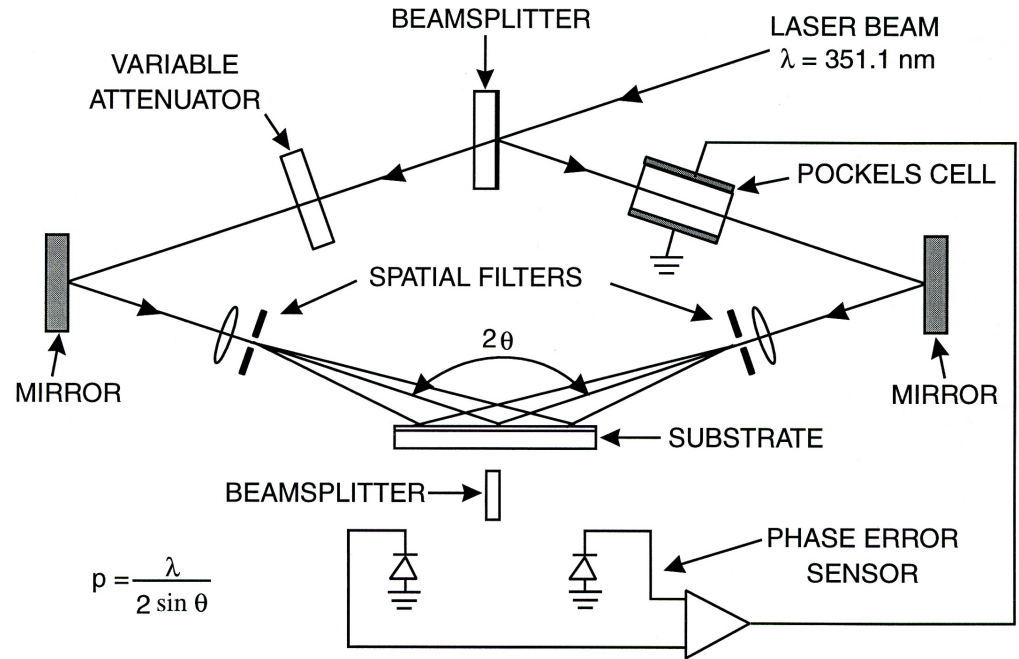
For each exposure, lock UV interference pattern to standard grating (on wafer) using Moire pattern

MLS demonstrates repeatability to less than ~200 ppm (within few weeks!)

Thinks he can achieve high aspect ratio by plasma etching rather than X-ray lithography

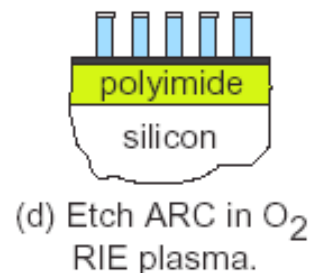
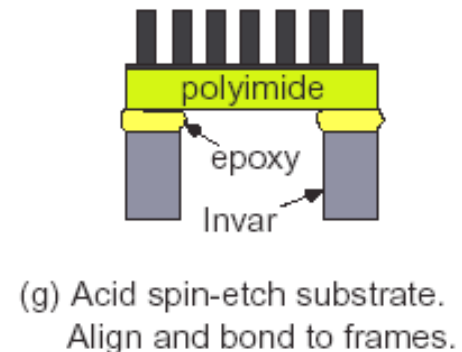
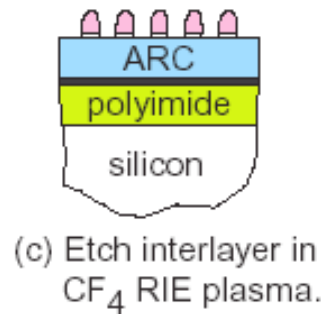
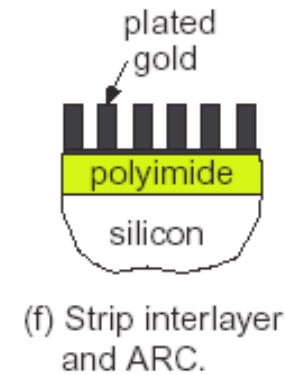
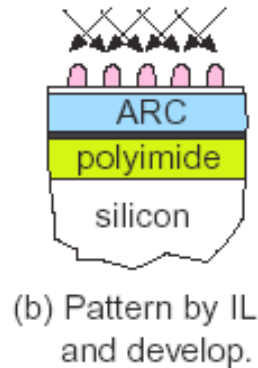
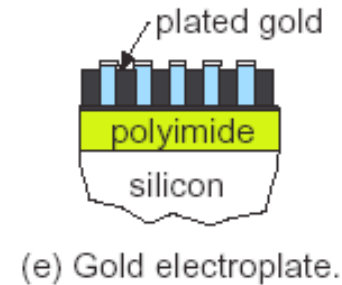
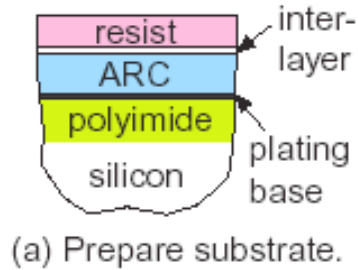
Now the masks have become the gratings

INTERFERENCE LITHOGRAPHY

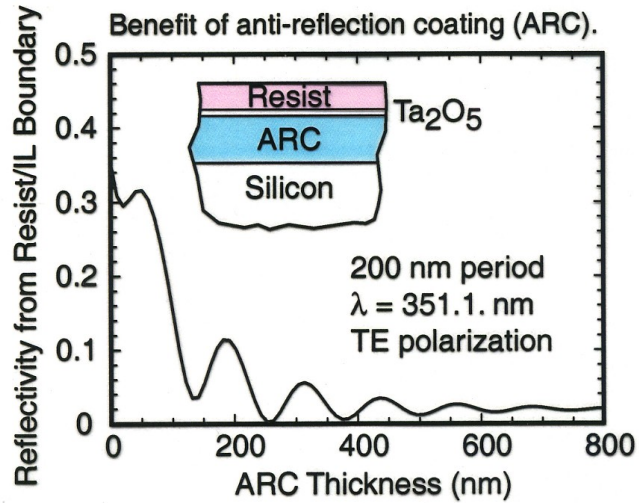


Simplified HETG Fabrication Process

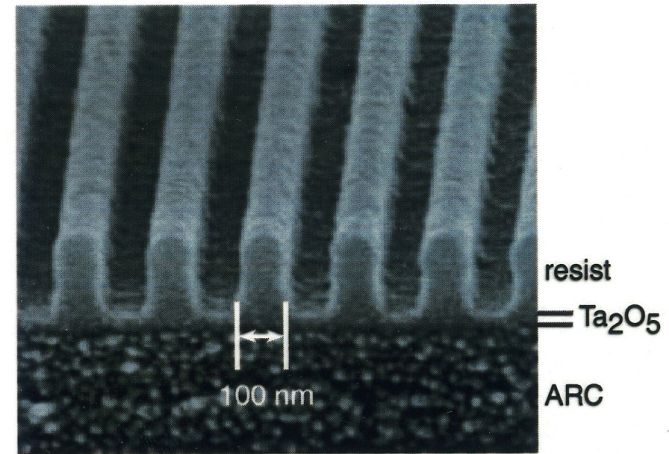
Dozens of technological innovations by Schattenburg and his team; several key patents for processes now widely in use by VLSI industry



