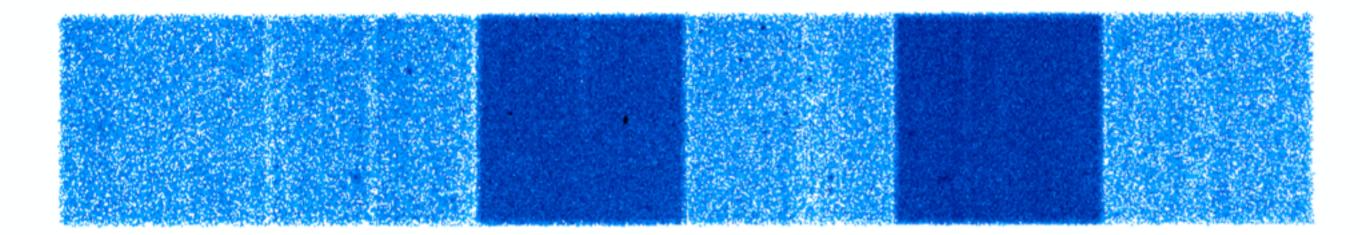
Rodolfo Montez Jr.

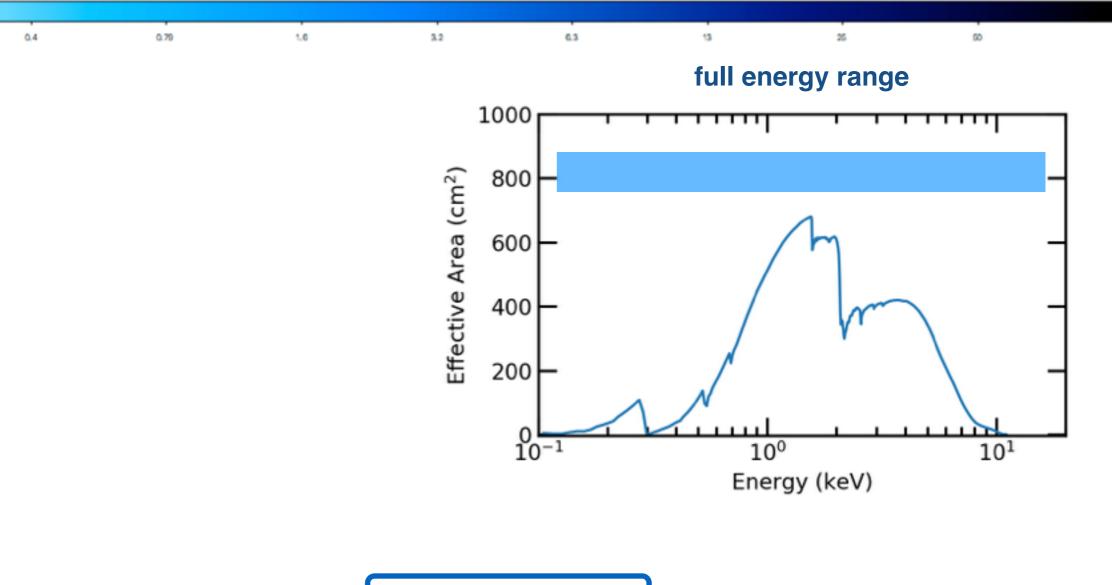
#### dmcopy "acis\_repro\_evt2.fits[...][...]" new\_file.fits

# your new best friend data model filters dmcopy "acis\_repro\_evt2.fits[...][...]" new\_file.fits event file output file

but it could be a fits image, ascii table, etc.

some filters will preserve the event list some will destroy it, options can give you more control

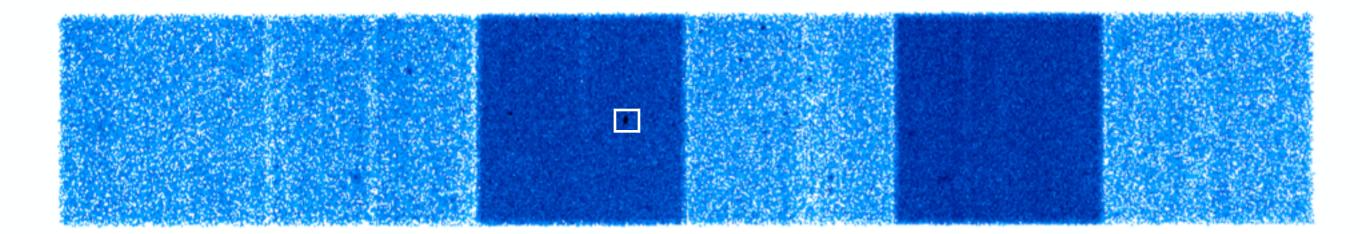


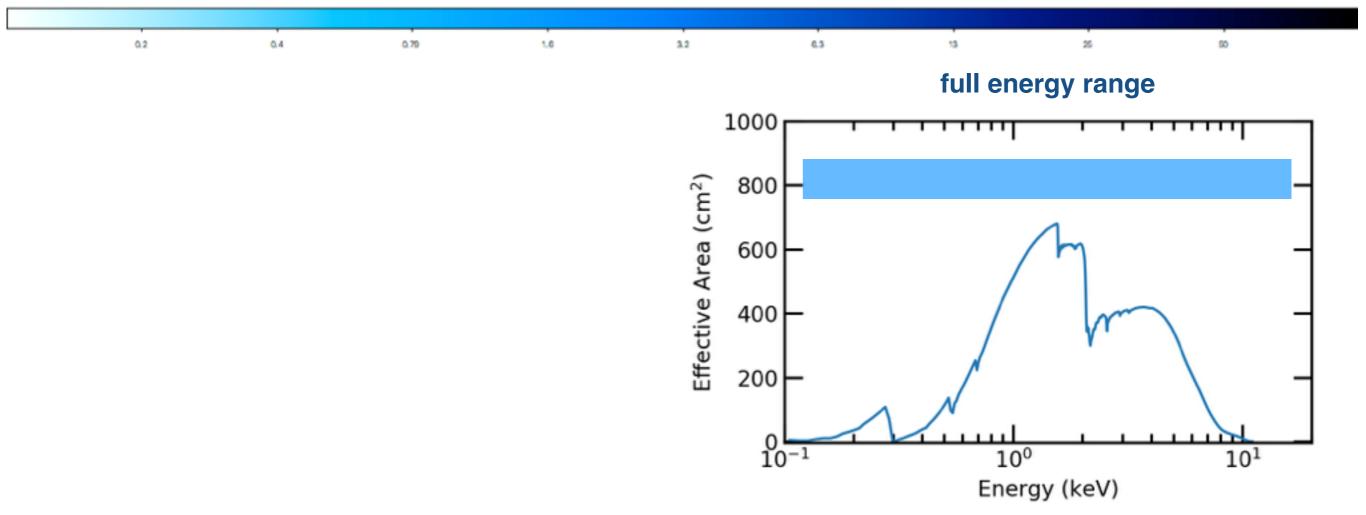


0.2

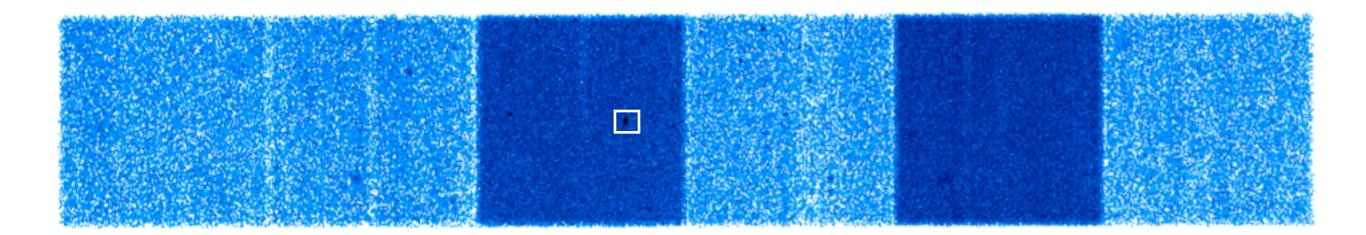
here is the bin filter used in dmcopy

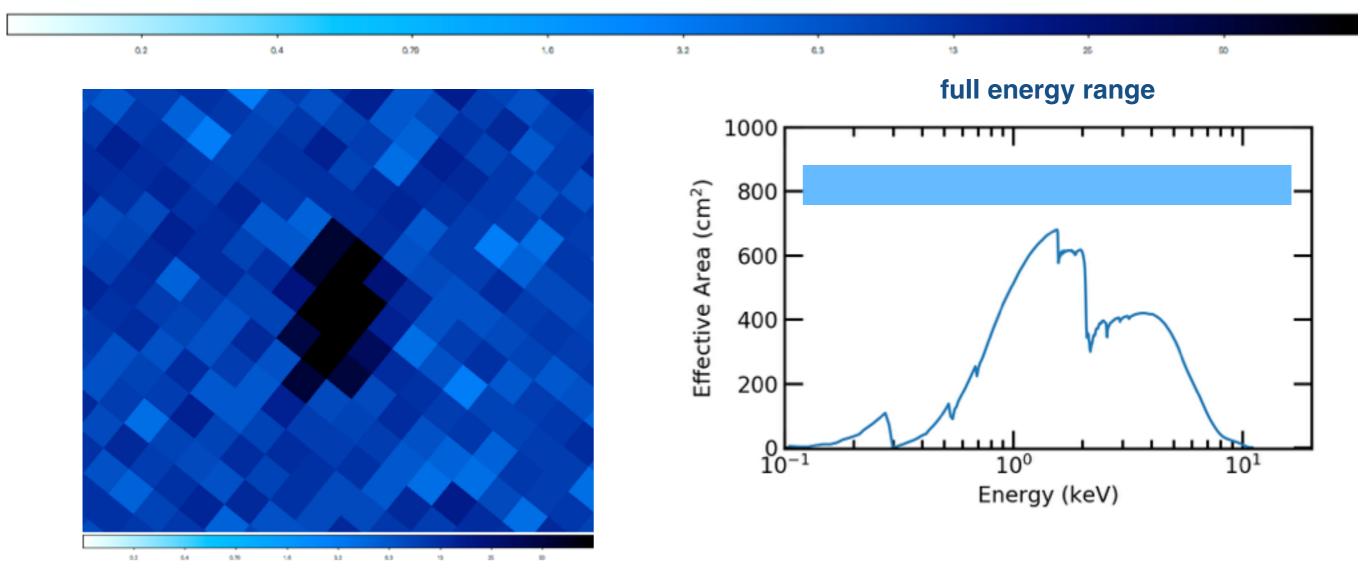
[bin x=::8,y=::8]



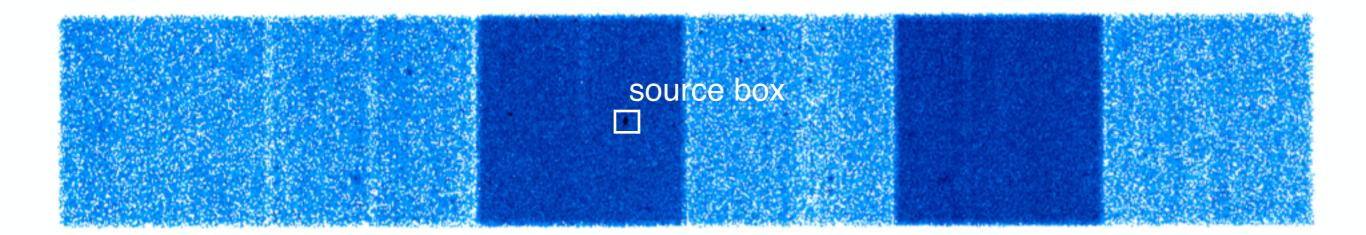


[bin x=::8,y=::8]





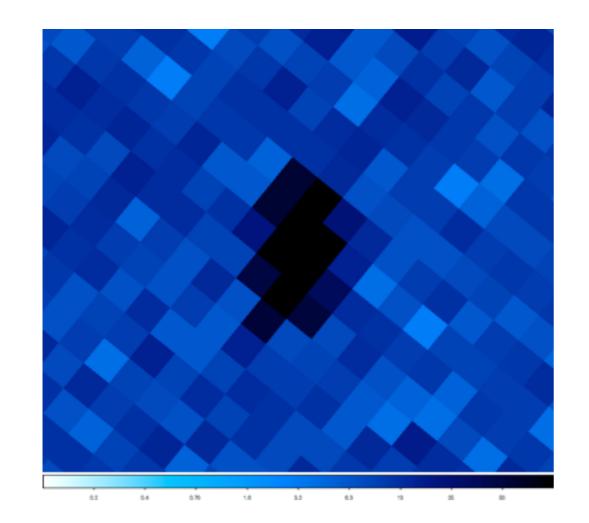
[bin x=::8,y=::8]



