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# **CIAO 3.2 Release Notes**

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# How CIAO 3.2 and CALDB 3.0.0 Affect Your Analysis

Some tools and calibration released in CIAO 3.2 and CALDB 3.0.0 affect analyses in progress, i.e. certain processing steps should be run again for greater accuracy. This section of the release notes describes those changes, as well as the affected tasks.

### **ACIS QE Degradation: Spatial Component**

• The ACIS QE degradation model has been enhanced to account for spatial variations in the contamination on the ACIS optical blocking filters. The contamination is now expressed as a function of time, energy, and ACIS chip coordinate.

For imaging analysis of extended sources or point sources far off-axis, there is a significant change in instrument and exposure maps when the new calibration is applied.

The spatial contamination correction is automatically incorporated into exposure maps and ARF files (made by mkinstmap/merge\_all, mkarf/psextract, mkgarf).

See the ACIS QE Degradation why topic for additional information.

# New method for creating ACIS imaging RMFs

• The tool mkacisrmf is a new addition in CIAO 3.2 and represents an entirely new method for creating ACIS response matrices. The tool is calibrated for ACIS chips 0–3, 5, 6, and 7 (the ACIS–I array, plus ACIS–S1, S2, and S3). The response files must be used with -120 C data that has the time-dependent gain adjustment and CTI correction (if available) applied.

The results from mkacisrmf and mkrmf are very similar. Most users will not lose any accuracy by finishing the analysis with RMFs created by mkrmf. The exceptions are the following cases:

• The new RMFs are 10–15% more accurate at energies lower than 1 keV. If your analysis is sensitive at this level, create a new RMF with mkacisrmf.

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Calibration for the ACIS-S1 (ccd\_id=5) back-illuminated chip has been greatly improved. Anyone doing imaging analysis on this chip should reprocess with mkacisrmf.

For cross-calibration consistency, users doing either of these analyses should reprocess their dataset with the new "phase 2" gain file (the default gain in CALDB 3.0.0).

The Creating ACIS RMFs in CIAO 3.2 why topic has further details on the new tool.

### **ACIS Hot Pixel Tools**

• New tools were introduced in CIAO 3.2 to identify and flag hot pixel and afterglow events in ACIS observations: acis\_build\_badpix, acis\_classify\_hotpix, and acis\_find\_hotpix. Also included is acis\_run\_hotpix, a script that runs the ACIS hot pixel tools in the correct order.

These tools supersede acis\_detect\_afterglow and replaced that tool in standard data processing in DS 7.4.0. acis\_run\_hotpix is better because it finds some afterglows that acis\_detect\_afterglow does not, it also finds pixels that are hot or have bad bias values, and it misidentifies far fewer legitimate X-ray events as bad.

The Identify ACIS Hot Pixels and Cosmic Ray Afterglows thread has more information on these tools and how to use them. Note that it is necessary to reprocess the data with acis\_process\_events in order to apply the new bad pixel file.

### **Grating Analysis Improvements**

• A number of CIAO tools have had improvements that will positively impact grating analyses: asphist, mkgarf, tg\_create\_mask, tg\_resolve\_events, and tgextract. The changes are detailed in the Tools section of these release notes.

### CALDB changes since the CIAO 3.1/CALDB 2.27 release

- CALDB v2.27 was released with CIAO 3.1 on 7 July 2004. For users who have not upgraded their CALDB since then, here is a summary of significant calibration updates. All of these changes are included in the CALDB 3.0.0 release; see the CALDB release notes for details on these and other changes.
- ACIS QE version 5 (CALDB v2.28): contains new QE information for for chips ACIS-5 and ACIS-7 (the back-illuminated chips). Users will be able to obtain more consistent flux measurements between the front- and back-illuminated chips with this update.

There is still a residual few percent difference between the FI and BI chips which will be accounted for in a future calibration update.

• *ACIS Time-dependent gain correction* (CALDB v2.29): new time-dependent gain calibration files for February, May, and August 2004 (epochs 18 and 19) were released. Plots in the ACIS Time-dependent Gain why topic illustrate how the new calibration files affect the magnitude of the correction.

# Tools

acis\_build\_badpix

#### • New Tool!

One of the new ACIS hot pixel tools; creates an observation-specific bad-pixel file.

For more information on using this tool, see the Identify ACIS Hot Pixels and Cosmic Ray Afterglows thread.

