**10 Billion** Years of Galaxy Cluster **Evolution** with Chandra

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### **Galaxy Cluster Evolution**

The timing of many major milestones in cluster formation/ evolution remain uncertain

Need wellselected samples of galaxy clusters at z > 1 to understand evolution



# **Combining SZ + X-ray Data: Selection**

- mm-wave surveys probing the Sunyaev Zel'dovich effect provide our best opportunity to study massive clusters in the early universe
  - Extend to much higher redshift than X-ray surveys
  - Significantly less biased to cluster dynamical state than X-ray



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Our strategy (2010-2024): Follow up most massive SPT-selected clusters at 0.3 < z < 1.7 with *Chandra* 







#### Michael McDonald





X-ray: NASA/CXC/SAO; Optical: NASA/ESA/STScI; IR: NASA/ESA/CSA/STScI/Milisavljevic et al., NASA/JPL/CalTech; Image Processing: NASA/CXC/SAO/J. Schmidt and K. Arcand

### Flores et al. 2021

(XMM-Newton + Chandra analysis of the most distant SZ-selected clusters at  $z \sim 1.5$ )



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Credit: NASA/CXC/GSFC/S.A.Walker, et al.

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Credit: X-ray: NASA/CXC/Univ. Waterloo/B.McNamara; Optical: NASA/ESA/STScl/Univ. Waterloo/B.McNamara; Radio: NRAO/Ohio Univ./L.Birzan et al.



### **Prospects for the Future: Larger Samples**





### **Prospects for the Future: AXIS Probe**



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Directly Probe Radio-Mode AGN Feedback to z ~ 2 Constrain ICM Metal Enrichment to z ~ 3





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# Summary

- The combination of mm-wave selection and X-ray follow-up with Chandra has enabled evolutionary studies of the ICM over 10 Gyr
  - See also talks by Steve Allen, Michael Reefe, Michael Calzadilla
- In the next 5-10 years we will discover galaxy clusters shortly after their birth (z ~ 3) with SZ surveys.
  - Understanding the physics that governs them requires new methodology/technology



### **Prospects for the Future: Multi-Observatory**



### **Prospects for the Future: Multi-Wavelength**