3.2

6.3

1.0



0.79

0.4

0.2

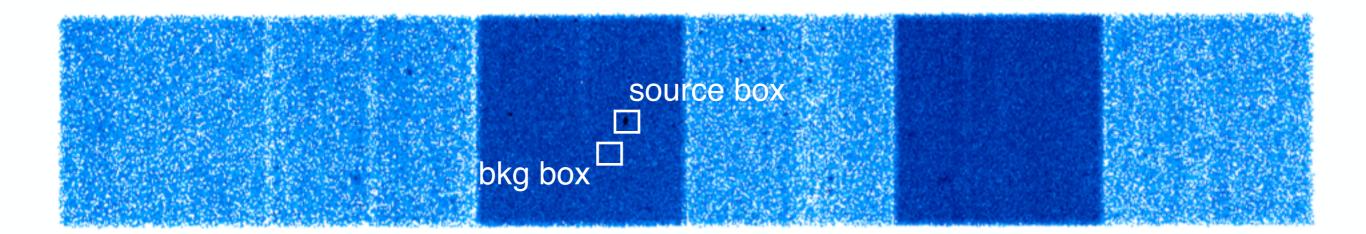
full energy range

25

13

[bin x=::8,y=::8]

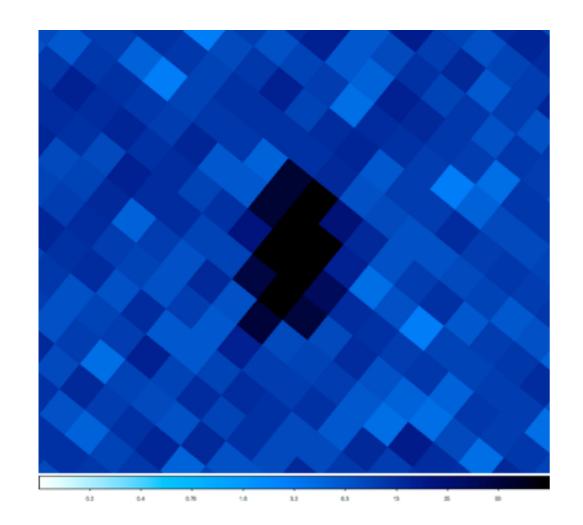
50



3.2

6.3

1.0



0.70

0.4

0.2

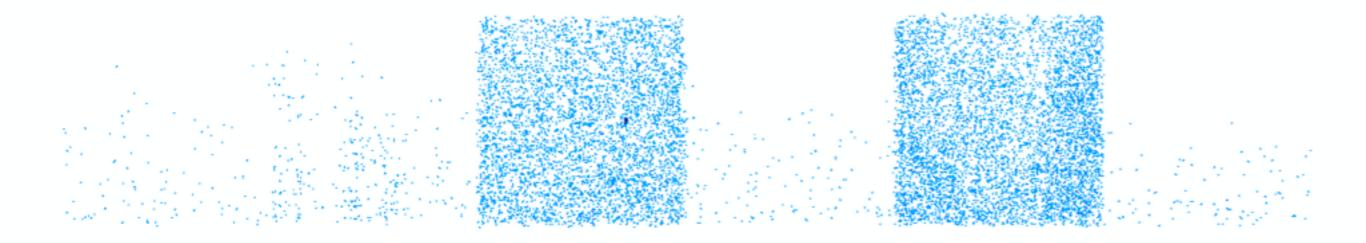
full energy range

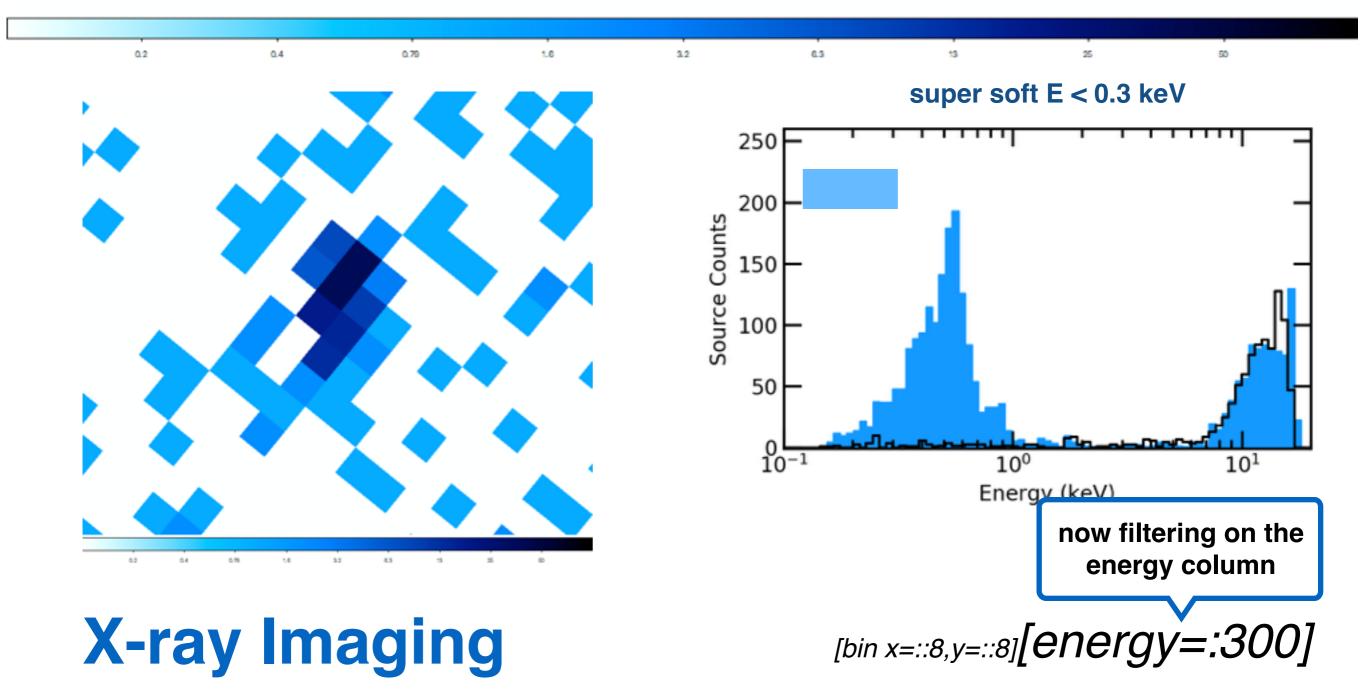
25

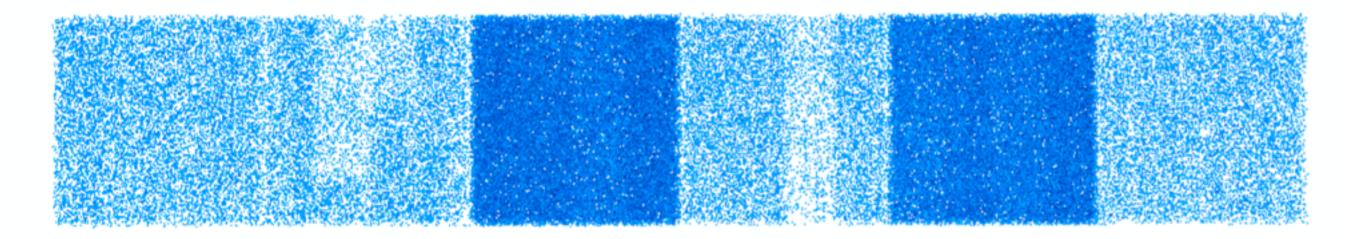
13

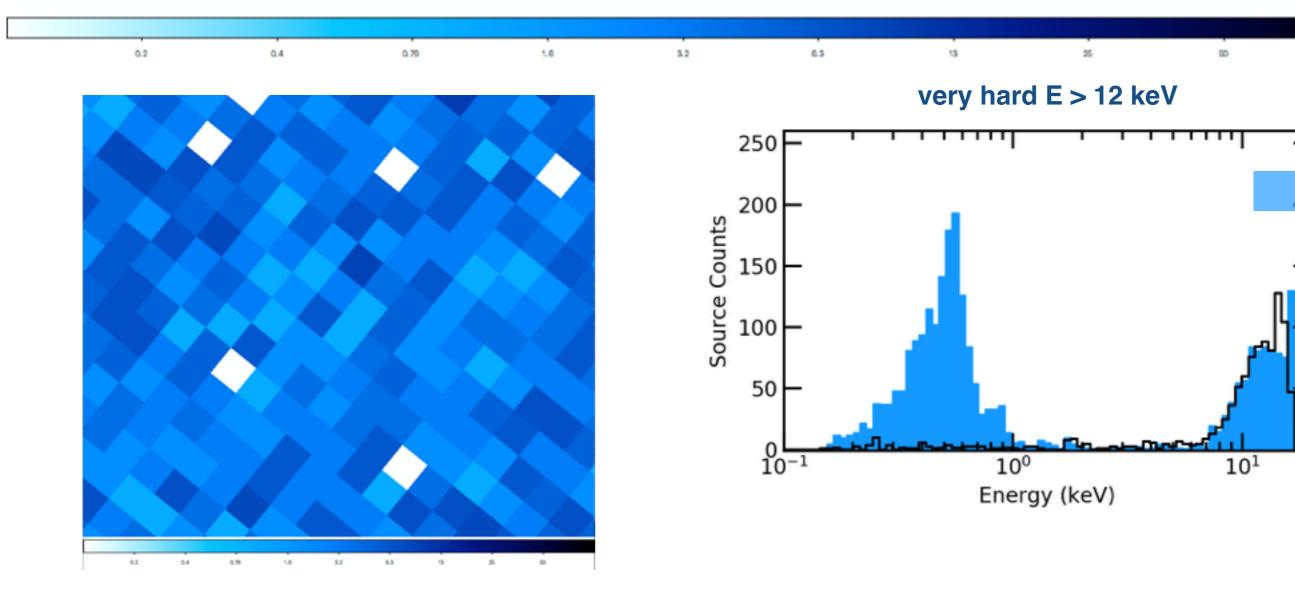
[bin x=::8,y=::8]

