## Quasars as high redshift standard candles

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# quasars



$$\log(L_X) = \alpha \log(L_{UV}) + \beta$$

$$L_v (2 \text{ keV}) \qquad \qquad L_v (2500 \text{ Å})$$

$$\alpha = 0.6$$



Risaliti & Lusso 2019, Lusso et al. 2020





# quasars



$$\log(L_{X}) = \alpha \log(L_{UV}) + \beta$$

$$L_{\nu} (2 \text{ keV}) \qquad \qquad L_{\nu} (2500 \text{ Å})$$

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## $\log [D_{L}(z)] = \log(F_{X}) - \alpha \log (F_{UV}) + \beta'$







Is the X-ray to UV relation constant with redshift ?

Are there selection effects in the sample ? Are the quasars in the cosmological sample really *average* quasars or are we "cherry-picking" the ones we like?

Are there systematic effects in the flux measurements ? Can reddening (extinction, host galaxy contamination) affect our sample?



Detailed spectral analysis (dust reddening, etc...)

#### OPTCAL/UV

#### X-ray



Signorini et al. 2023a

Nardini et al. 2019, Signorini et al. 2023

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Quasar spectral properties as a function of redshift: analysis



Trefoloni et al. 2024





Comparison with SN Ia in the common redshift









Relation slope vs. redshift



Lusso+20: 2,400 quasars: 1,600 XMM + 800 Chandra

Chandra@25



Relation slope vs. redshift



2024: 2,000 quasars: <u>All Chandra (from CSC 2</u>

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Lower dispersion with better data and better analysis

130 sources,  $\delta = 0.12$ 

30 sources,  $\delta = 0.08$ 



Sacchi et al. 2022





Quasars as standard candles:

- No satisfactory physical model for the disk-corona connection
- Dispersion still higher than that of supernovae

### BUT

- Slope not evolving with redshift
- Spectral properties non evolving with redshift
- Perfect match with supernovae in the common redshift range
- Dispersion decreasing with better flux measurements
- Dispersion entirely explained with "external" effects (inclination, variability)

