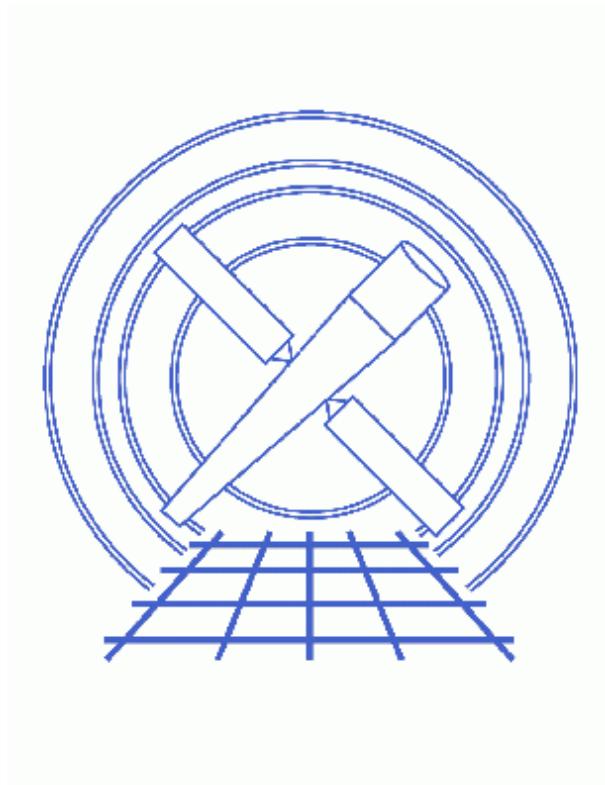


# Estimate Source Counts in an Image



## ***CIAO 3.4 Science Threads***

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# Estimate Source Counts in an Image

## *CIAO 3.4 Science Threads*

### Overview

**Last Update:** 1 Dec 2006 – updated for CIAO 3.4: CIAO version in warning

**Synopsis:**

A quick means for estimating source counts, which may be useful as a first step in a more detailed analysis procedure. The thread is not intended to provide accurate photometric results, for which careful exposure and PSF corrections are necessary.

**Purpose:**

To estimate net source counts in user-defined regions of event lists or image files.

**Read this thread if:**

you would like to find the number of counts in an HRC or ACIS imaging observations; running this thread on LETG and HETG observations is *not* recommended.

**Related Links:**

- Analysis Guide: [HRC Imaging](#)

Proceed to the [HTML](#) or [hardcopy \(PDF: A4 | letter\)](#) version of the thread.

---

## Getting Started

**Sample ObsID used:** 1838 (ACIS-S, G21.5-09)

**File types needed:** evt2

There are essentially two steps required to estimate net counts:

1. Define source and background regions. The ds9 display tool is recommended for interactively creating these regions. Alternatively, a source list (e.g. the output of one of the [detect tools](#) or a list of objects from an astronomical catalog) may be used.

## Estimate Source Counts in an Image – CIAO 3.4

2. Use the CIAO tool `dmextract` to determine counts and number of pixels for each region and to compute net counts for the source regions that have associated background regions.

In the following examples, we restrict the energy range of the events:

```
unix% dmcoppy "acisf01838N001_evt2.fits[energy=300:8000]" acis_1838_evt2.fits
```

---

## Estimating Source Counts


`dmextract` can be used to bin on vector columns, such as `sky`. This allows it to perform spatial extractions in regions in order to extract counts. Optional background files and background regions may also be input, in which case `dmextract` will compute net counts as well. Errors can be computed using either Gaussian or Poisson statistics or input via a variance map. For more details on all of these options, refer to [ahelp dmextract](#).

### A Simple Example

For another simple method of finding source counts, see the [Using Analysis Scripts](#) section of the [SAOImage ds9](#) thread.

Display the file:

```
unix% ds9 acis_1838_evt2.fits &
```

and create regions by left-clicking on the image. Then use the the "Get Info..." option in the Region menu to find out the dimensions of the regions. More information on creating and modifying regions is given [below](#). Consider a single source region `circle(4072.96,4248.00,20)` and background region `annulus(4072.96,4248.00,86,114)` as displayed in [Figure 1](#) .