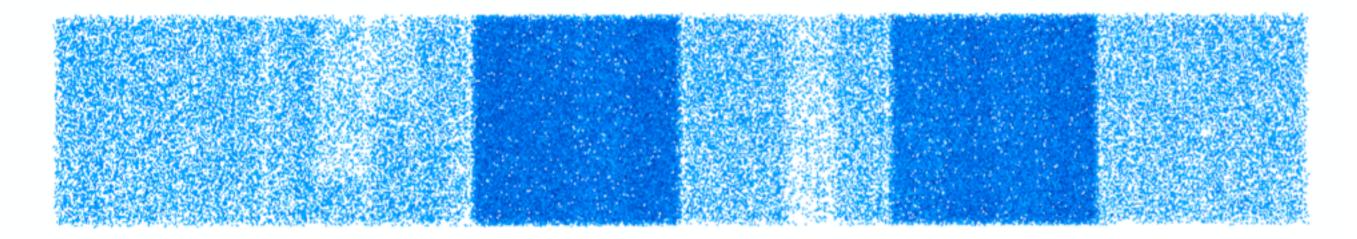
50

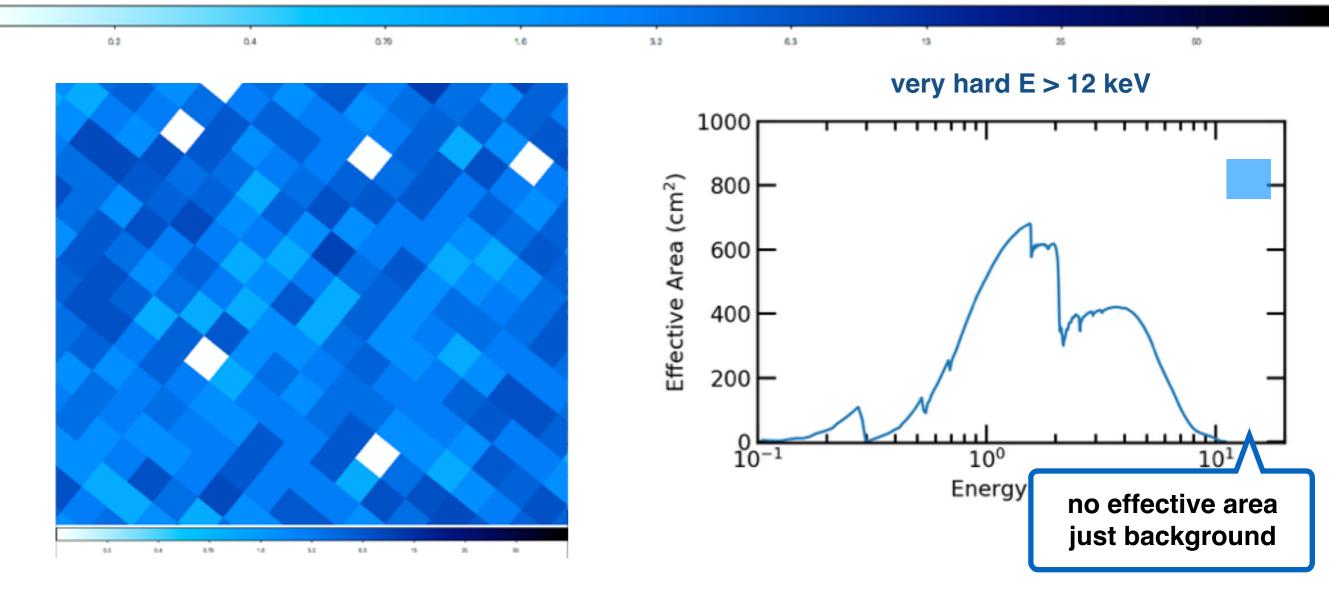


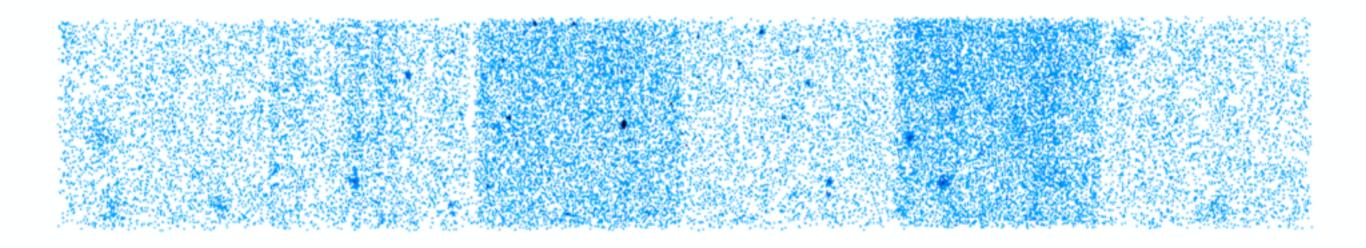


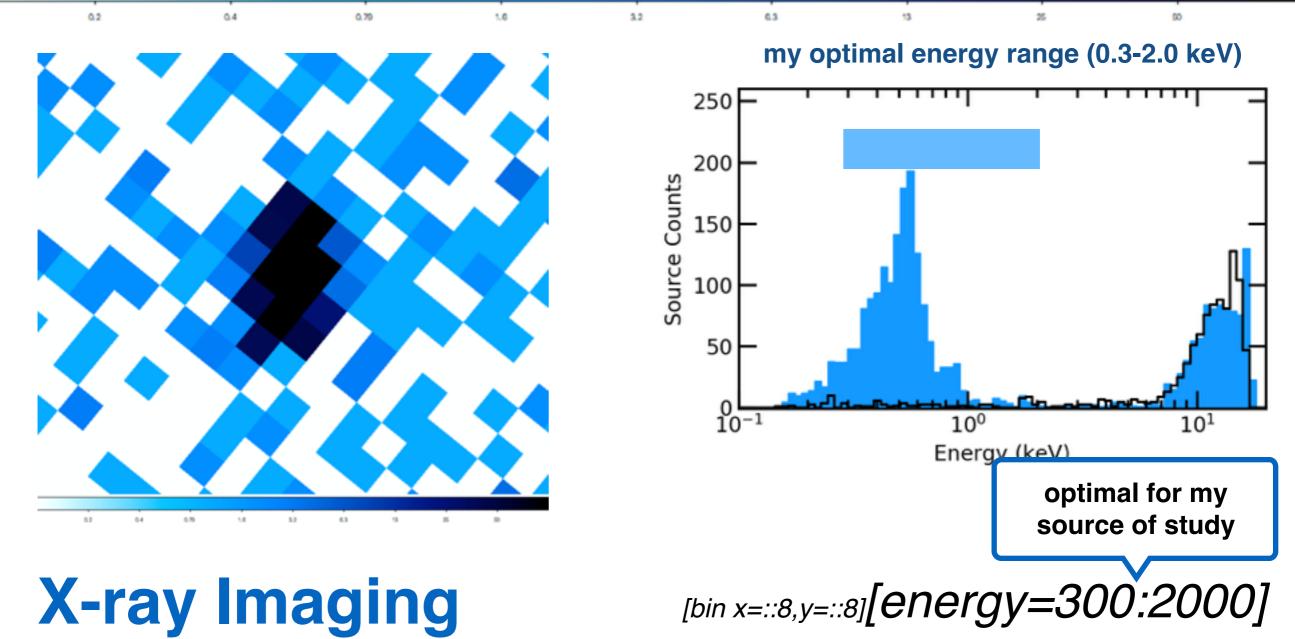


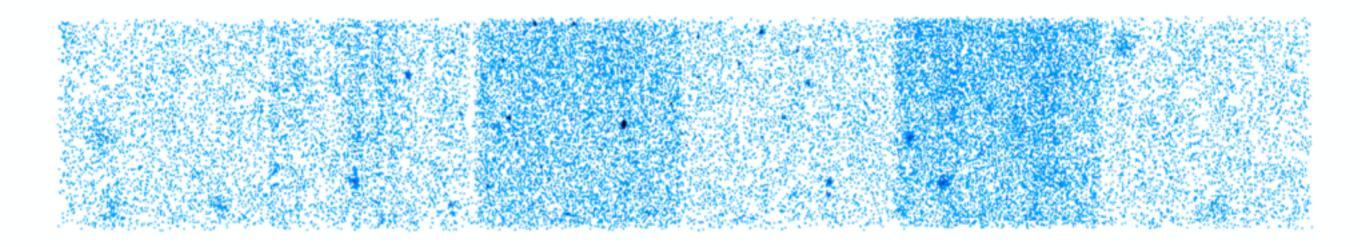


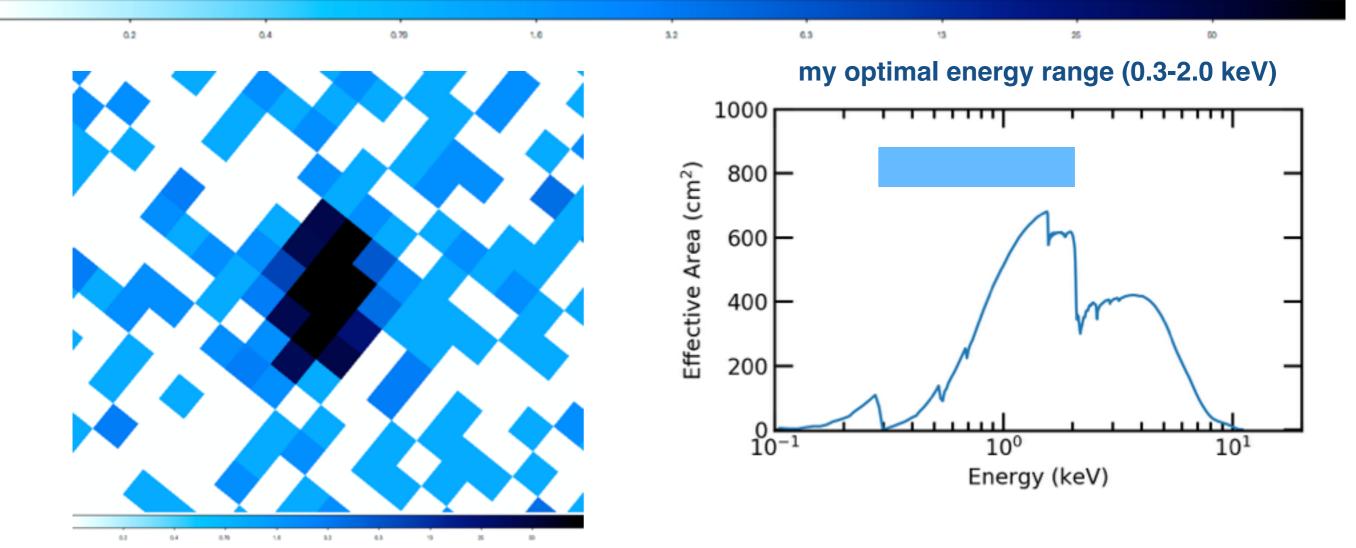