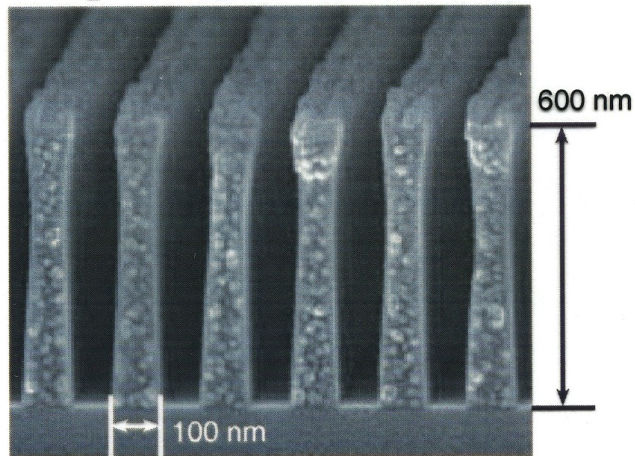
Gold Transmission Grating Fabrication Process



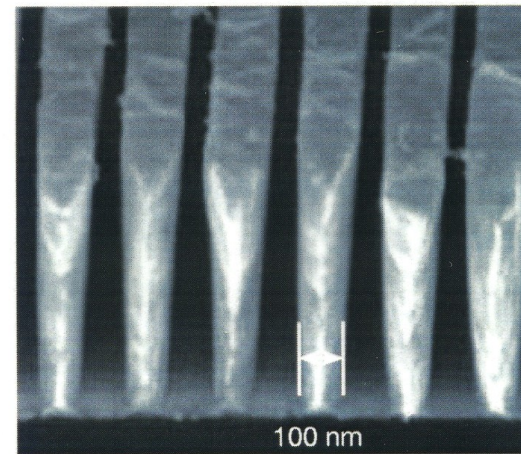
Grating after interference lithography.



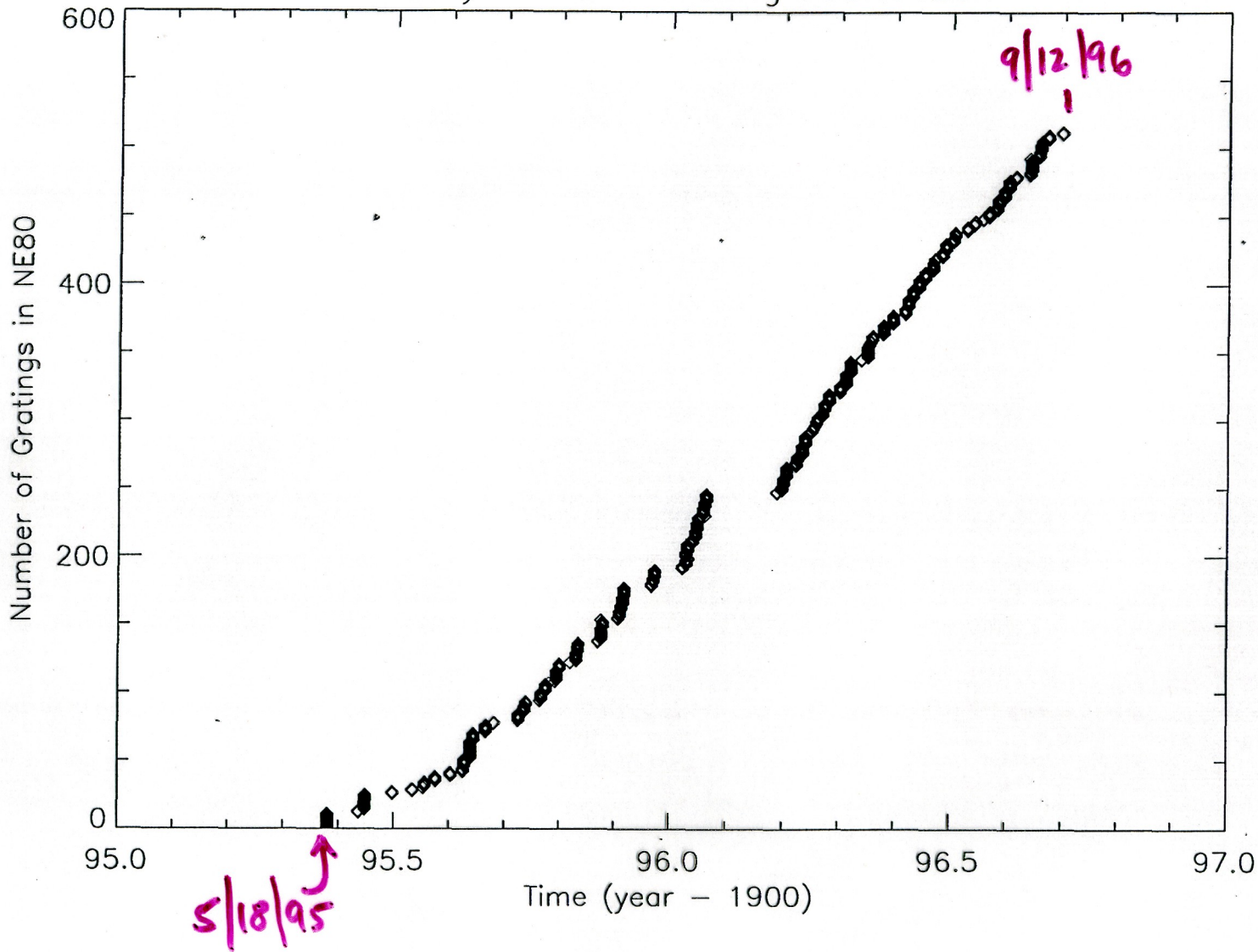
Grating after oxygen plasma RIE of ARC.



Grating after gold plating and resist stripping.