#### acis\_classify\_hotpix

#### • New Tool!

One of the new ACIS hot pixel tools; determines whether a suspicious pixel is hot or affected by an afterglow.

For more information on using this tool, see the Identify ACIS Hot Pixels and Cosmic Ray Afterglows thread.

#### acis\_find\_hotpix

#### • New Tool!

One of the new ACIS hot pixel tools; identifies pixels that have an unusually large (or small) number of events.

For more information on using this tool, see the Identify ACIS Hot Pixels and Cosmic Ray Afterglows thread.

#### acis\_process\_events

- unneeded cc\_chipy parameter removed
- Bug fix: if the string "qp" is in the output filename the tool segmentation faults (fix in dslib)
- updates to allow afterglow bits to be set from hotpixel-ammended bad pixel file.

#### acis\_run\_hotpix

#### • New Tool!

A script which runs the three hot pixel tools (acis\_build\_badpix, acis\_find\_hotpix, and acis\_classify\_hotpix) in the correct order.

For more information on using this tool, see the Identify ACIS Hot Pixels and Cosmic Ray Afterglows thread.

#### acisreadcorr

• Bug fix: default value of bgroup parameter is changed to 10

#### aconvolve

- enabled support for 1D images
- FFT speed improvement
- Bug fix: errors instead of crashing when kernelspec parameter is left blank
- Bug fix: added error checking to kernelspec for extra/missing parentheses or commas

#### arestore

• changed order of parameters make consistent with other tools

#### asphist

- Bug fix: the aspect histogram can now be made for nominal aimpoints which do not fall on a detector element at all during the observation. This was done intentionally for grating observation of Sco X-1 to prevent detector damage by an intense zero order image, while allowing two of the 4 HETG spectra to fall partially on the ACIS-S array.
- Bug fix: nominal aim falling into a chip gap caused asphist to fail
- The asphist computation time has been reduced by about a factor of two by removing an unnecessary second pass through the data.

The speed improvement updates required replacing the max\_bin parameter with two new parameters: delta\_xy and delta\_roll.

• Bug fix: corrected typo in error message ("DTF" in place of "LTF")

#### axbary

• increased keyword precision in the output file

#### celldetect

- output source list has improved metadata (TLMIN/MAX)
- Bug fix: added "double" checking: will still output duplicate srcs, but the double flag/column will be set for one of them to make removing the duplicates easier

### csmooth

- small speed improvement in FFT
- Bug fix: bad redirect in the sclmax parameter ()sclmax changed to )sclmin)

#### destreak

- The mask parameter which used to be for a region to ignore during the processing has been generalized so more than just spatial filters can be applied. The upshot being that you can now destreak a L1 event file with mask="[status=0,grade=0,2,3,4,6]" and it will use only those values when processing the data.
- implemented a maximal-row-loss algorithm; new parameter: max\_rowloss\_fraction controls the max number of rows per frame to be streaking.

### dmcoords

• dmcoords does not automatically take into account drifts in the Chandra geometry when calculating sky coordinates. These drifts were negligible near the beginning of the mission, but have grown to the order of 10 arcseconds and so cannot be neglected. The Using dmcoords to Convert between CHIP and SKY Coordinates thread shows how to account for this change.

A new parameter (asolfile) was added to use the aspect solution files to account for the geometry drifts in the calculations.

• Bug fix: the "set nominal" command now respects the celfmt parameter to use either HMS or DEG format

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• Bug fix: made internal help consistent with interactive mode setting for SIM/DISPLACE

#### dmcopy

- Bug fix: improvements for handling image cubes (fix in the Data Model library)
- Bug fix: set pixels outside of region filter to NULL for >2D images (fix in the Data Model library)

#### dmdiff

- The tool now checks:
  - ♦ images (1D and 2D)
  - ♦ WCS
  - ♦ DSS
  - ♦ Units
  - ♦ Comments
  - ♦ Datatypes
  - ♦ Missing keys
  - ♦ Ranges

Many new parameters have been added to support the extended comparisons. See the dmdiff ahelp file for a complete list.