To extract counts in the source region and compute net counts:

```
unix% punlearn dmextract
unix% pset dmextract infile="acis_1838_evt2.fits[bin sky=circle(4072.96,4248.00,20)]"
unix% pset dmextract outfile=1838_simple.fits
unix% pset dmextract bkg="acis_1838_evt2.fits[bin sky=annulus(4072.96,4248.00,86,114)]"
unix% dmextract
Input event file (acis_1838_evt2.fits[bin sky=circle(4072.96,4248.00,20)]):
Enter output file name (1838_simple.fits):
```

The contents of the parameter file may be checked using [plist dmextract](#).

The output may be examined using [dmlist](#):

```
unix% dmlist "1838_simple.fits[HISTOGRAM]" cols,data

-----
Columns for Table Block HISTOGRAM
-----

ColNo  Name                Unit          Type          Range          Position
  1    sky(X,Y)           pixel         Real8         -Inf:+Inf      Position
  2    EQPOS(RA,Dec)     deg           Real8         -360.0:        360.0         Position
  3    SHAPE              String[16]    String[16]    -Inf:+Inf      Region shape type
  4    R[2]               pixel         Real8(2)      -Inf:+Inf      Radius
```

## Estimate Source Counts in an Image – CIAO 3.4

5	ROTANG[2]	pixel	Real8(2)	-Inf:+Inf	Angle
6	COMPONENT		Int2	-	Component number
7	AREA	pixel**2	Real4	-Inf:+Inf	Area of extraction
8	EXPOSURE	s	Real8	-Inf:+Inf	Exposure time of source
9	COUNTS	count	Real8	-Inf:+Inf	Counts
10	ERR_COUNTS	count	Real8	-Inf:+Inf	Error on counts
11	COUNT_RATE	count/s	Real8	-Inf:+Inf	Rate
12	COUNT_RATE_ERR	count/s	Real8	-Inf:+Inf	Rate Error
13	BG_AREA	pixel**2	Real8	-Inf:+Inf	Background Area of Extr
14	BG_EXPOSURE	s	Real8	-Inf:+Inf	Exposure time of backgr
15	BG_ERR	count	Real8	-Inf:+Inf	Error on Background cou
16	BG_COUNTS	count	Real8	-Inf:+Inf	Background Counts
17	BG_RATE	count/s	Real8	-Inf:+Inf	Background Rate
18	BG_SUR_BRI	count/pixel**2	Real8	-Inf:+Inf	Background Counts per
19	BG_SUR_BRI_ERR	count/pixel**2	Real8	-Inf:+Inf	Error on background c
20	NET_COUNTS	count	Real8	-Inf:+Inf	Net Counts
21	NET_ERR	count	Real8	-Inf:+Inf	Error on Net Counts
22	NET_RATE	count/s	Real8	-Inf:+Inf	Net Count Rate
23	ERR_RATE	count/s	Real8	-Inf:+Inf	Error Rate
24	SUR_BRI	count/pixel**2	Real8	-Inf:+Inf	Net Counts per square
25	SUR_BRI_ERR	count/pixel**2	Real8	-Inf:+Inf	Error on net counts p

-----

World Coord Transforms for Columns in Table Block HISTOGRAM

-----

ColNo	Name	
7:	CEL_AREA	= +0 [arcsec**2] +0.2421 * (AREA -0)
18:	BG_CEL_BRI	= +0 [count/arcsec**2] +4.1311 * (BG_SUR_BRI -0)
19:	BG_CEL_BRI_ERR	= +0 [count/arcsec**2] +4.1311 * (BG_SUR_BRI_ERR -0)
24:	CEL_BRI	= +0 [count/arcsec**2] +4.1311 * (SUR_BRI -0)
25:	CEL_BRI_ERR	= +0 [count/arcsec**2] +4.1311 * (SUR_BRI_ERR -0)

-----

Data for Table Block HISTOGRAM

-----

ROW	sky(X,Y)	EQPOS(RA,Dec)	SHAPE
1	( 4072.960, 4248.0) ( 278.3893176799, -10.5692082869) Circle	[ 20.0	1
	0] [ 0	0]	
	1256.6370849609 7854.4664748687	7523.0	
	86.7352292901 0.95779898279160	0	
	17592.917968750 7854.4664748687	24.9198715888	
	621.0 0.07906329500380 0.03529829452414		
	0.00141647176625 7478.6428551599	86.7534919622	
	0.95215160432459 0.01104511582547	5.9513147787	
	0.06903623408891		

The counts information is given in the last section of the output.