History of HETG Grating Fabrication



HETG Timeline: the second decade

- 1989 HETG Accepted for AXAF
- 3 AXAF Restructured to AXAF-I and AXAF-S;
- 4 HETG Systems Requirements Review (SRR)

- 1993 Hampshire Instruments ceases operations; X-ray lithography abandoned

- 8 Preliminary Design Review (PDR)
- 1995 Critical Design Review CDR

- 1996 Deliver & Calibrate Completed HETG

- 1999 Chandra Launch!

With profound admiration and gratitude for the HETG Team

Fabrication



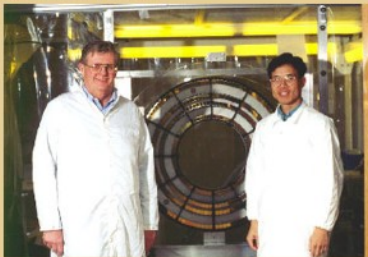
Back row:
Bob Fleming, Mark Schattenburg, Roger Millen, Bob Sisson,
Hank Smith.
Front row:
Rich Aucoin, Jeanne Porter, Jane Prentis, Pat Hindle.

Testing



Dick Elder, Bill Forbes, Bob Laliberte, Ed
Warren, Mike Enwright.

Mechanical



Don Humphries, Chris Pak

Support



Kim Farrell, Dave Breaslau.

Science



Kathy Flanagan, Mike McGuirk, Mark Schattenburg, Claude
Canizares, Dan Dewey, Dick Elder.

Inspiration



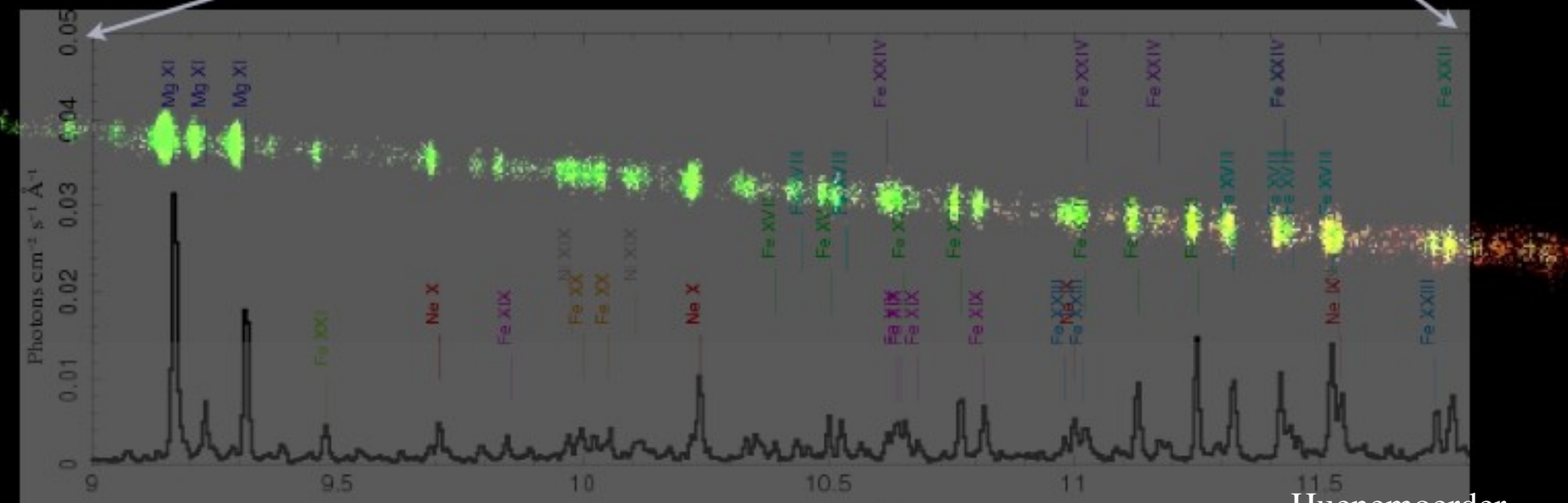
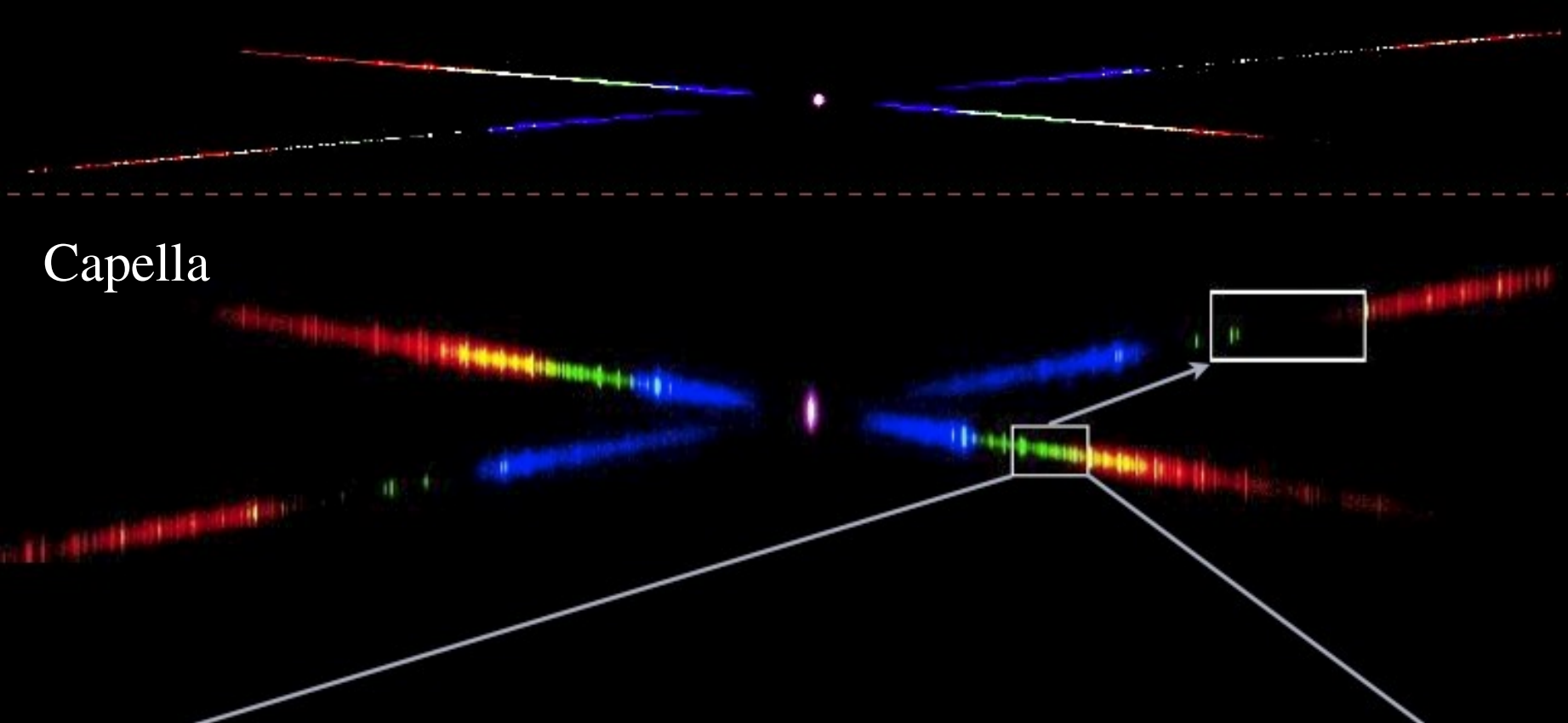
Al Levine, Claude Canizares, Gene Galton,
Angie[for Tom] Markert



