

High Energy Transmission Grating (HETG) Spectra in Continuous Clocking Mode

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Scope:

- Why use CC-mode for HETG observations
- HETG CC-Mode Observations of X-Ray Binaries
- Trap maps for ACIS CC-mode Calibration
- Problems we recognized so far: OSIPs, Si K, CTI......
- TE and CC-mode with all Flight Grades Transmitted
- Observations of very bright sources
- Future Actions and Recommendations



Why use CC-mode for HETG observations



-- Timing studies-- Pileup mitigation



-- Continuum reconstruction: New pileup model for bright sources (Davis 2008)



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X-Ray Binaries observed in HETG CC-Mode



source	Yoff	Z_{SIM}	rate
	arcmin	mm	cts/s
Sco X-1	0.000	-15.000	1430.0
GX 5-1	0.167	-11.300	505.0
GRO J1655-40	0.330	-7.500	492.0
H 1743-322	0.330	-7.500	492.0
GRS 1915+105	-1.330	-4.000	366.0
Cvg X-2	1.167	-6.140	331.0
Cvg X-1	-1.330	-4.000	266.0
GX 349-2	0.167	-7.500	225.0
XTE J1650-500	-1.330	-4.000	245.0
XTE J1550-564	-1.330	10.000	200.0
4U 1812-12	0.000	-6.040	182.3
4U 1820-30	1.167	-6.140	180.0
4U 1636-53	0.167	-6.140	165.0
4U 1630-472	0.330	-7.500	160.0
GX 339-4	0.330	-7.500	160.0
4U 1735-44	0.167	-7.500	150.0
XTE J1817-330	0.330	-4.500	200.0
PSR B0833-45	0.000	-3.000	88.7
GX 9+1	0.000	-3.000	86.2
4U 1705-44	0.330	-7.490	70.0
GX 340+0	0.330	-7.490	60.0
GX 17+2	0.330	-4.000	58.1
Cyg X-3	0.167	-6.800	40.0
4U 1746-37	-0.330	-4.000	24.9
4U 1728-34	0.167	-4.000	20.3
Her X-1	0.330	-4.000	19.7
SAX J1808.4-3658	0.330	-4.000	20.0
GRS 1747-312	-0.330	-6.140	12.1
RAPID BURSTER	0.000	0.000	5.6
GX 1+4	-0.330	-5.860	4.0
CRAB PULSAR-CC	0.000	0.000	2.0
4U 1323-619	-0.330	-4.000	1.1

35 sources observed so far mostly X-ray Binaries total exposure: ~ 2 Msec

Observe sources the instruments were NOT designed for

Primary application of cc-mode was for maximum time resolution

The use of Graded mode heavily restricts instrument level calibration

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Current status of ACIS CC-mode Calibratio



Development of trailing pix. algorithm using ECS data in cc-mode:





Current status of ACIS CC-mode Calibratio



Development of trailing pix. algorithm using ECS data in cc-mode:





128

16

4

32

8

1

2

charge

transfer

charge transfer

ACIS flight grades (fltgrade) record wated of the 8 pixels surrounding a local maximum have charge in them, above the split threshold.

Many CC mode events have charge in the upper pixe. (flugrade 64) or both upper and lower pixel (fltgrade 66)

Seven flight grades, including 66 and 255, are discarded on board. ACIS fltgrade 66



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Frame Drops in Bright X-Ray Binaries





Use Graded Mode to minimize diff. exosure loss via frame drops in the telemetry stream

loose flight grades through trailing charge for good











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ACIS properties in HETG CC-Mode Spectra





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How to Address the Problem



What has been done so far:

-find all descrete changes in the effective area.
---> OSIP problems, Si K edge problem
-find mismatches between HETG flux dispersions
---> Diagnose CTI, CCD, Grade related problems
-cross-correlate observations with other instruments, i.e. RXTE, Suzaku, etc...

----> Heavyly model dependent, source flux dependent, dependent on external instruments calibrations

What is currently done::

apply current versions of CTI corrections to observations preformed in cc-faint mode





Obs Sequence: 4U1957-11 ----> 10ks TE ----> 20ks CC ----> 10ks TE



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Obs Sequence: 4U1957-11 ----> 10ks TE ----> 20ks CC ----> 10ks TE





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Wednesday, September 30, 2009









CC-mode HETG, MEG +1 TE compare before vs after









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Wednesday, September 30, 2009









CC-mode HETG, MEG -1 TE compare before vs CC with CTI

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CC-mode HETG, MEG -1 TE compare before vs CC with CTI









CC-mode HETG, MEG +1 TE compare before vs CC with NCTIW66









CC-mode HETG, MEG -1 TE compare before vs CC with NCTIW66

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CC-mode HETG, MEG +1 TE compare before vs CC with NCTIW66

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Rate Dependent Issues:



Si K edge problem: > 50 mCrab:

- depression between 6.2 -- 6.74 A (Si K edge) in BI devices
- edge mismatch (shift) in FI devices

HEG +/- vs MEG +/- mismatch:

- separate count rate dependence from low rate CTI

What is currently done::

evaluate Si K edge depth vs. source rate





Bright X-ray sources with $N_h < 3 \times 10^{21} \text{ cm}^{-2}$



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The Si K Edge in CC-Graded Mode: HEG vs MEG

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CC-Graded Mode: MEG vs HEG

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All observations should use "Faint" NOT "Graded"

- include non-standard flight grades, i.e. g66
- consider creating new G+ mode, which includes g66
- compensate anticipated loss of data through telemetry saturation via increased exposure rather on-board grading Observations ::

-moderately brightsourese in cc-F mode: 4U 1957-11 -- 50 mCrab (CAL, cycle 10, 20 ks) GX 13+1 -- 90 mCrab (GTO, cycle 11, 40 ks) GX 17+2 -- 150 mCrab (GO, cycle 11, 10 ks) -- 250 mCrab (CAL, cycle 10, 20 ks) Calibration Workshop, Cambridge MA, Sept. 21, 2009