- added better checking for Inf and NaN in image data
- Bug fix: ignoring subcomponent tolerance ranges

#### dmextract

- Bug fix: in ltc1 or ltc2 mode, the value of TIMEDEL is set to the time resolution of the lightcurve. Similarly TIMEPIXR is set to 0.5 since the value in the TIME column is the mid-bin.
- Bug fix: read the TLMIN, TLMAX, and TDBIN keywords when a binning specification is not given
- added new way of handling binning on EXPNO
- fixed problem with uncertainities when user supplies variance image
- name of the vector column from the input image is kept (was hard-coded to be POS)

### dmfilth

- added NULL, NaN, and DSS image checking
- change to randseed parameter limits to allow value of 0

### dmgroup

- Bug fix: allow BIN to work on ascending or descending data (e.g. to bin wavelengths that decrease)
- Bug fix: BIN\_WIDTH always working on row/channel instead of on xcolumn data
- Bug fix: problem with bin=::<value>
- Bug fix: problem with BIN\_FILE when non-integer xcolumn is used
- Bug fix: last element now marked as quality=2 when SNR grouping (fix to the group library)
- Bug fix: typo in the error message for grouptype=NUM\_BINS

### dmgroupreg

• Bug fix: tool name is included in error messages (uses errlib)

### dmgti

- Bug fix in smoothing: last n/2 rows were off-by-one where n is the smoothing factor/size
- allow .[0-9]+ to be parsed as a number without requiring lead 0
- Bug fix: casting when converting numeric values from one format to another with (int)
- Bug fix: <array>\*\*<scalar> giving array all equal to first element/pixel
- Bug fix: not using correct array index in expressions (e.g. sum=(phas[2]+phas[3])). Note that setting individual elements in an array still does not work, e.g. (phas[2]=32)

### dmhedit

• Bug fix: empty lines in the file list cause problems

#### dmhistory

• New functionality!

A new action parameter has been added to determine what dmhistory should do with the information it obtains. The default is get as before (i.e. send it to the outfile); new options are pset (set the parameter file for the requested tool) and put (the parameter file for the given tool is written to the data file).

• Bug fix: tool keeps trailing whitespace when necessary

## dmimg2jpg

• Bug fix: modifed region labels (showlabel parameter). corrected problems where lines were too long and source numbers would appear in the wrong place

### dmimgcalc

- allow .[0-9]+ to be parsed as a number without requiring lead 0
- Bug fix: casting when converting numeric values from one format to another with (int)
- Bug fix: <array>\*\*<scalar> giving array all equal to first element/pixel
- Bug fix: tool will error is impout expression does not contain an "impout=" variable
- Bug fix: if the input file is missing, dmimgcalc dies with a segmentation fault
- Bug fix: table in second extension opened instead of image
- Bug fix: typo in error message ("Could [not] correctly parse expression")

### dmimghist

- output includes bin\_low and bin\_hi columns to output denotes the lower end and (non-inclusive) upper limit, respectively, of pixel values in a bin
- added NULL, NaN, and DSS image checking

### dmimgthresh

• added NULL, NaN, and DSS image checking

### dmlist

• Bug fix: listing bits with array option not working (fix in the Data Model library)

#### dmmakereg

- Bug fix: there is better error reporting when the tool cannot parse a region
- Bug fix: ASCII output when region contains Points

#### dmregrid

#### • New functionality!

It is possible to specify the bin grid in either image or physical coordinates. A new parameter (coords\_sys) controls whether bin is interpreted as logical (the default, old behavior) or physical.

Images covering different physical pixel ranges and at different resolutions (binning) can now be combined without applying the x/yoffset parameters.

- $\bullet$  allow for non–integer bin and min/max x/y
- Bug fix: clobber would not clobber outfile that had extension supplied

#### dmsort

• Bug fix: corrected behavior when filtering columns and copyall=yes

#### dmstat

- Note that dmstat only works on images if they are 2-dimensional.
- improved handling of min/max location values

#### dmtcalc

- Bug fix in smoothing: last n/2 rows were off-by-one where n is the smoothing factor/size
- allow .[0-9]+ to be parsed as a number without requiring lead 0
- Bug fix: casting when converting numeric values from one format to another with (int)
- Bug fix: <array>\*\*<scalar> giving array all equal to first element/pixel
- Bug fix: not using correct array index in expressions (e.g. sum=(phas[2]+phas[3])). Note that setting individual elements in an array still does not work, e.g. (phas[2]=32)
- Bug fix: the tool SEGVs when setting bits in a file (Linux only)
- Bug fix: will return non-zero exit status if unable to evaluate expression

### get\_src\_region

- restored to CIAO (had been removed at the CIAO 3.0.0 release). get\_src\_region determines a background threshold value from an image or event file and outputs regions with counts higher than that threshold.
- now able to process non-integer (floating point) images
- new parameter (binning) to control image histogram binning specification
- added NULL, NaN, and DSS image checking

#### mkacisrmf

• New Tool!