Tom Markert

1948-1996

Capella



Hydrogen order

SNR E0102-72

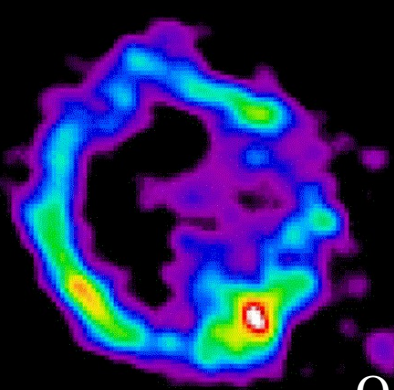
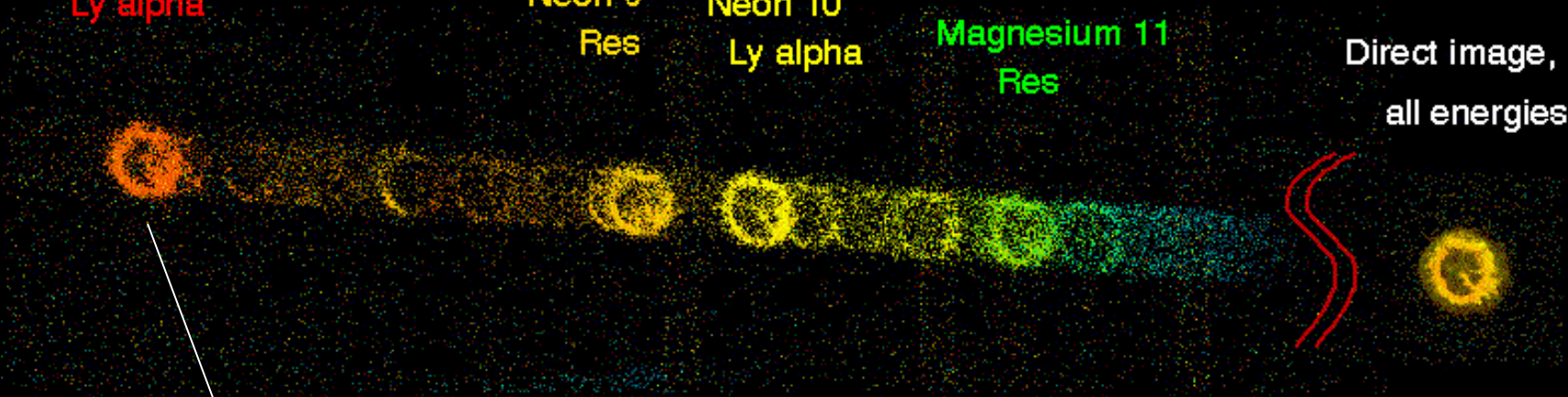
Oxygen 8
Ly alpha

Neon 9
Res

Neon 10
Ly alpha

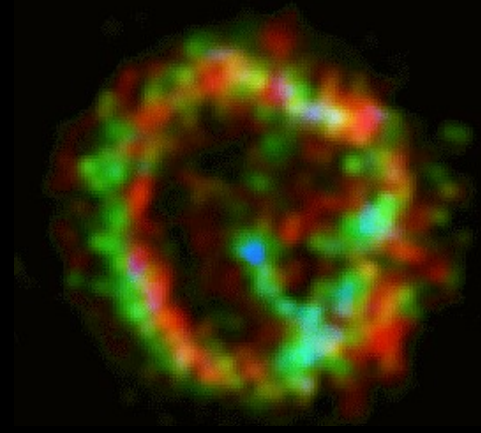
Magnesium 11
Res

Direct image,
all energies



O VIII Ly α

1800 kms^{-1}
900 kms^{-1}
-900 kms^{-1}
-1800 kms^{-1}

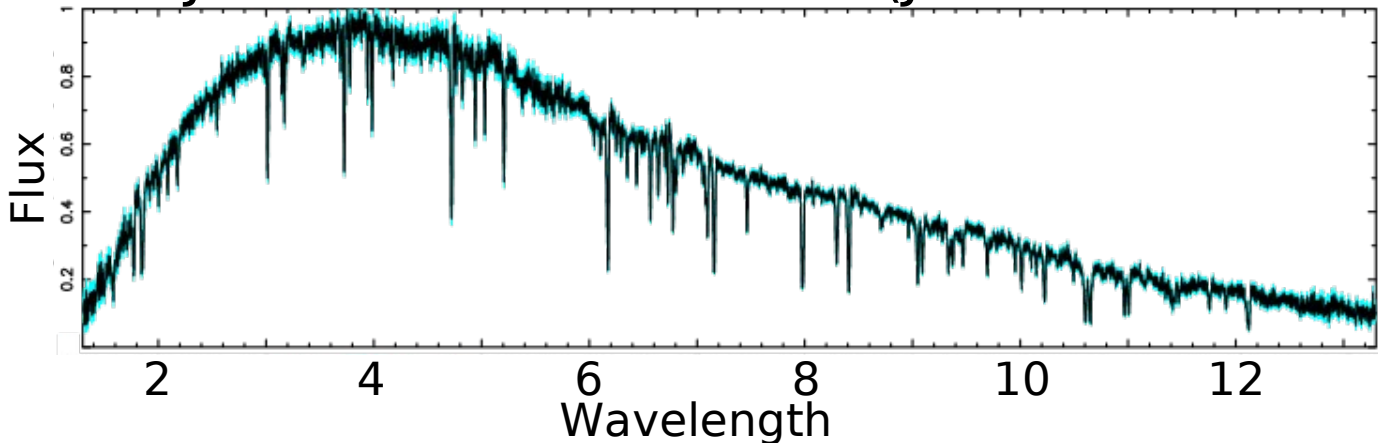


Flanagan et al.