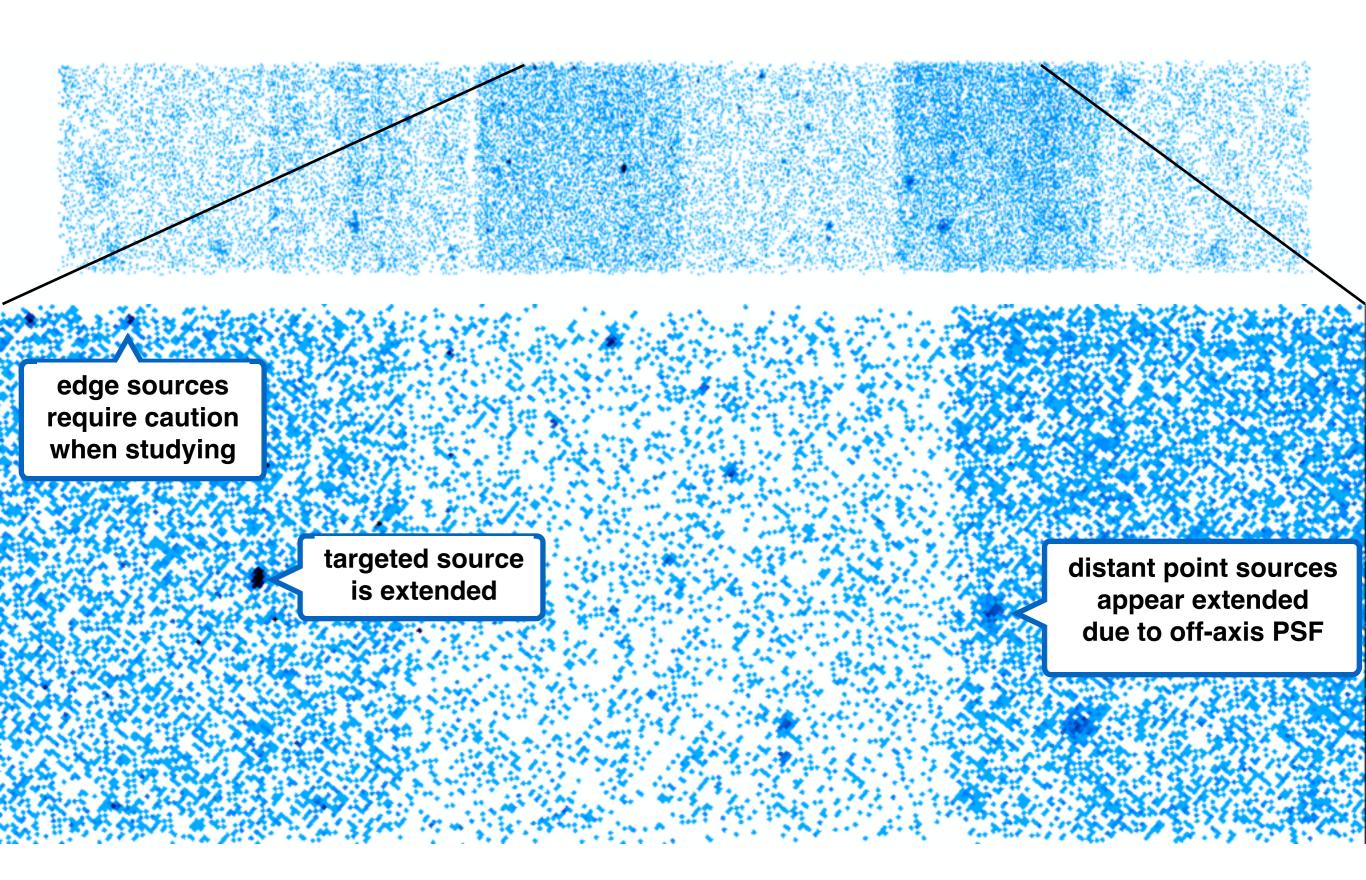


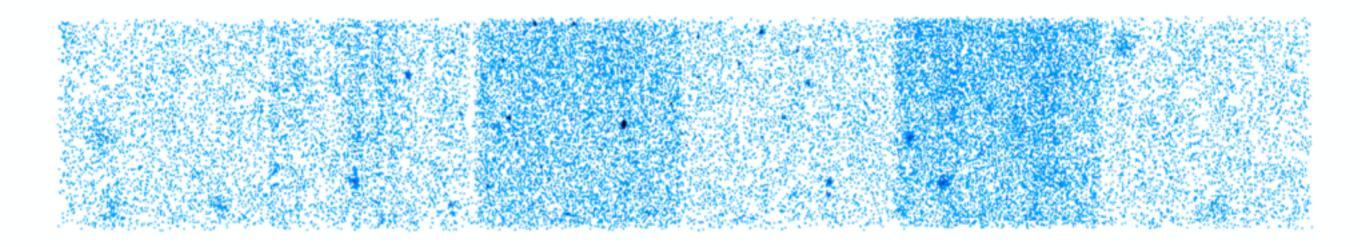


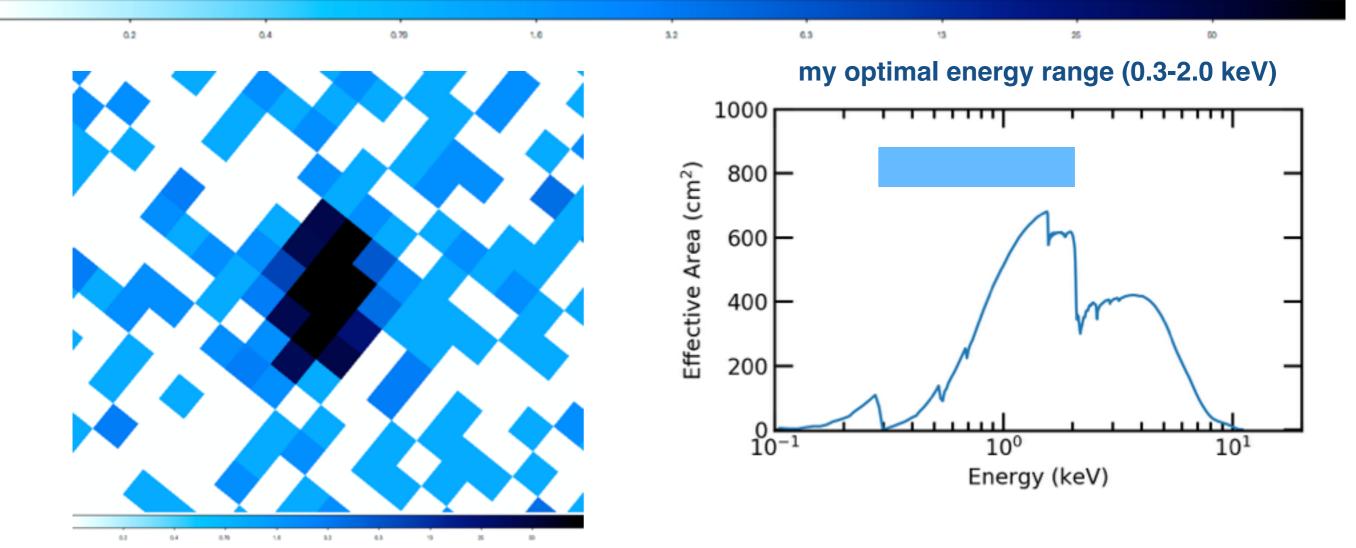


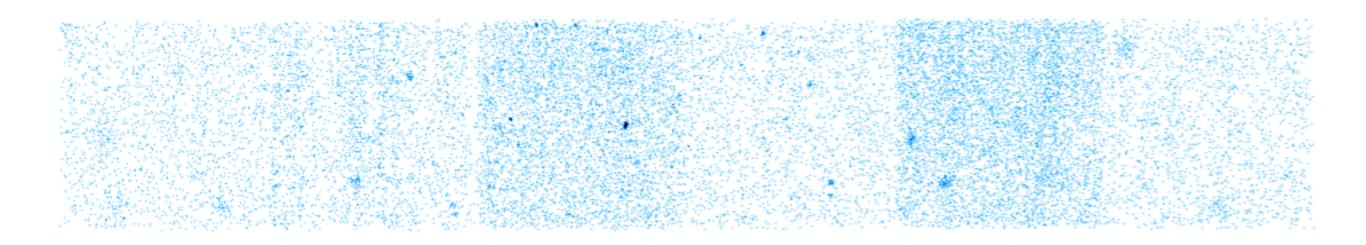


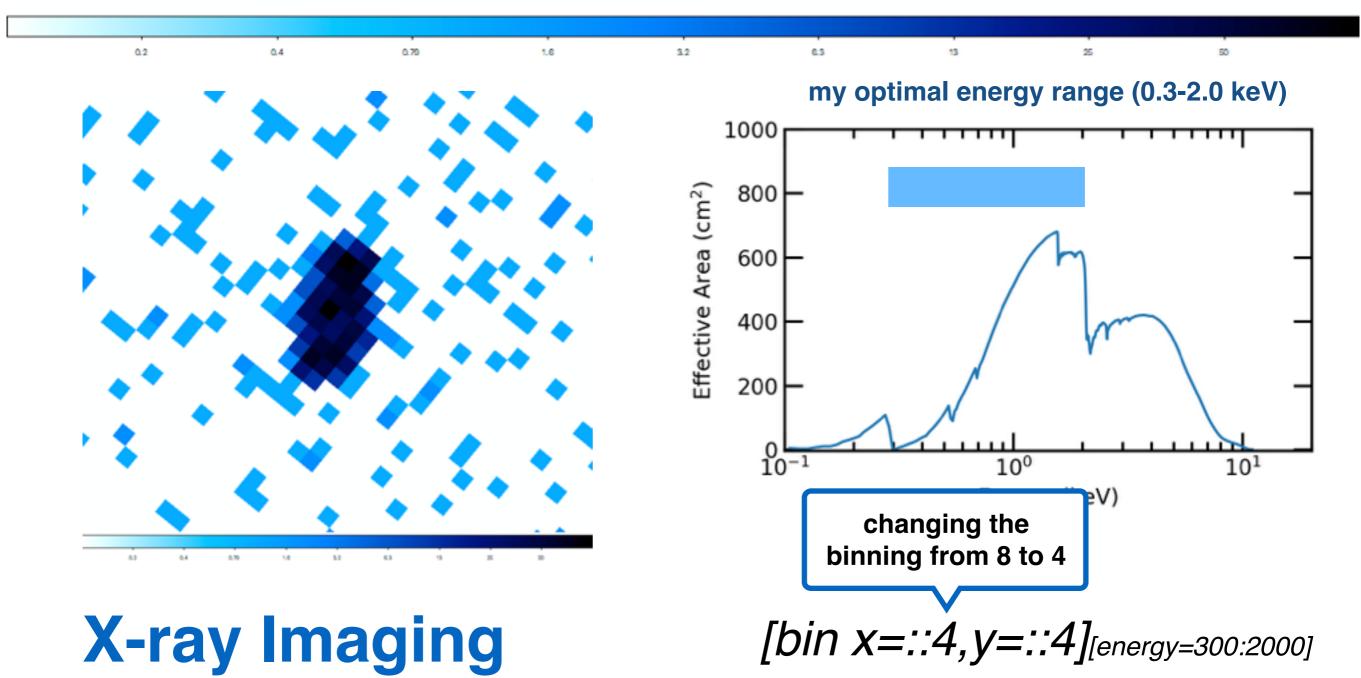


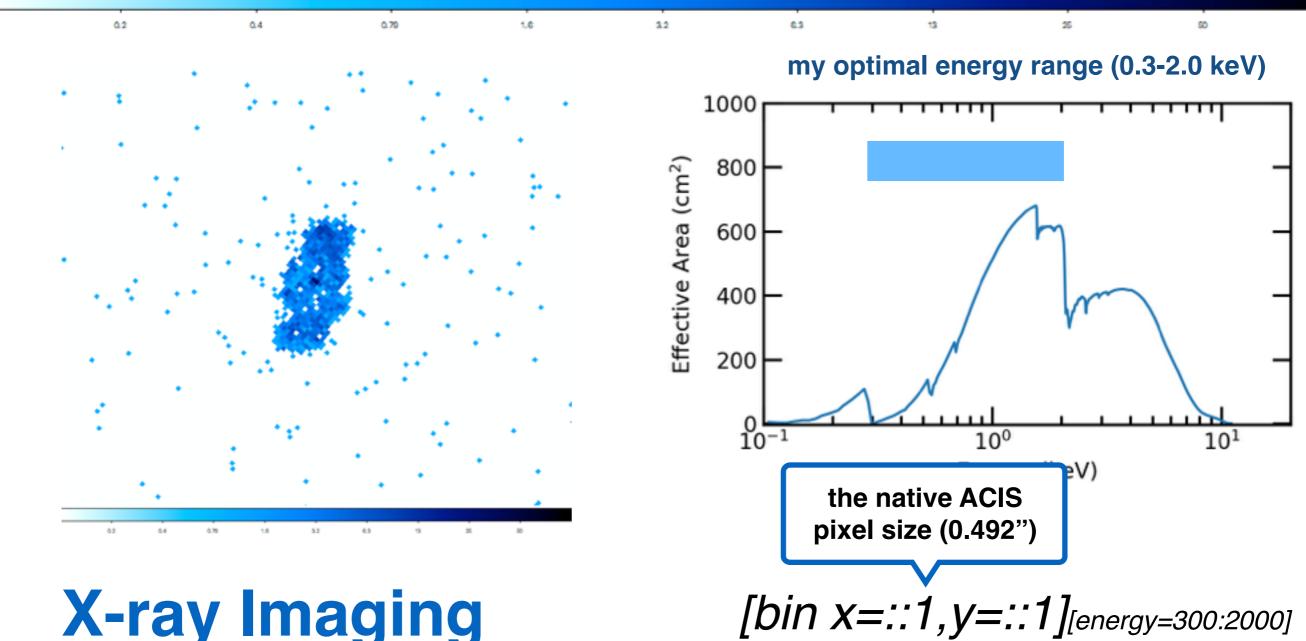