The tool mkacisrmf is a new addition in CIAO 3.2 and represents an entirely new method for creating ACIS response matrices. Unlike its predecessor mkrmf, this new tool separates the RMF calculation process into two components: an "ideal" component which describes the CCD spectral

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response prior to the effects of CTI, and a spatially varying component which incorporates the changes in the response produced by CTI.

The algorithm contained in mkacisrmf is based upon the algorithm developed by Alexey Vikhlinin and implemented in his calcrmf2 tool. As compared to the original tool (mkrmf) it includes a number of enhancements, including an improved interpolation scheme to calculate the response at intermediate energies between the available calibration points. All functionality of mkrmf is available including the ability to produce weighted response matrices for arbitrary spatial regions.

For more information on this tool, see the Creating ACIS RMFs in CIAO 3.2 why topic.

#### mkrmf

• Bug fix: tool crashed if input FEF filename contained an error

#### mkwarf

- Bug fix: added support for 0th order gratings
- Bug fix: entire response region is bad or outside of the mask

### psf\_project\_ray

- new parameter, detsubsysmod, (similar to the mkwarf detsubsysmod parameter) used to set ARDLIB BPMASK qualifier; was previously hard-coded as BPMASK=0
- the tool will now work with multi-OBI event files (RA\_PNT vs RA\_NOM keywords when available and different)
- Bug fix: passes correct detector name to ARDLIB for HRC data

#### reproject\_events

- Bug fix: if a GTI does not exist for the first event in the file, the remainder of the file is skipped
- Bug fix: rand\_seed not getting initialized correctly, which could cause problems with the background files

### skyfov

- Bug fix: for problems seen with near-zero roll angles. More robust convex hull algorithm uses 8 points at each corner at min/max ra/dec/roll to determine polygon, and the polygon is not limited to 4 corners.
- Bug fix: file is empty or ObsID does not match

### stk\_build

• upper limit of verbose range in parameter file is 5

#### stk\_count

- new parameter: echo determines if the retrieved value is printed to the screen in addition to being stored in the parameter file
- upper limit of verbose range in parameter file is 5

#### stk\_read\_num

- new parameter: echo determines if the retrieved value is printed to the screen in addition to being stored in the parameter file
- upper limit of verbose range in parameter file is 5

#### tg\_create\_mask

- Mask geometric parameters have been changed to allow floating-point regions. This is to support better extractions in crowded fields when fractional pixels are important for the zero order radius or diffraction arm widths.
- Bug fix: handling of width and radius parameter; values in the output R column are expected to be different by up to half a pixel. Changed datatype of those parameters from integer to read/float.

#### tg\_resolve\_events

- The spatial filter mask now supports polygonal regions instead of just boxes. This has no impact on pipeline processing, which will still use default regions from tg\_create\_mask. The enhancement is to support development of better background region definition and HEG–MEG extraction regions which can be customized near zero order to avoid overlap to shorter wavelengths.
- The coordinate tranformation errors have been fixed for cases when the zero order instantaneous (or mean) position is off the detector array. Since grating event coordinates have their origin at the zero order, this deficiency prevented proper assignment of wavelengths and orders. (There is only one observation to date for which this was intentionally the case: Sco X-1, offpointed in order to avoid the risk of detector damage.)

### tgdetect

- Bug fix: naming of temporary files works correctly
- new parameters for temporary and diagnostic files: temproot, keeptemp, and keepexit
- added "double" checking: will still output duplicate srcs, but the double flag/column will be set for one of them to make removing the duplicates easier (change made in celldetect)

#### tgextract

• Source spectral binning region max and min in the cross-dispersion direction were adjusted (to a half-width of 0.002 degrees) to compensate for an anomalous residual rotation of the LETG spectrum when imaged onto the ACIS-S array. The expansion of the region will prevent excessive clipping of the spectrum.

The cause of the residual rotation is under investigation. When resolved, we will revert to the previous default half–width of 0.00133 degrees.