Disk Winds in Black Hole X-ray Binaries

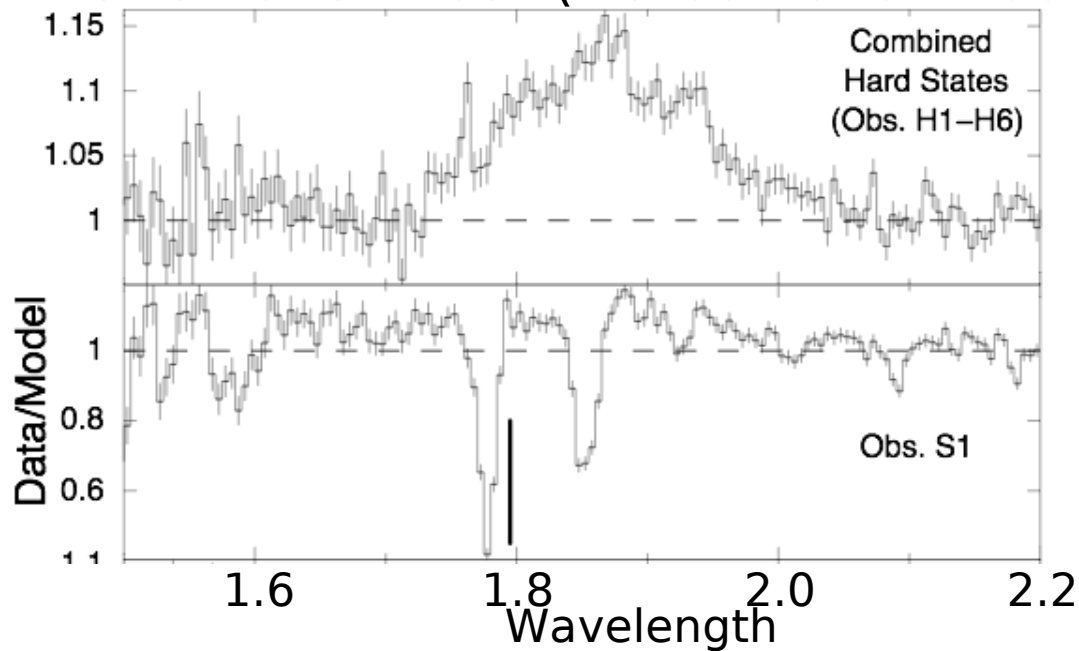
GROJ1655-40

(J. Miller et al. 2008)



Magnetic
Fields Drive
Disk Winds

GRS 1915+105 (Neilsen et al. 2009)



Disk
Winds
Suppress
Jets

914 analysis ready observations

308 distinct objects:

623 HETG/ACIS

91 LETG/ACIS

200 LETG/HRC

Verification & Validation (V&V):

All products manually examined for zeroth order position, extraction region, & confusing sources.

Recently reprocessed to apply CALDB 4.1

TGCat


RESULTS: Found 4 matching extractions

+/-	Links	YobsidA	YobjectA	YinstrumentA	YgratingA	Yra(h:m:s)A	Ydecl(d:m:s)A	Ydate_obs(y-m-d t)A	Yexposure(s)A
<input type="checkbox"/>	0-p-v-s	1926	Vela X-1	ACIS	HETG	09:02:6.838	-40:33:16.920	2001-02-11 21:20:17	84286.8
<input type="checkbox"/>	0-p-v-s	102	Vela X-1	ACIS	HETG	09:02:6.847	-40:33:16.776	2000-04-13 09:57:52	28018.8
<input type="checkbox"/>	0-p-v-s	1927	Vela X-1	ACIS	HETG	09:02:6.845	-40:33:16.812	2001-02-07 09:57:17	30033.5
<input type="checkbox"/>	0-p-v-s	1928	Vela X-1	ACIS	HETG	09:02:6.845	-40:33:16.776	2001-02-05 05:29:55	30261.3

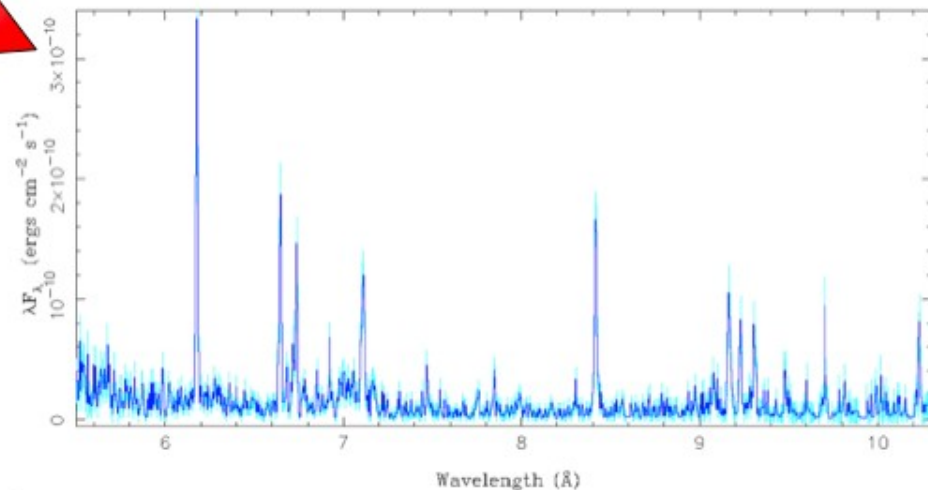
Flux Spectrum (Energy (keV))