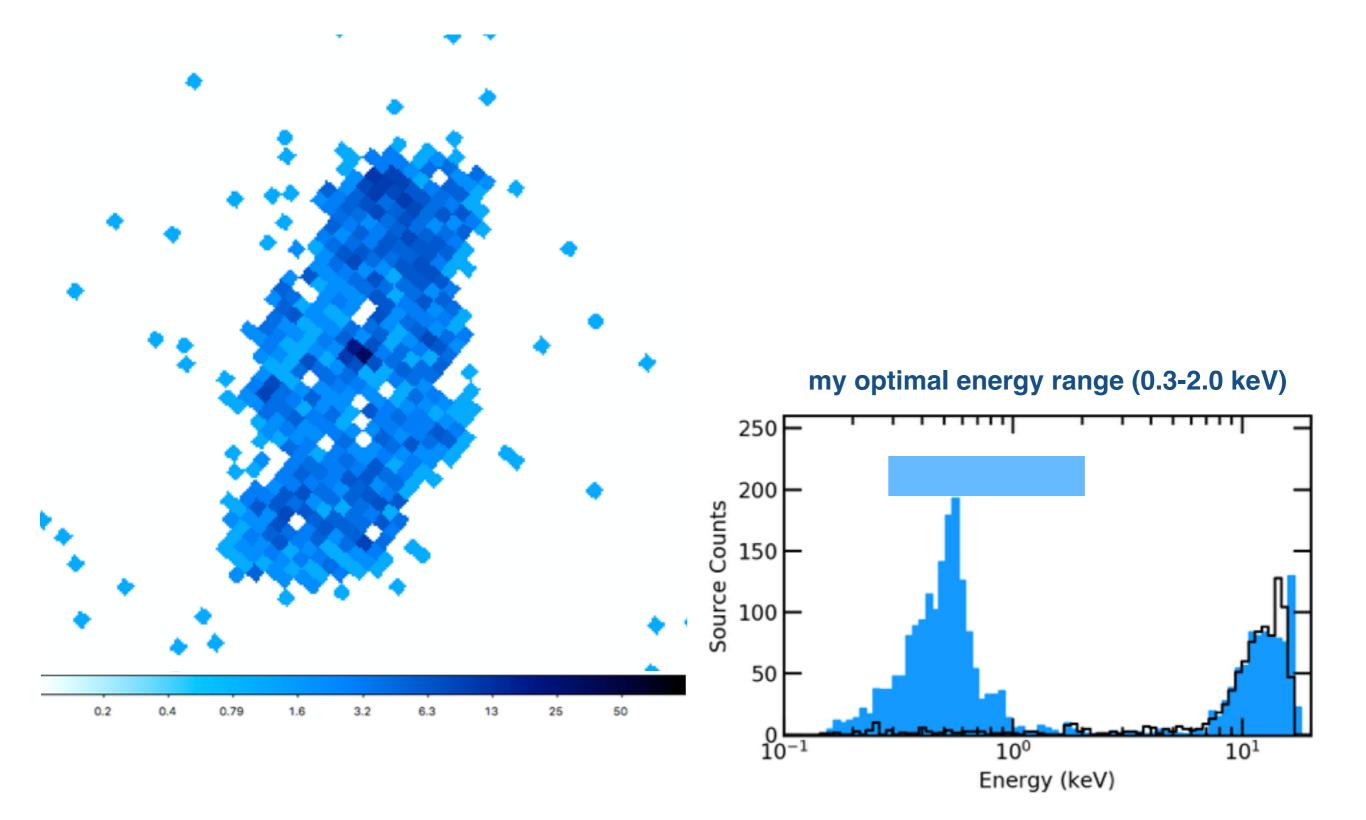




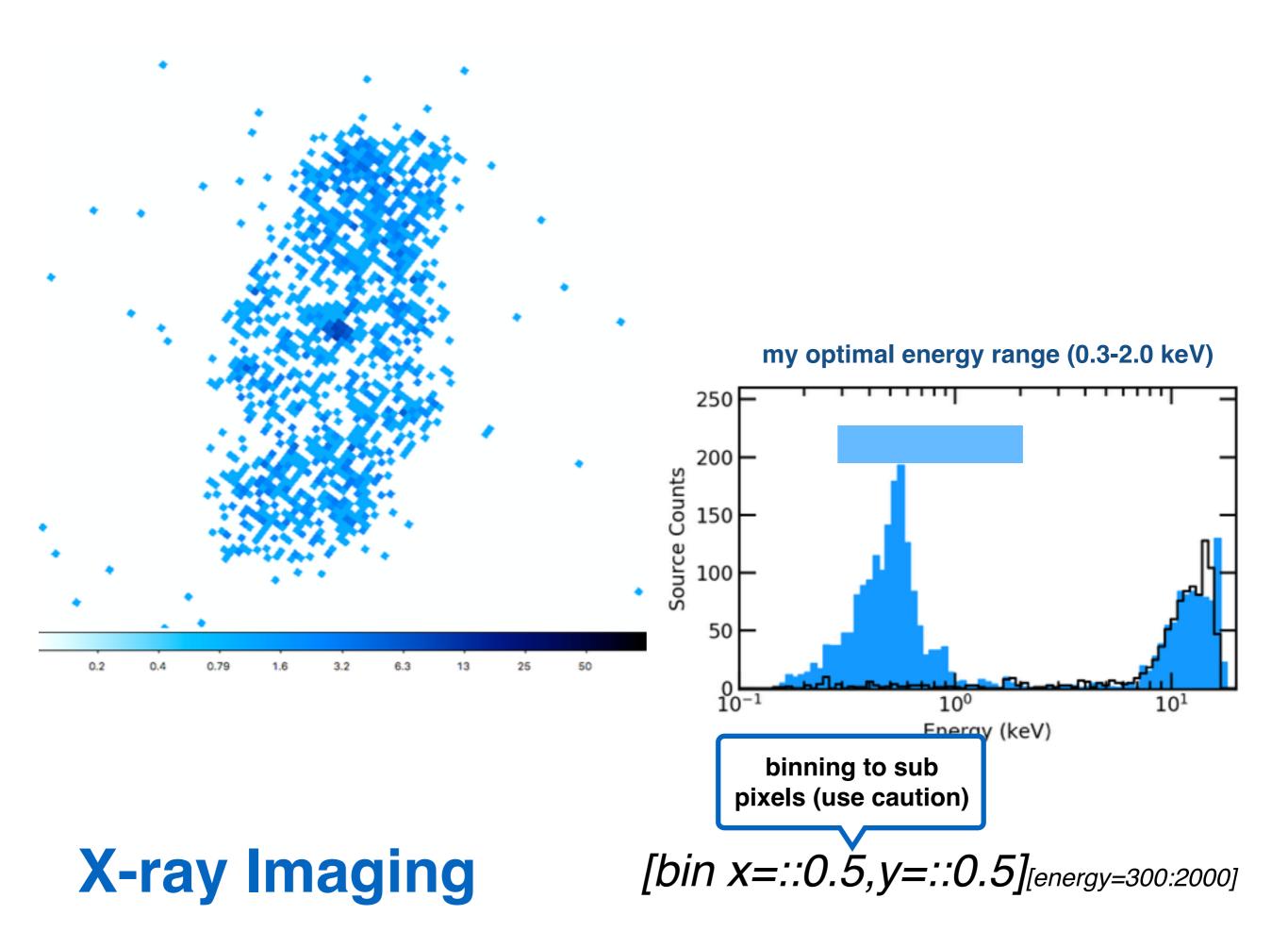


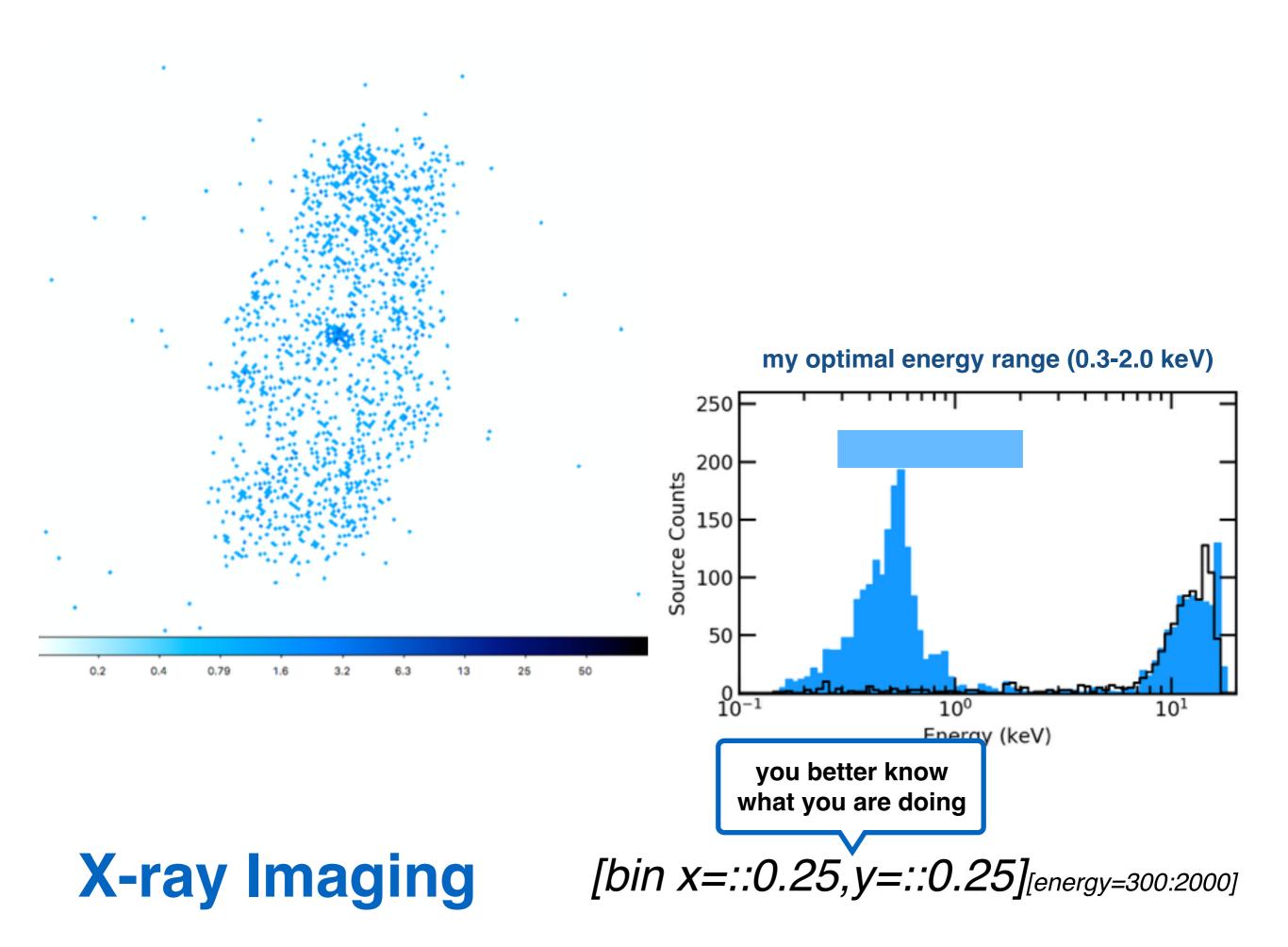


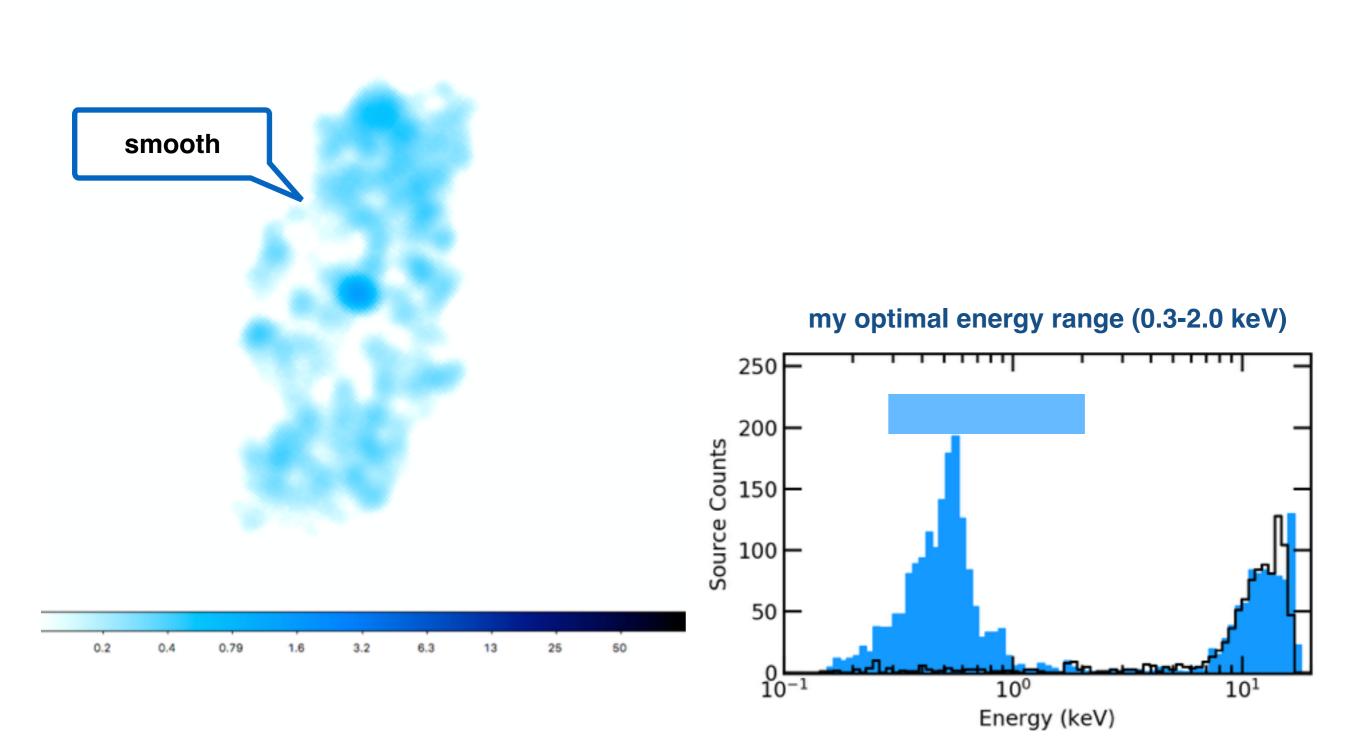




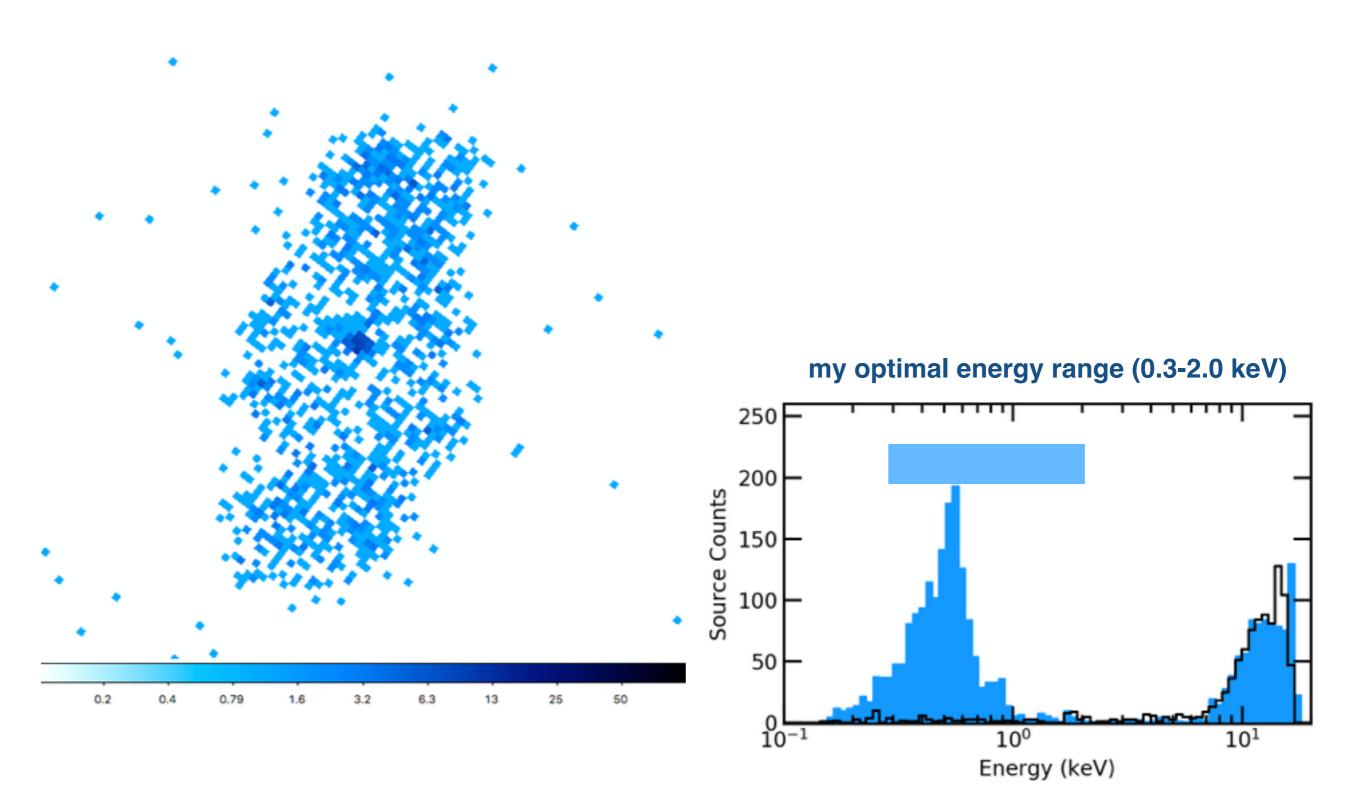