The LETG/ACIS–S rotation compensation technical note (http://space.mit.edu/CXC/docs/technote\_2004.06.11\_la\_rotation.ps.gz) describes the problem in more detail.

### vtpdetect

- Bug fix: fixed problem with WCS not getting attached to column
- Bug fix: changed ASCII region output to be compatible with dmcopy & ds9
- $\bullet$  added TLMIN/TLMAX keywords to output POS column
- Bug fix: added warning when recursive routine goes into an infinite loop with event list input

#### wavdetect

- some speed improvement in FFT convolutions
- Bug fix: will exit with non-zero status if the tool fails

#### wrecon

• added TLMIN/TLMAX keywords to output POS column

# **Parameter Files**

Since some of the parameter files have new entries or default values, it is recommended that users delete all old parameter files before running CIAO 3.2. The FAQ on cleaning up parameter files has more information.

#### acis\_process\_events

• unneeded cc\_chipy parameter removed

#### acisreadcorr

• Bug fix: default value of bgroup parameter is changed to 10

#### ardlib

• To support the spatial contamination model, the ardlib.par AXAF\_ACIS\_CONTAM\_FILE parameter has been replaced by a set of 10 parameters (one per CCD): AXAF\_ACIS0\_CONTAM\_FILE to AXAF\_ACIS9\_CONTAM\_FILE. The default value of these parameters is CALDB.

#### arestore

• changed order of parameters make consistent with other tools

#### asphist

• The speed improvement updates required replacing the max\_bin parameter with two new parameters: delta\_xy and delta\_roll.

#### csmooth

• Bug fix: bad redirect in the sclmax parameter ()sclmax changed to )sclmin)

#### destreak

• implemented a maximal-row-loss algorithm; new parameter: max\_rowloss\_fraction controls the max number of rows per frame to be streaking.

### dmcoords

• A new parameter (asolfile) was added to use the aspect solution files to account for the geometry drifts in the calculations.

### dmdiff

• Many new parameters have been added to support the extended comparisons. See the dmdiff ahelp file for a complete list.

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• change to randseed parameter limits to allow value of 0

### dmhistory

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• new parameter (binning) to control image histogram binning specification

## psf\_project\_ray

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### tgdetect

• new parameters for temporary and diagnostic files: temproot, keeptemp, and keepexit

# Data Model

## Ability to perform weighted binning

• In the DM syntax, we have always been able to make a 2D image by picking two columns from a table (e.g. an event list):

unix% dmcopy "evt.fits[ccd\_id=3][bin time=1000:2000:0.4,energy=300:1000:2]" time\_energy.img

This defines a 2500 x 350 pixel image where the x-axis is time in 0.4 second pixels and the y-axis is energy in 2 eV pixels. For every row in the table which passes the filter (e.g.  $ccd_id=3$ ), we add 1 to the relevant pixel in the image.

In the new "weighted binning", you can specify a "weighting column" from the table by adding its name to the end of the binning spec, preceded by a semicolon:

unix% dmcopy "evt.fits[ccd\_id=3][bin x=4000:5000:2,y=3000:4000:2;*energy*]" subset.img

This means that instead of adding 1 to a pixel for each row that passes, we add the value of the energy column in that row. (Note that the energy column has a floating–point data type, and the output image is automatically forced to that type).

This will be particularly useful with tables which are already histograms or tabulated functions: suppose you have a table with colums WAVELENGTH and FLUX; this will let you make a 1D FITS image from the table at any desired resolution, for use with spectral programs that work on images rather than tables. Or, suppose you have a table which gives a grid of parameter values NH, ALPHA and goodness–of–fit values CHI2:

unix% dmcopy "grid.table[bin nh=19.0:22.0:0.1,alpha=-2.0:2.0:0.05;chi2]" grid.image

### **Improved Keyword Support**

• Support for FITS WCS Rotation keywords (CROTA2, PC, CD) and unoffical IRAF keywords (LTM, LTV)

All X-ray sky images have north at top; since all X-ray satellites perform aspect reconstruction from event data, we have the freedom to reorient the pixels. In contrast, optical astronomical images retain their original pixelization and use rotation information in the WCS header keywords to specify the alignment between north and the pixel x-axis. We didn't handle this correctly in earlier releases of CIAO; now we correctly propagate the rotation metadata.

Users should note, however, that region files using the ellipse and rotbox shapes still define the rotation angle in those shapes as being with respect to pixel x, not celestial east, and so are not portable between different datasets.