press "go" to operate on selections: limit download print

Go Change Columns

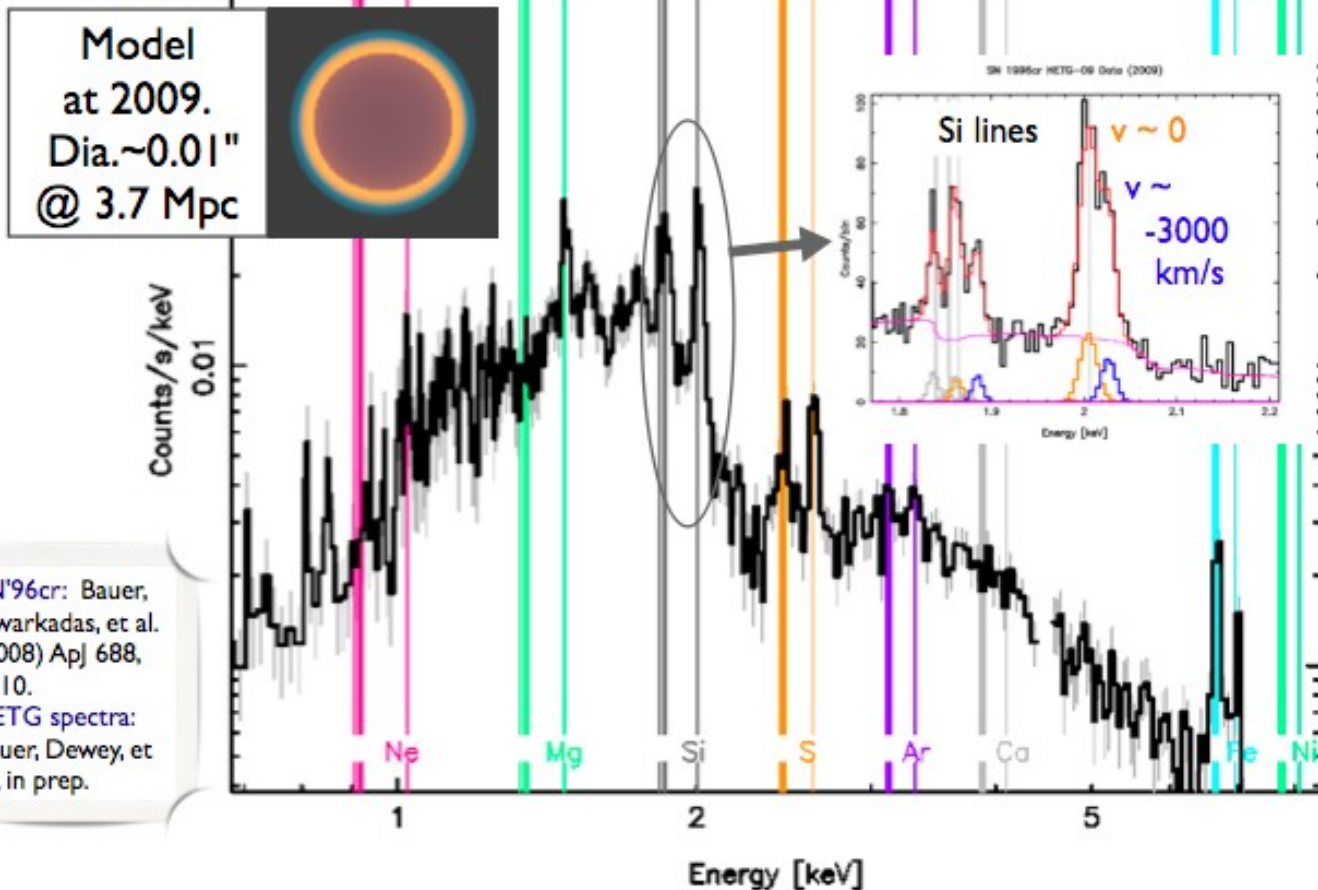


Search, sort, select,
preview, plot,
download, ...



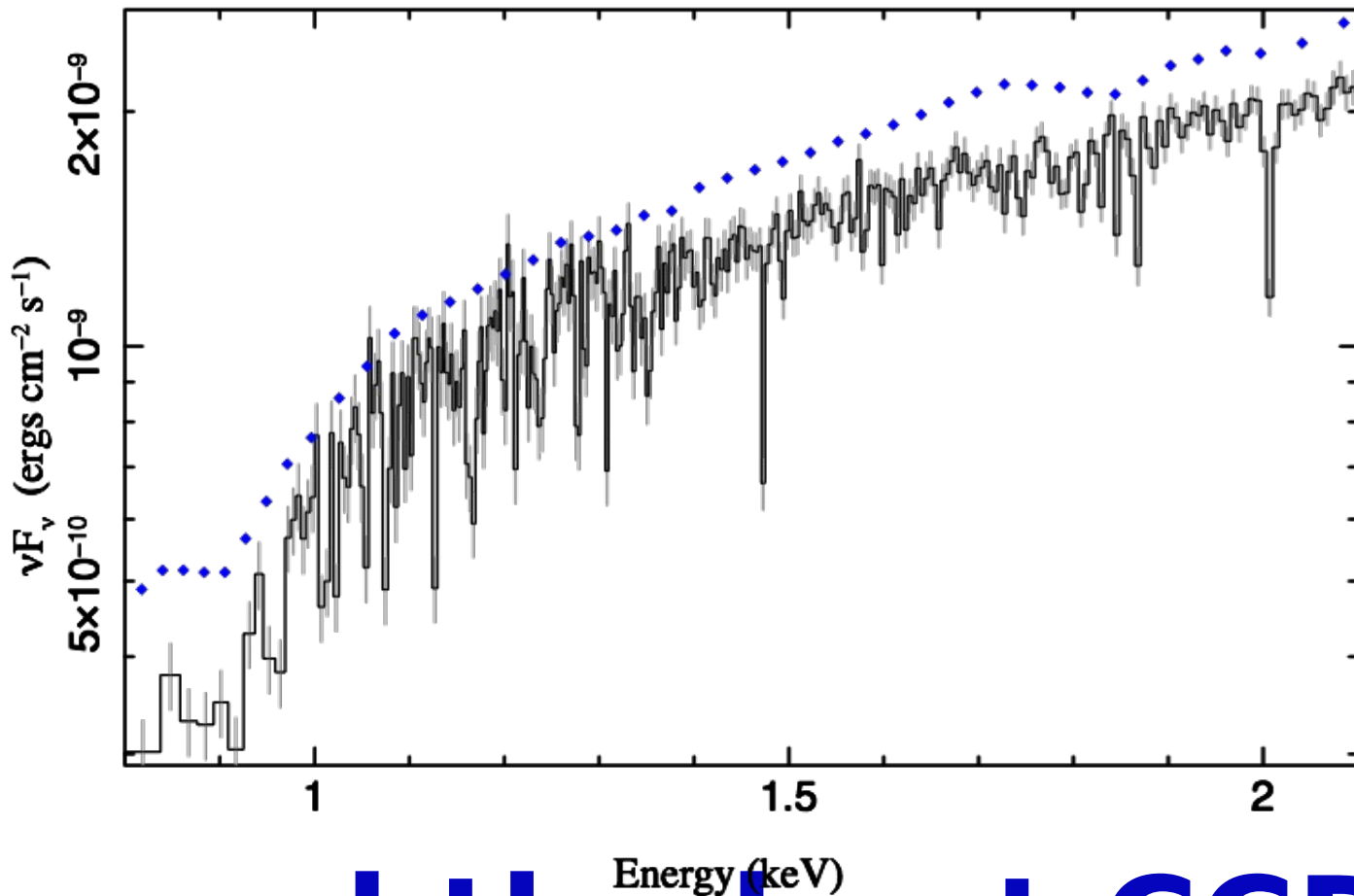
SN 1996cr w/HETG (485 ks) circa 2009 (PI: F.E. Bauer)

Flux increased years after explosion: bubble-shell CSM structure.
Line shapes w/o red-shifted emission: SN core blocks backside.