[bin x=::1,y=::1][energy=300:2000]



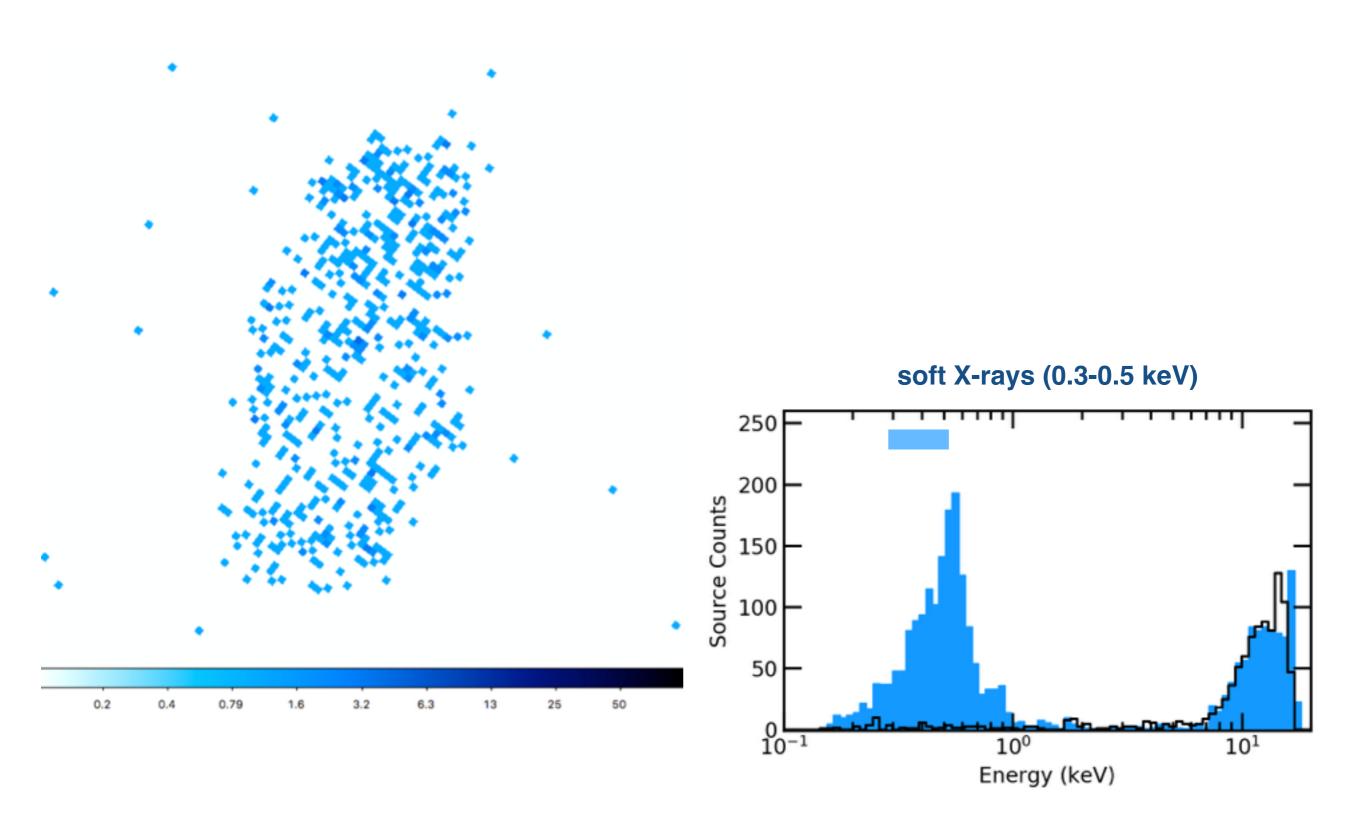




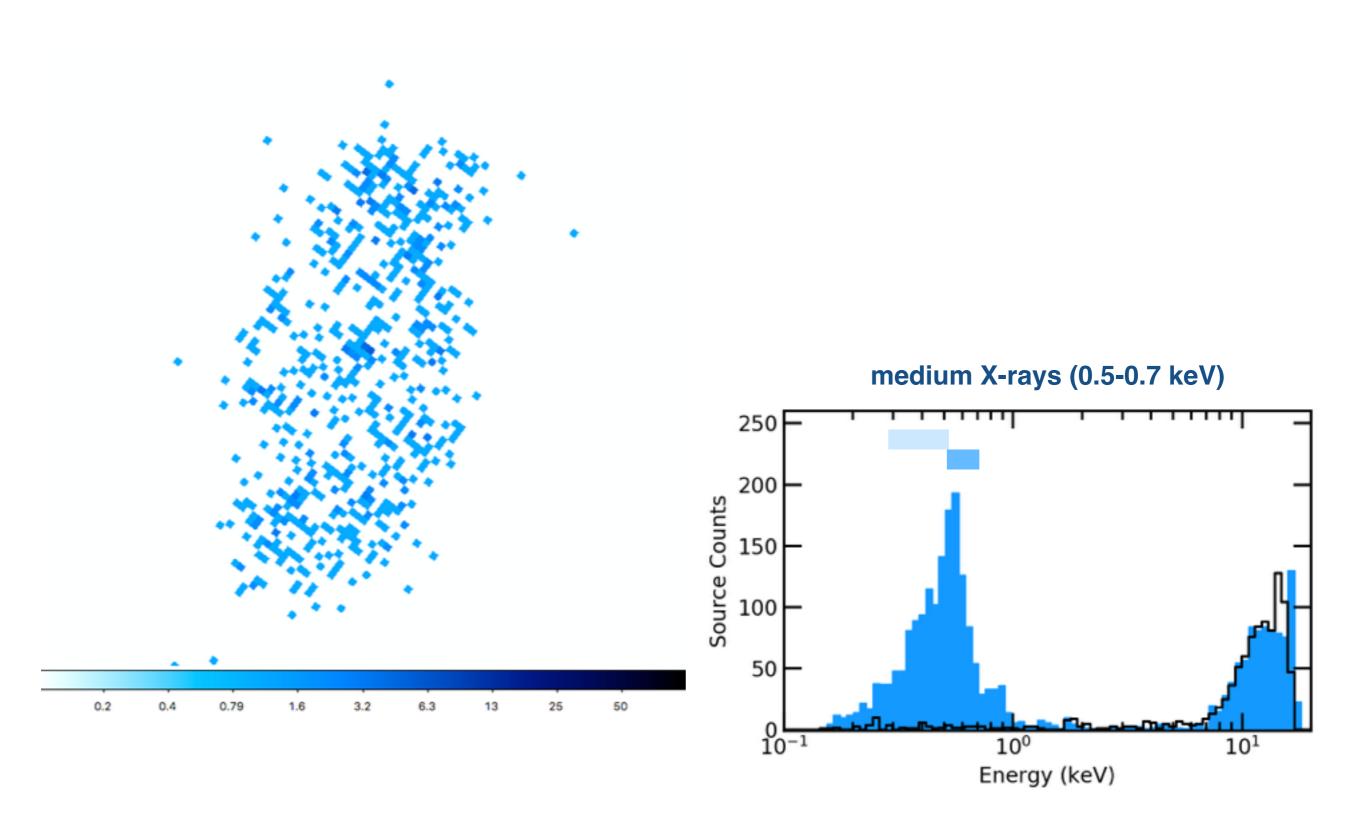
**X-ray Imaging** [bin x=:0.25, y=:0.25][energy=300:2000]



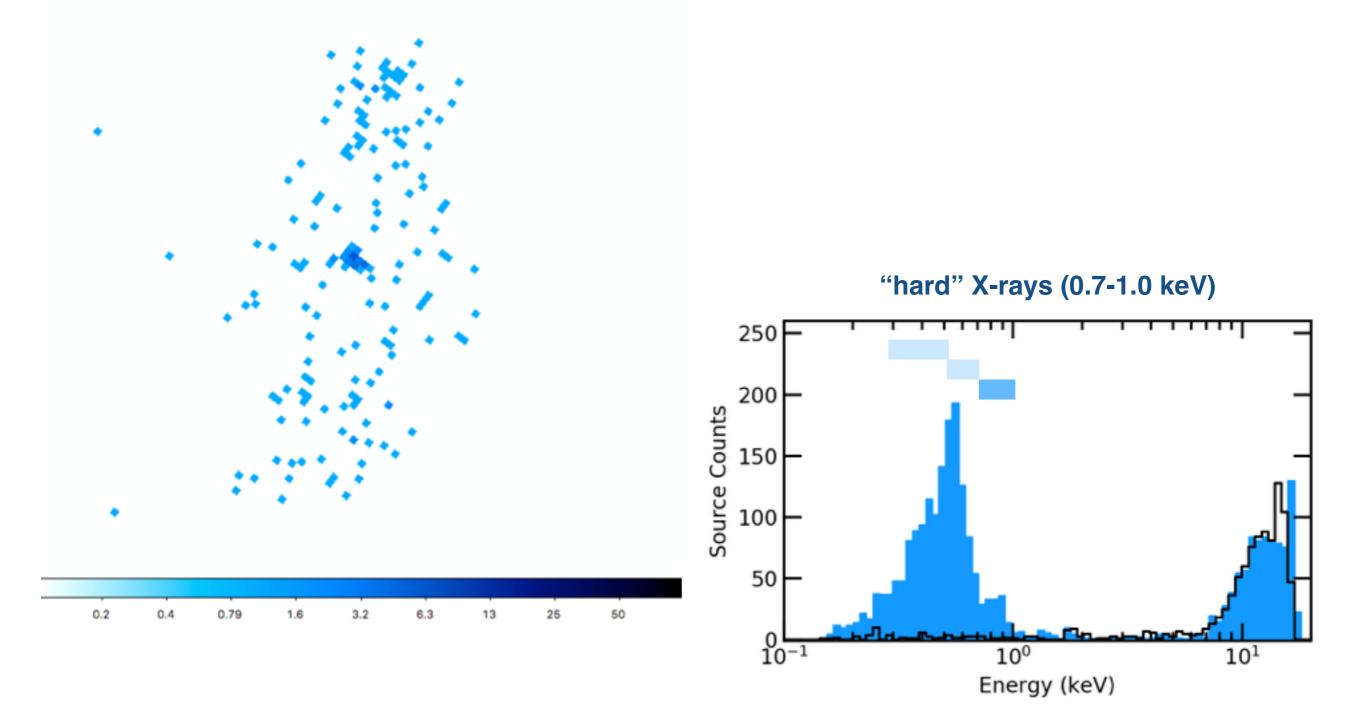
[bin x=::0.5,y=::0.5][energy=300:2000]