We have also added support to recognize IRAF's nonstandard LTM and LTV coordinate keywords to define physical coordinate systems.

### **Virtual Column Filtering/Binning for Tables**

• It is now possible to filter on a coordinate as well as a column. For example:

unix% dmcopy "evt.fits[phi=0.00:0.02]" onaxis.fits

selects only photons within 0.02 deg of the axis. Using "dmlist evt.fits cols" you can see that phi is not one of the columns in the event list, it's just a coordinate, which is defined in the header by a coordinate transformation on the detx and dety columns. The filtering applies the transform "on-the-fly." (In CIAO 3.1 it was possible to do "dmlist "evt.fits[cols phi]" data", but you couldn't filter on a coordinate.)

*Warning 1:* you can't omitting the lower bound, e.g. [phi=:0.02], because the lower/upper bound range of coordinates arenot currently stored in the file header.

*Warning 2:* filtering on 2D regions (e.g. of RA and Dec) is problematic for reasons to do with the spherical nature of these coordinates. Our regions (circle, polygon, etc.) are always defined as shapes on the tangent plane; you should continue to use

unix% dmcopy "evt.fits[sky=circle(14:13:11.2,-00:02:23,5')]" ...

rather than [eqpos=circle...] or [(ra,dec)=circle...]. 1D filters in degrees, such as "ra=270.0:275.0", are fine.

### **Other Enhancements**

• Allow users to type-cast images using [option type=i4] syntax:

unix% dmcopy "integer\_image.fits[opt type=r4]" real\_image.fits

• Some improvements for handling image cubes, e.g.

unix% dmcopy "cube.fits[#2=12][bin #1,#3]"

makes a 2D image from plane 12 of the second axis of a 3D image.

### **Bug Fixes**

- The Data Model bug page lists all currently known bugs, as well as those fixed in CIAO 3.2.
- dmlist will display logical values as T/F even when opt=raw (instead of 1/0)
- correctly handle files with only NULL and GTI blocks
- print bit columns properly in array mode (fix affects dmlist)
- print all elements of a string array when opt=array (fix affects dmlist)
- allow numbered image axes (e.g. #1,#2) grouping using bin or filter syntax
- proper support for variable–length arrays: if you dmcopy an RMF file, it will retain the correct FITS variable–length format
- fixed bug in column selection on second component of a coordinate vector, e.g. evt.fits[cols y]
- set pixels outside of region filter to NULL for >2D images
- keep the LONP/LAT coordinate system on DETX/DETY (THETA, PHI) when copying
- allow image enlargement using bin filter; restrict range/region filters to image field
- catch bad filter syntax, such as "a:b:c" when not binning
- error out when given a dataset with no blocks. This was reported as a tool crashing when given a directory instead of a filename, which it interpreted as a dataset with no blocks.
- propagate subspace region blocks like GTI blocks: if an event file includes keywords saying that another block (HDU) in the file is a region describing the area covered by the event file, that block will get copied along with the event block

# Sherpa

## **Convolution with PSF Models**

- Improvements to the file-based PSF models (fpsfld, fpsf2d) and the point-source models (ptsrcld, ptsrc2d) have been made.
- The ptsrc1d and ptsrc2d have been updated to accommodate both FITS and ASCII files.
- A bug in file-based PSF convolution (fpsf models) has been corrected.

## **Fitting Methods**

• The Levenberg-Marquardt fitting method has improved handling of problem datasets.

## **Error Estimation**

- Improvements have been made to the projection method to handle the case of overflow and underflow errors:
  - the projection method now catches overflows and underflows, and calculates the errors for the remaining parameters rather than immediately exiting, and
  - the allowable range on many model parameters has been restricted to reduce the chances of overflow occurring.
- The error-estimation commands can now be run without first performing a fit. This allows users to load in a previous fitting session (saved with the "SAVE" command) and continue their analysis. This only holds for the Sherpa commands (e.g. PROJECTION, INTERVAL-UNCERTAINTY) and not the S-Lang versions (e.g. run\_proj(), run\_intunc()).

# **S–Lang Integration**

- In a number of cases, cleaned up extraneous output from S-Lang functions. These include functions such as get\_fit(), run\_proj(), and sherpa\_eval().
- Sherpa can now be used with the S-Lang "require()" function; this change is in conjunction with the S-Lang module changes for provide and require.