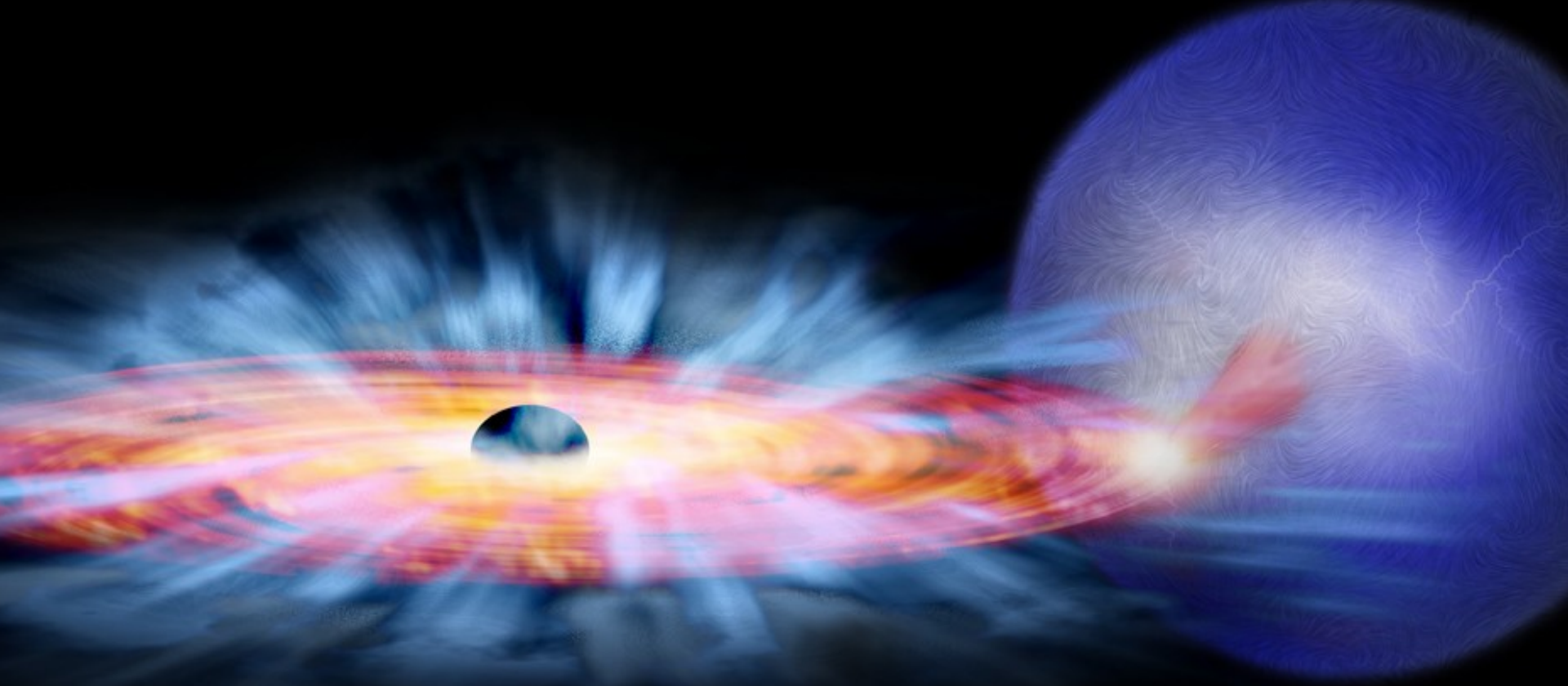
SN'96cr: Bauer, Dwarkadas, et al. (2008) ApJ 688, 1210.
HETG spectra: Bauer, Dewey, et al., in prep.

Cyg X-1 Disk & Atmosphere Viewed with HETG ...



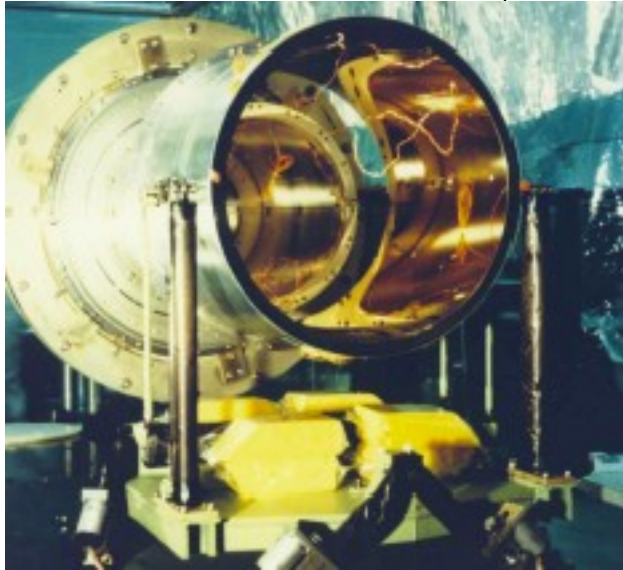
... and the best CCDs
(Suzaku)