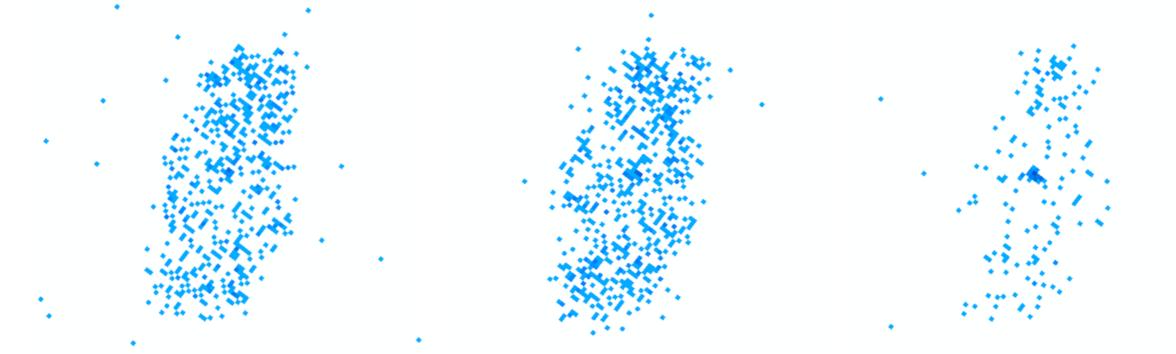
[bin x=::0.5,y=::0.5][energy=300:500]



[bin x=::0.5,y=::0.5]**[energy=500:700]** 

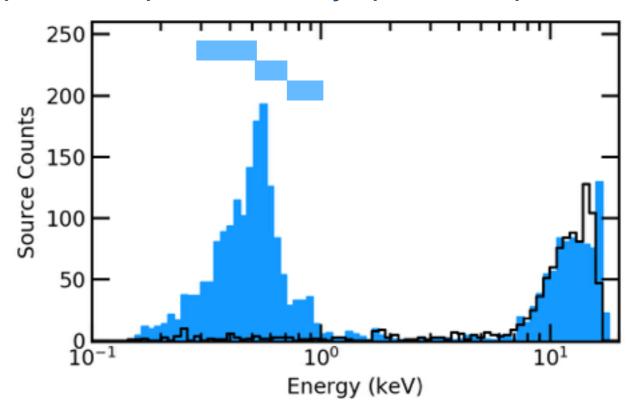


[bin x=::0.5,y=::0.5][energy=700:1000]



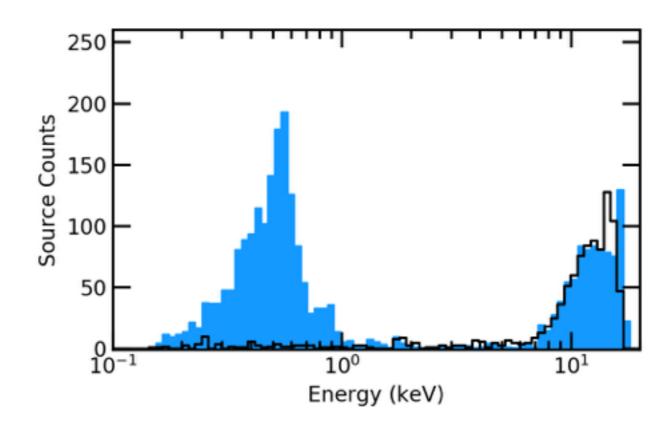
soft X-rays (0.3-0.5 keV)

medium X-rays (0.5-0.7 keV) "hard" X-rays (0.7-1.0 keV)



dmcopy energy filtering

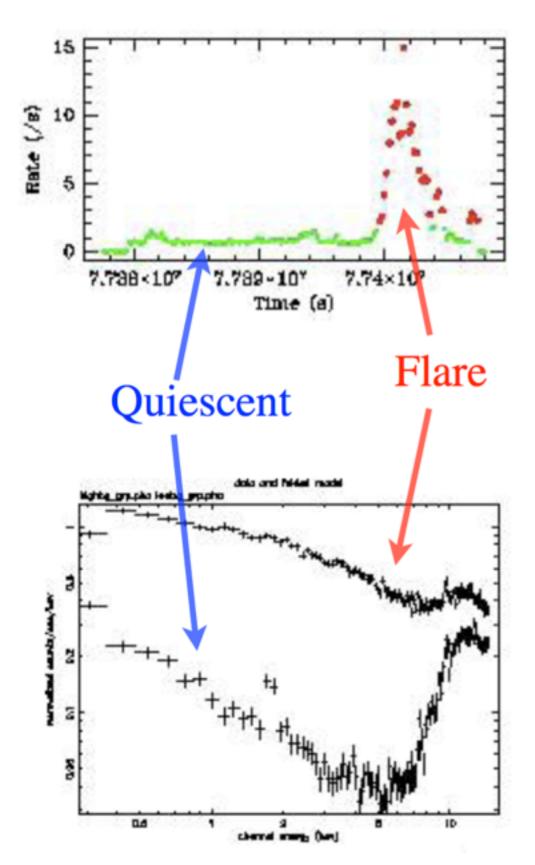
Energy Filtering (ciao → *dmcopy*)

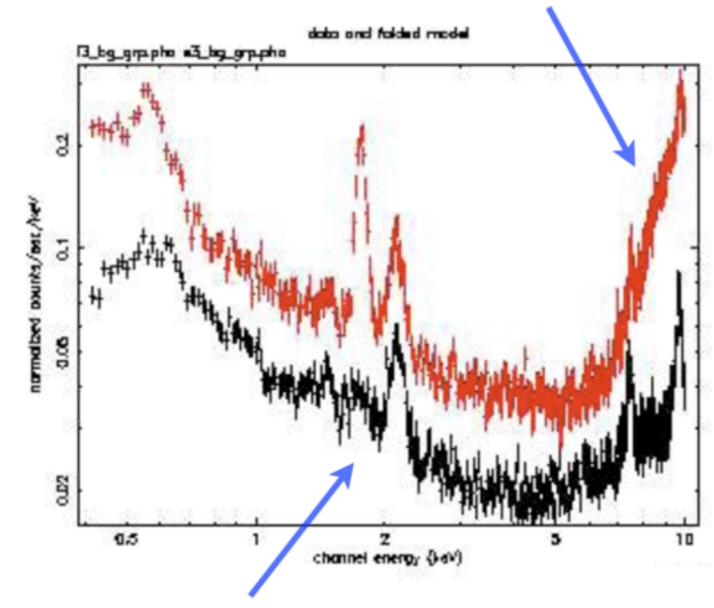


- Energy Filtering (ciao → *dmcopy*)
- Background Flares
  - 1. ciao  $\rightarrow$  *dmextract* (make light curve)
  - 2. chips  $\rightarrow lc\_clean(...)$  (id high bg periods)
  - 3. ciao  $\rightarrow$  *deflare* (remove high bg periods)

http://cxc.harvard.edu/ciao/threads/flare/

**BI** Quiescent





#### FI Quiescent

[time=START\_TIME:END\_TIME]

- Energy Filtering (ciao → *dmcopy*)
- Background Flares
  - 1. ciao  $\rightarrow$  *dmextract* (make light curve)
  - 2. chips  $\rightarrow lc\_clean(...)$  (id high bg periods)
  - 3. ciao  $\rightarrow$  *deflare* (remove high bg periods)

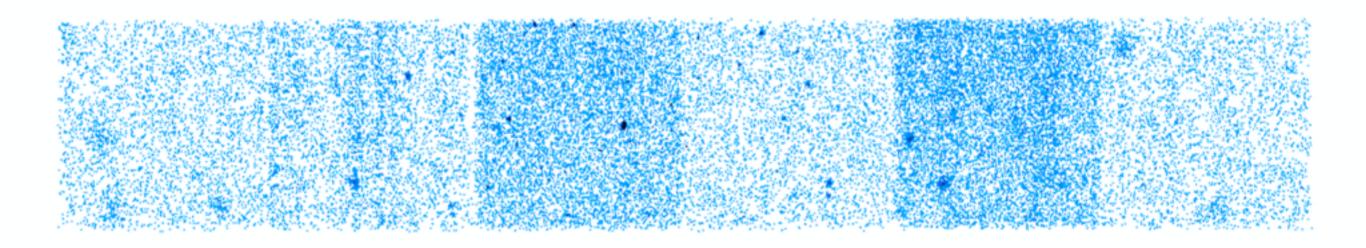
http://cxc.harvard.edu/ciao/threads/flare/

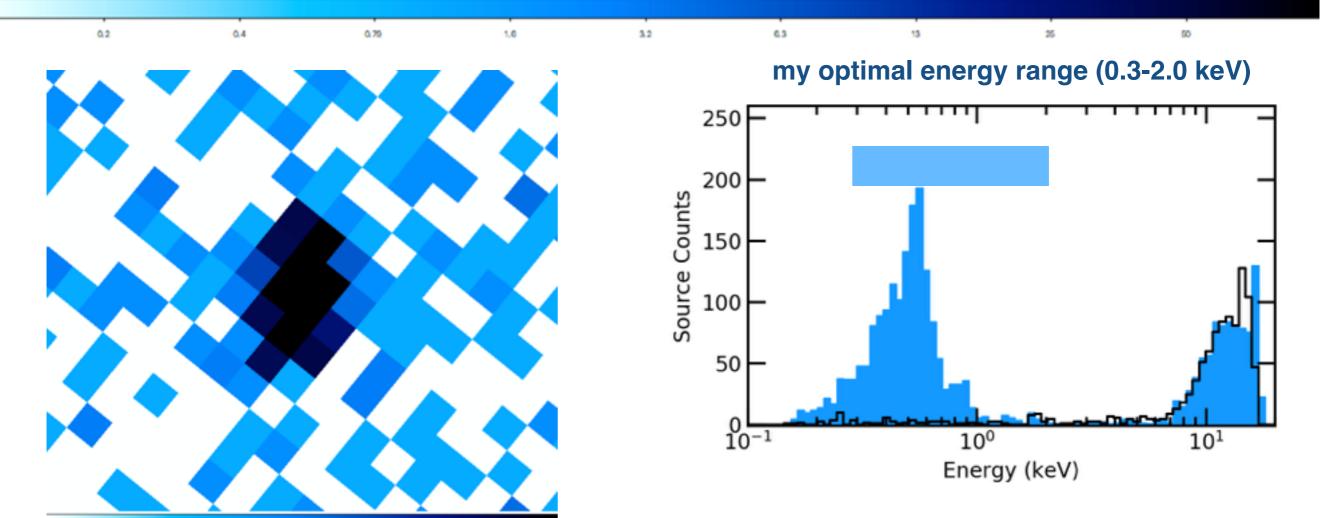
- Energy Filtering (ciao → *dmcopy*)
- Background Flares
  - 1. ciao  $\rightarrow$  *dmextract* (make light curve)
  - 2. chips  $\rightarrow lc\_clean(...)$  (id high bg periods)
  - 3. ciao  $\rightarrow$  *deflare* (remove high bg periods)

http://cxc.harvard.edu/ciao/threads/flare/

- Blank-sky Background
  - 1. remove high bg periods (ciao/chips/ciao)
  - 2. ciao  $\rightarrow$  blanksky (blank bg tailored to obs)

http://cxc.harvard.edu/ciao/threads/acisbackground/





3.2

6.5

15

35

1.6

6.79

6.4

#### What's the Flux?

yesteryear

#### **Creating a Fluxed Image**

- 1. dmcopy (create counts image)
- 2. mkinstmap (create instrument map) [repeat]
- 3. mkexpmap (create exposure map) [repeat]
- 4. dmregrid (combine all the maps)
- 5. dmimgcalc (divide counts image by exp map)