## **Miscellaneous Changes**

- In the jdpileup model, some default parameter values and limits have been changed:
  - $\blacklozenge$  value of alpha changed from 0.95 to 0.5
  - $\bullet$  f is now that the function for the function of the functi
  - the upper limit for the n parameter is 100 (was 2048)
  - ♦ the lower limit for the g0, ftime, and fracexp parameters is 1e-120 (was 0)
  - value of fracexp changed from 1 to 0.987
- The CIAO 3.2 version of Sherpa links in the XSPEC v11.3.1 model library, and includes interfaces to the new models in this library. The XSpec model function help file lists the models which are not available in CIAO 3.2.
- The CIAO 3.2 version of Sherpa makes use of version 0.4.0 of the S–Lang XPA module, in its CIAO 3 compatibility mode. This means that the ds9 package can not be used with Sherpa.
- Sherpa no longer overrides the ChIPS redraw setting. This means that ChIPS commands will be executed immediately, so users no longer need to issue the REDRAW command.

### **Bug Fixes**

- The Sherpa bug pages list all currently known bugs, as well as those fixed in CIAO 3.2.
- get\_fit always returned goodness-of-fit information for all datasets
- $sherpa_eval()$  does not return -1 on a failed use command
- due to a round–off/casting problem, creating an ARF in double precision causes the source model to contain 0's when it shouldn't
- if a parameter value is at the maximum boundary, running projection on it may result in a core dump (Linux only)

# Libraries

#### ardlib

• To support the spatial contamination model, the ardlib.par AXAF\_ACIS\_CONTAM\_FILE parameter has been replaced by a set of 10 parameters (one per CCD): AXAF\_ACIS0\_CONTAM\_FILE to AXAF\_ACIS9\_CONTAM\_FILE. The default value of these parameters is CALDB.

#### cxcparam

- use strncpy to make sure string buffer doesn't overflow
- parser did not set error condition when parameter file had too many field or field was not parseable
- crash when bad parameter file lived in the system part of the PFILES search path, so it could not be copied to the user path

### dslib

- changed length of ds\_autoname string routines
- corrected "kernal" -> "kernel" spelling
- tweaked autonaming to avoid \_extname when filename begins with an underscore
- event file names containing "qp" that are not qpoe files caused crash
- fixed autonaming so that empty strings do not cause array bound reads
- added range to WCS struct to allow detect tools to include TLMINTLMAX on output columns

#### errlib

• Updated version to reflect CIAO 3.2.

### fl\_acisio

• allow afterglow bits to be set from hotpixel-ammended badpixel file

### grplib

- allow BIN to work on ascending or descending data (e.g. to bin wavelengths that decrease)
- BIN\_WIDTH always working on row/channel instead of on xcolumn data
- problem with bin=::<value>
- problem with BIN\_FILE when non-integer xcolumn is used
- last element now marked as quality=2 when SNR grouping (fix to the group library)
- use grouptypeval when binspec is unspecified
- typo in the error message for grouptype=NUM\_BINS

## hdrlib

• Updated version to reflect CIAO 3.2.

### mtl\_libs

- SEGVs when setting bits in a file (Linux only)
- allow .[0-9]+ to be parsed as a number without requiring lead 0
- casting when converting numeric values from one format to another with (int)
- <array>\*\*<scalar> giving array all equal to first element/pixel
- not using correct array index in expressions (e.g. sum=(phas[2]+phas[3])). Note that setting individual elements in an array still does not work, e.g. (phas[2]=32)
- minor speed improvements
- table in second extension opened instead of image

### regionlib

- ASCII output when region contains Points
- problem writing ASCII files with rectangles and circles

### rmflib

• crash if input FEF filename contained an error

### stklib

- changed formatting internally for pgrid, rgrid, and lgrid strings
- added igrid to force integer type grid
- problem with lgrid giving 1 more than necessary
- calling stk\_delete\_current when current==0 caused crash
- stk\_build on an @ file that doesn't exist will now return a NULL stack pointer
- stk\_change\_current exhibited memory problems

### tcdio

- enabled support for 1D images
- added warning when image has 0 axes

### tcdlib

• minor code cleanup that gives some noticeable speed improvements in FFT convolutions

# **S-Lang Modules**

## Changes for provide and require

• To keep up to date with S-lang enhancements, CIAO modules were upgraded so that they could be loaded with "provide" and "require"; this also affected the varmm and chips modules. The changes were made to be backwards-compatible such that loading a module with "require" is the same as loading it with "import". Other changes were made to varmm so that varmm-based software automatically set proper slang load paths and loaded the slsh resource file (slsh.rc). The slsh resource file will not be loaded by varmm explicitly if the the file is not located in the

\$ASCDS\_INSTALL/etc directory.