Understanding the Atmosphere Crucial for Constraining the Physics

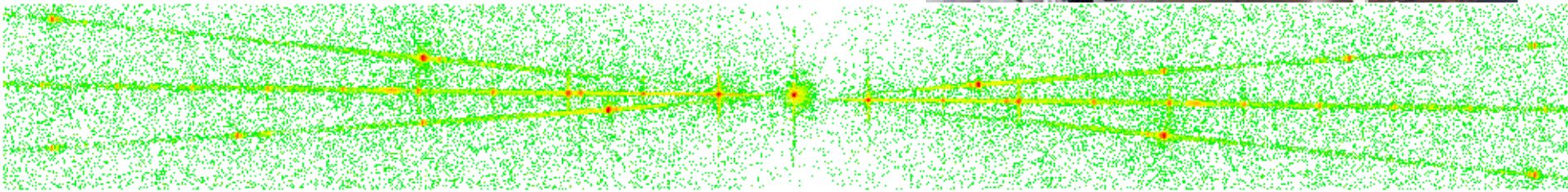
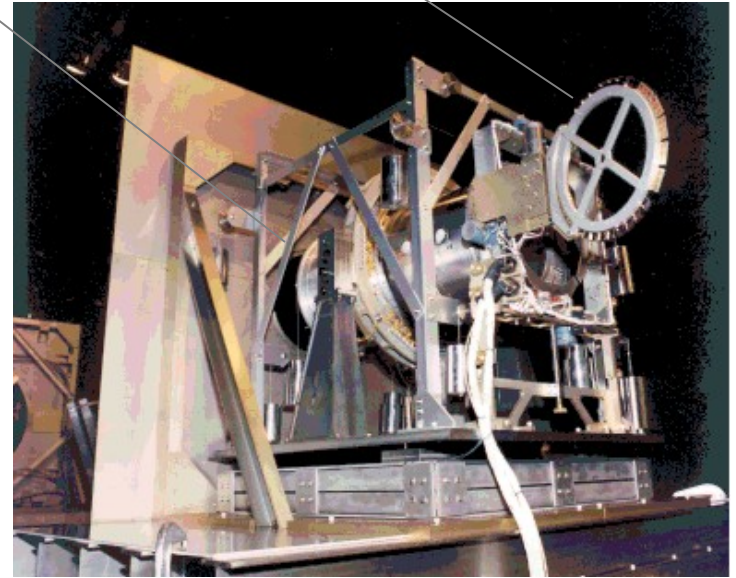


**Magnetic driving? Transitions from
Disks to Jets?**

Test Mirror Assembly (TMA)



Grating facets on wheel (in open position)

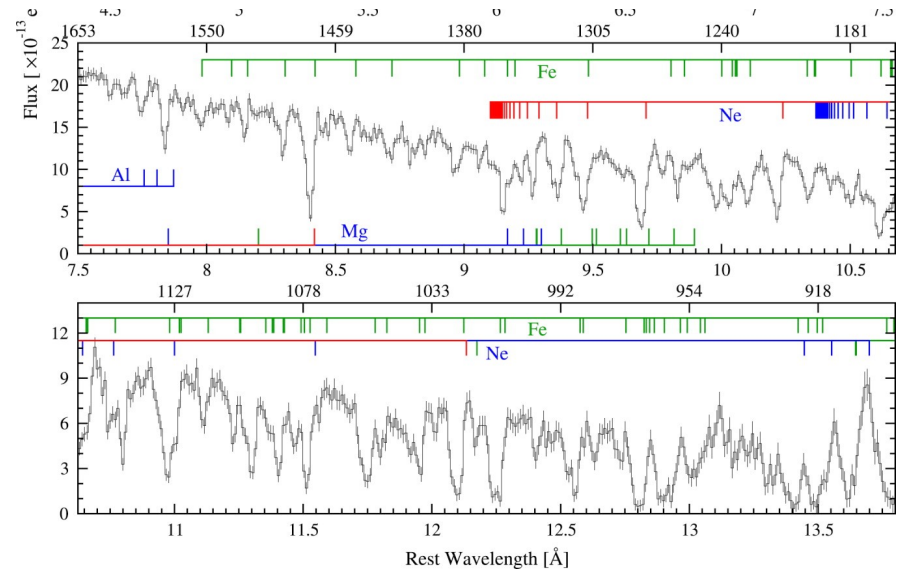
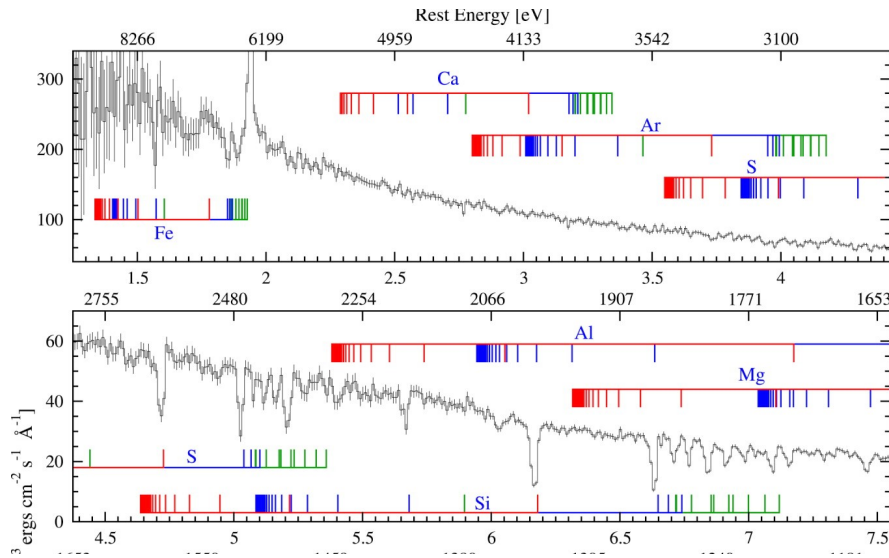


1996 TMA Objective Grating Assembly (TOGA) test at MSFC/XRCF

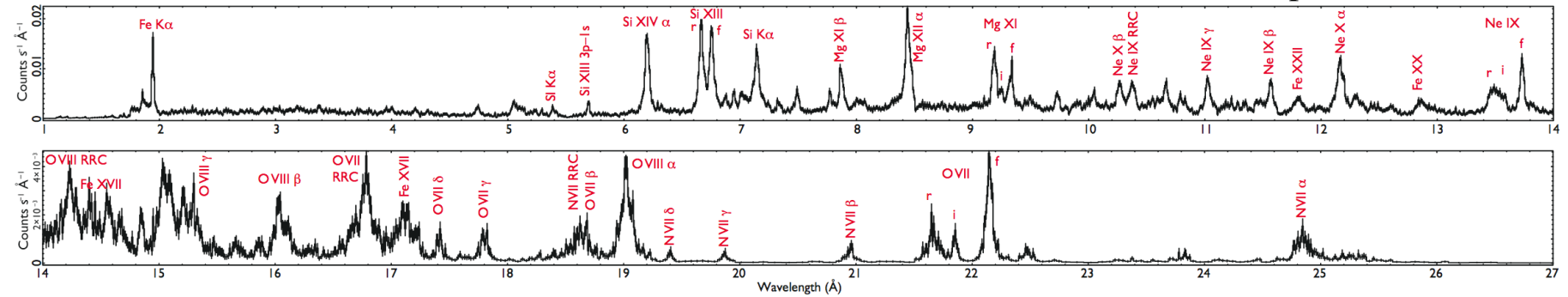
Twelve years after initial proposal, the first real evidence that grating assembly would perform as expected!!

“[expletive deleted]!!! I might even use the gratings!” --- Leon van Speybroeck

Environments of Active Galactic Nuclei



NGC 3783 Kaspi et al. 2002



NGC 1068 Evans et al. 2009