# **Creating a Fluxed Image**

- dmcopy (create counts image)
  feel good
- mkinstmap (create instrument map) [repeat]
  send helpdesk ticket
- 3. mkexpmap (create exposure map) [repeat] send apologetic helpdesk ticket
- 4. dmregrid (combine all the maps) send frantic helpdesk ticket
- 5. dmimgcalc (divide counts image by exp map)





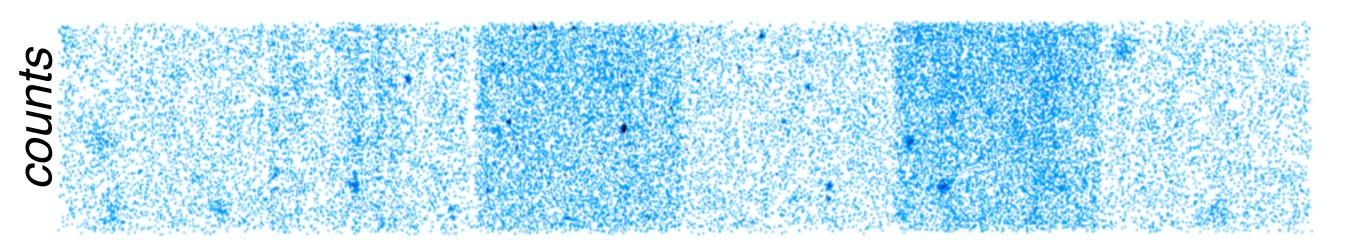
#### What's the Flux?

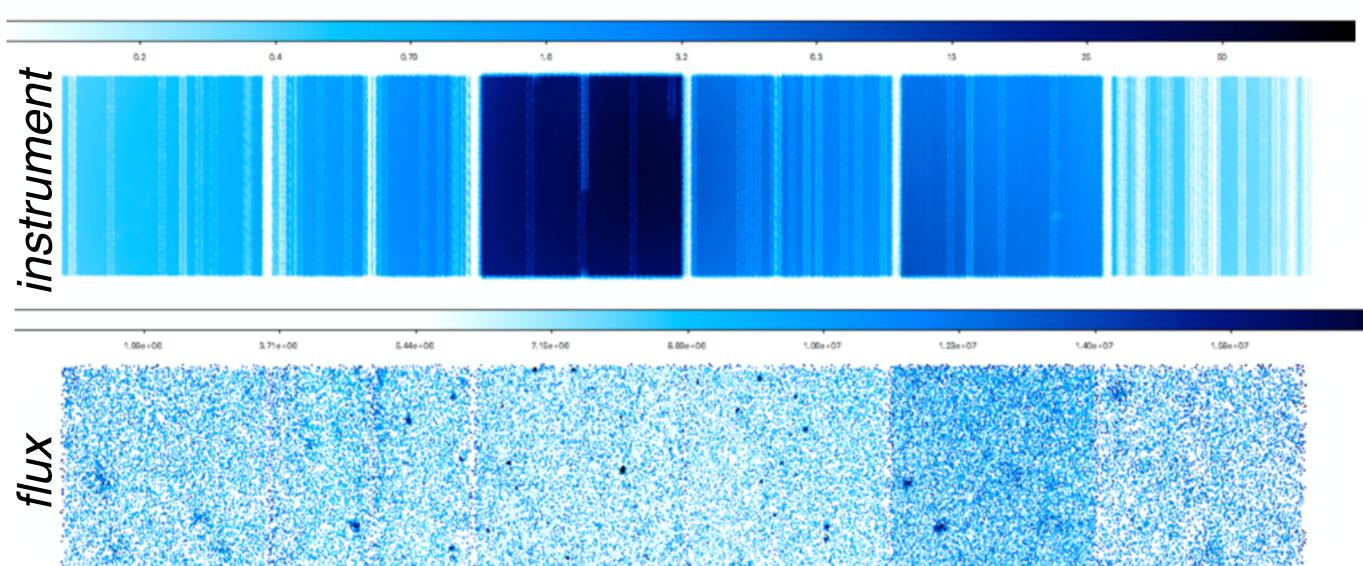
fluximage

#### expmap, fluxed image, etc. output directory l fluximage evt2.fits output/ event file

script will locate the required ancillary files (asol, bpix, msk, etc.)

#### **Flux Images**

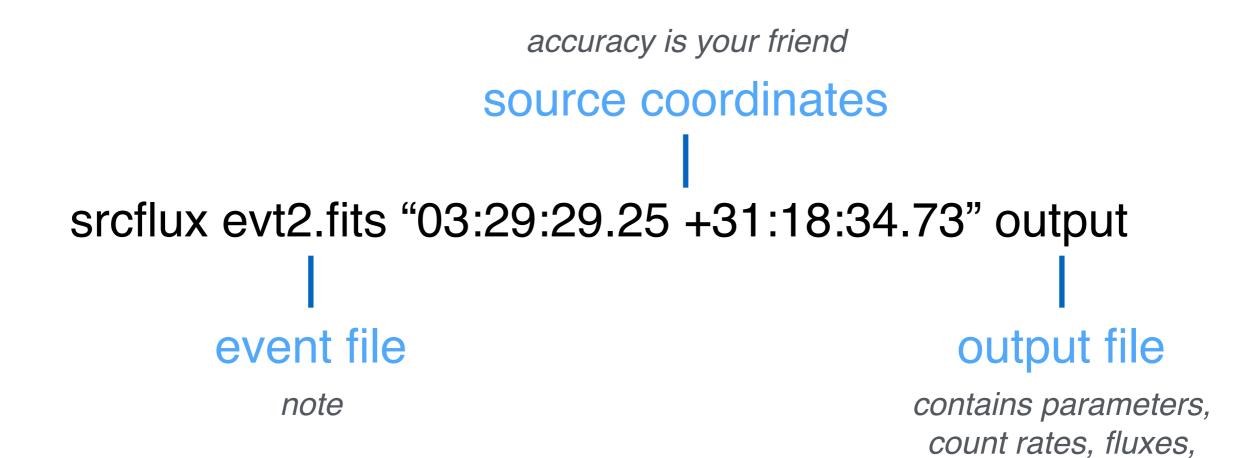






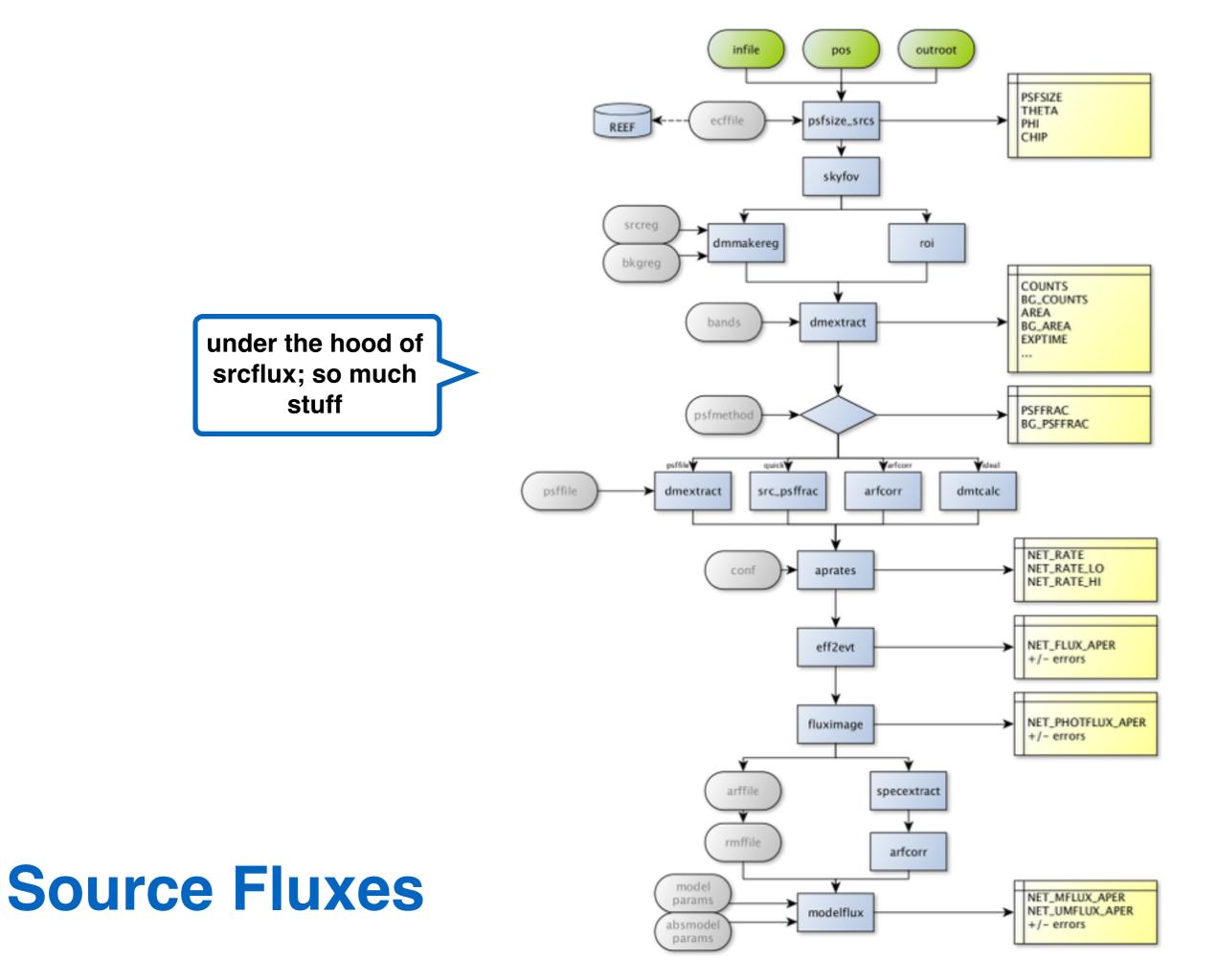
#### What's the Flux?

srcflux



and model flues

#### **Source Fluxes**



#### srcflux infile = repro/acisf06436 repro evt2.fits pos = 03:29:29.250 +31:18:34.73 outroot = single/run1 bands = broad srcreq = bkgreg = bkgresp = yes psfmethod = ideal psffile = conf = 0.9rmffile = arffile = model = xsphabs.abs1\*xspowerlaw.pow1 paramvals = abs1.nH=0.0;pow1.PhoIndex=2.0 absmodel = absparams = abund = angrfovfile = asolfile = mskfile = bpixfile = dtffile = ecffile = CALDB parallel = yes nproc = INDEF tmpdir = /tmp clobber = noverbose = 1 $mode = \alpha l$ Extracting counts Setting Ideal PSF : alpha=1 , beta=0 Getting net rate and confidence limits Getting model independent fluxes Getting model fluxes Getting photon fluxes Running tasks in parallel with 4 processors. Running eff2evt for single/run1 broad 0001 src.dat Running aprates for single/run1 broad0001 rates.par Running eff2evt for single/run1 broad 0001 bkg.dat Making response files for single/run1 0001 Running modeflux for region 1 Adding net rates to output Appending flux results onto output Appending photflux results onto output Computing Net fluxes Adding model fluxes to output Scaling model flux confidence limits

#### Summary of source fluxes

#### Position

#### **Source Fluxes**

output

3 29 29.25 +31 18 34.7 Rate Flux Mod.Flux 0.5 - 7.0 keV Value 90% Conf Interval 0.0398 c/s (0.0381,0.0415) 5.17E-13 erg/cm2/s (4.94E-13,5.39E-13) 4.38E-13 erg/cm2/s (4.2E-13,4.57E-13) Summary of source fluxes

Position

3 29 29.25 +31 18 34.7 Rate Flux show me the flux! Mod.Flux 0.5 - 7.0 keV Value 90% Conf Interval 0.0398 c/s (0.0381,0.0415) 5.17E-13 erg/cm2/s (4.94E-13,5.39E-13) 4.38E-13 erg/cm2/s (4.2E-13,4.57E-13)

#### **Source Fluxes**