### **Module Conflicts**

• Bug fix: checks error lib before trying to initialize. Reported as "If functions are called in a certain order, a benign warning is printed to the screen" (fix to caldb module)

#### caldb

- checks error lib before trying to initialize.
- fixed array over-run
- changes for provide/require; back-compatible with import

#### group

- fixed array over-run
- problem where grpBinFile() caused crash (seen on os-x)
- changes for provide/require; back-compatible with import

#### paramio

- paramopen take NULL for the file name and it will get it from the argv list. This enables "@@" opening and removes the need to have taskname in both argv array and individual input.
- paccess now has two modes of operation. The new mode is enabled when the first argument is a Param\_File\_Type variable, in which case the function returns 1 if the second argument is the name of a variable defined in the parameter file, or 0 otherwise.
- fixed array over-run
- changes for provide/require; back-compatible with import

### pixlib

- added optional "dtheta" argument to pix\_set\_simoffset() routine
- added a new routine, pix\_fpc\_to\_msc() which is a wrapper around fpc\_to\_mnc() and mnc\_to\_msc() to get off-axis angles
- fixed array over-run
- changes for provide/require; back-compatible with import

### region

- new regExtent function: calculates the bounding box of a region
- new regRegionString function: complements regPrintRegion routine
- fixed array over-run
- changes for provide/require; back-compatible with import

#### stack

- return NULL when @ file is not valid
- stk\_read\_next and stk\_read\_num return NULL at the end of a stack
- corrected usage message of stk\_expand\_n
- fixed array over-run
- changes for provide/require; back-compatible with import

# **Analysis Scripts**

The CIAO contributed scripts package is available from the CIAO Scripts page and is considered a required part of the installation. The script installation instructions explain where the scripts should be unpacked within the CIAO directory tree.

## **New Additions**

- Four scripts have been moved from the CIAO distribution to the scripts package:
  - ♦ psextract
  - ♦ acis\_fef\_lookup
  - ♦ add\_grating\_orders
  - ♦ add\_grating\_spectra

Users that follow the script installation instructions will not notice any difference in using these scripts between CIAO 3.1 and 3.2.

### Changes for provide and require

• A number of scripts were updated in conjunction with the "provide" and "require" S-lang enhancement; see the S-Lang Modules section of these release notes for details. Again, this change will be transparent to users.

### get\_sky\_limits

• use require for loading modules instead of import

#### paramest.sl

- updated for fixes to run\_regproj/run\_regunc() and run\_intproj/run\_intunc(); removes an error message seen in CIAO 3.1
- the script can be loaded via require("paramest");
- use require for loading modules instead of import

### setplot.sl

• use require for loading modules instead of import

### sherpa\_plotfns.sl

- added a provide("sherpa\_plotfns"); line
- use require for loading modules instead of import
- Use get\_fit() rather than get\_goodness()

### sherpa\_utils.sl

- the script can be loaded via require("sherpa\_utils");
- use require for loading modules instead of import

### tg\_osort\_img

- new path to ISIS: \$ASCDS\_INSTALL/contrib/share/
- removed '() = mk\_osum\_img...; 'since there is no return value from mk\_osum\_img()

# **Documentation**

### **New Threads**

- Creating ACIS RMFs with mkacisrmf
- Identify ACIS Hot Pixels and Cosmic Ray Afterglows

## **Updated Threads**

- Extract ACIS Spectra for Pointlike Sources and Make RMFs and ARFs
- Step-by-Step Guide to Creating ACIS Spectra for Pointlike Sources
- Extracting Extended Source Spectra and Responses
- Coadding Spectra and Weighted Responses

## Why Topics

- New: Creating ACIS RMFs in CIAO 3.2
- Updated: ACIS QE Degradation

## **CALDB Website**

• The Chandra Calibration Database website has been updated for the CALDB 3.0.0 release.

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URL: http://cxc.harvard.edu/ciao3.4/releasenotes/ciao 3.2 release.html Last modified: 11 August 2006