Alternatively, `prism` may also be used to examine the output:

```
unix% prism 1838_simple.fits &
```

as shown in Figure 2 .

## Regions vs. Stacks of Regions

Region descriptors may also be input via files, rather than typed on the command line:

```
unix% dmextract infile="acis_1838_evt2.fits[bin sky=region(source.reg)]" outfile=1838_simple_2.fits \
  bkg="acis_1838_evt2.fits[bin sky=region(background.reg)]"
```

where

```
unix% cat source.reg
circle(4072.96,4248.00,20)

unix% cat background.reg
annulus(4072.96,4248.00,86,114)
```

However, if you want to extract counts from a number of source regions contained in a single file, then you must input the region file as a stack. If you have two region files:

```
unix% more stack.reg
circle(4072.96,4248.00,20)
circle(4244,4094,6)

unix% more stackbgd.reg
annulus(4072.96,4248.00,40,60)
annulus(4244,4094,10,30)
```

then to compute net counts in each region separately:

```
unix% dmextract infile="acis_1838_evt2.fits[bin sky=@stack.reg]" outfile=1838_stack.fits \
  bkg="acis_1838_evt2.fits[bin sky=@stackbgd.reg]"
```

Examine the output as before:

```
unix% dmlist "1838_stack.fits[cols counts,area,bg_counts,bg_area,net_counts,net_err]" data
-----
Data for Table Block HISTOGRAM
-----
```

ROW	COUNTS	AREA	BG_COUNTS	BG_AREA	NET_COUNTS
1	7523.0	1256.6370849609	5072.0	6283.1855468750	6508.59998
2	94.0	113.0973358154	32.0	2513.2741699219	92.56000

If you use the `sky=region(stack.reg)` syntax:

```
unix% dmextract infile="acis_1838_evt2.fits[bin sky=region(stack.reg)]" outfile=1838_region.fits \
  bkg="acis_1838_evt2.fits[bin sky=region(stackbgd.reg)]"
```

`dmextract` will interpret the list of regions as a single, connected region. Using this syntax with `verbose > 0` will print a warning:

```
# dmextract (CIAO 3.4): dsDMEXTRACTREGCOMPWERR -- WARNING:Region #1
contains more than 1 component. Only the first component will be
described in the region columns of the output file.
```

This returns a single row with the sum of the counts in all the individual regions:

```
unix% dmlist "1838_region.fits[cols counts,area,bg_counts,bg_area,net_counts,net_err]" data
```

## Estimate Source Counts in an Image – CIAO 3.4

```
-----
Data for Table Block HISTOGRAM
-----
```

ROW	COUNTS	AREA	BG_COUNTS	BG_AREA	NET_COUNTS
1	7617.0	1369.7343750	5104.0	8796.4589843750	6822.0

### Exposure Corrections

Exposure maps may be applied to both source and background regions. In this case, in the calculation of the net counts from the source, the background counts are normalized not only by the ratio of the *areas* of the source and background regions, but also by the ratio of the *mean exposures* in the source and background regions (be sure to read the caveat at the end of this section). The following threads give complete instructions on generating exposure maps:

- [Create an ACIS Exposure Map for a Single Chip](#)
- [Create an ACIS Exposure Map for Multiple Chips](#)
- [Create an HRC Exposure Map](#)

Both normalized [ $\text{cm}^2 \cdot \text{counts}/\text{photon}$ ] and unnormalized [ $\text{cm}^2 \cdot \text{sec} \cdot \text{counts}/\text{photon}$ ] exposure maps may be used as input to `dmextract`. Bin the data to a FITS image and generate exposure maps that are congruent to that image. To isolate the S3 chip in this dataset:

```
unix% dmcopyp "acis_1838_evt2.fits[bin x=3696.5:4720.5:1,y=3872.5:4896.5:1]" 1838_s3.fits
```

Use `xygrid=3696.5:4720.5:#1024,3872.5:4896.5:#1024` in the `mkexpmap` step to create an *unnormalized* exposure map, `s3_expmap.fits`, then run `dmextract` and examine the results:

```
unix% dmextract infile="1838_s3.fits[bin sky=@stack.reg]" outfile=1838_stackexp.fits \
    bkg="1838_s3.fits[bin sky=@stackbgd.reg]" exp=s3_expmap.fits bkgexp=s3_expmap.fits

unix% dmlist "1838_stackexp.fits[cols counts,area,exposure,bg_counts,bg_area,net_counts]" data

-----
Data for Table Block HISTOGRAM
-----
```

ROW	COUNTS	AREA	EXPOSURE	BG_COUNTS	BG_AREA	NET_COUNTS
1	7522.0	1252.0	7854.4664748687	5059.0	6272.0	6512.1358710676
2	93.0	113.0	7854.4664748687	31.0	2516.0	91.6077107005

The `NET_COUNTS` are calculated as the counts in the source region [`COUNTS`] minus the counts in the background region [`BG_COUNTS`] (appropriately normalized by the areas [`AREA/BG_AREA`] and the mean exposure maps in source and background regions) and divided by the mean exposure map in the source regions.

#### ***Caveat on Exposure Corrections:***

Normalizing counts by mean exposure in regions may lead to errors if there are large exposure variations in the region which are not accompanied by similar variations in counts. Consider a bright point source at the center of a large region whose exposure varies strongly near the boundaries. That loss of exposure is not reflected in the counts, which are concentrated near the point source, but would strongly affect the mean exposure. In cases such as this, it is better to flat field the image by the exposure map and than extract flat-fielded counts in the region. A variance map should also be computed and used to calculate errors in the region.

---


## Defining Source and Background Regions

Source and background regions can easily be defined interactively for small numbers of regions. If the number of sources is large, however, it may be preferable to create a source list.

### Interactive Definition

The event file may be viewed directly with ds9:

```
unix% ds9 acis_1838_evt2.fits &
```

To create a region, left-click once on the image. The default region shape is circle; to select a different shape, use the "Region -> Shape" menus in ds9. The shape must be set *before* creating the region. Click again to make the region "active;" in Figure 3 , the circular region is active, but the rectangular one is not.

To change the size of the region, click and drag on the anchor points which appear when the image is active.

To save the region:

- Create the region(s) to save.
  - Region -> File Format-> Ciao
  - Region -> File Coordinate System -> Physical
  - Region -> Save Regions... -> Save As "source.reg"
- 

### Using a Source List

The source lists output by the [detect tools](#) can also be used to define regions for estimating source counts. These regions can be read into ds9 using the "Load Regions..." option (if given a FITS file, ds9 automatically looks for a block named REGION.) The source regions are saved in the SRCLIST block in the DETECT tools output. Note that the source list must end in `.fits` to be recognized by ds9.

To display the region, either rename the block to REGION:

```
unix% dmcoppy "s3_img_src.fits[srclist][region]" sources.fits
```

*or* specify the block name when reading it into ds9:

```
Region -> Load Regions... -> s3_img.fits[srclist]
```

---

Parameters for `/home/username/cxcds_param/dmextract.par`

```
#-----
#
# DMEXTRACT -- extract columns or counts from an event list
#
#-----
```



## Estimate Source Counts in an Image – CIAO 3.4

```
infile = acis_1838_evt2.fits[bin sky=circle(4072.96,4248.00,20)] Input event file
outfile = 1838_simple.fits      Enter output file name
  (bkg = acis_1838_evt2.fits[bin sky=annulus(4072.96,4248.00,86,114)]) Background region file
  (error = gaussian)           Method for error determination(poisson|gaussian|<variance file>)
(bkgerror = gaussian)         Method for background error determination(poisson|gaussian|<variance file>)
  (bkgnorm = 1.0)              Background normalization
  (exp = )                      Exposure map image file
  (bkgexp = )                   Background exposure map image file
(sys_err = 0)                  Fixed systematic error value for SYS_ERR keyword
  (opt = phal)                  Output file type: phal
(defaults = ${ASCDS_CALIB}/cxo.mdb -> /soft/ciao/data/cxo.mdb) Instrument defaults file
  (wmap = )                     WMAP filter/binning (e.g. det=8 or default)
(clobber = no)                 OK to overwrite existing output file(s)?
(verbose = 0)                   Verbosity level
  (mode = ql)
```

---

## History

23 Dec 2004 reviewed for CIAO 3.2: no changes

19 Dec 2005 updated for CIAO 3.3: default value of dmextract error and bkgerror parameters is "gaussian"

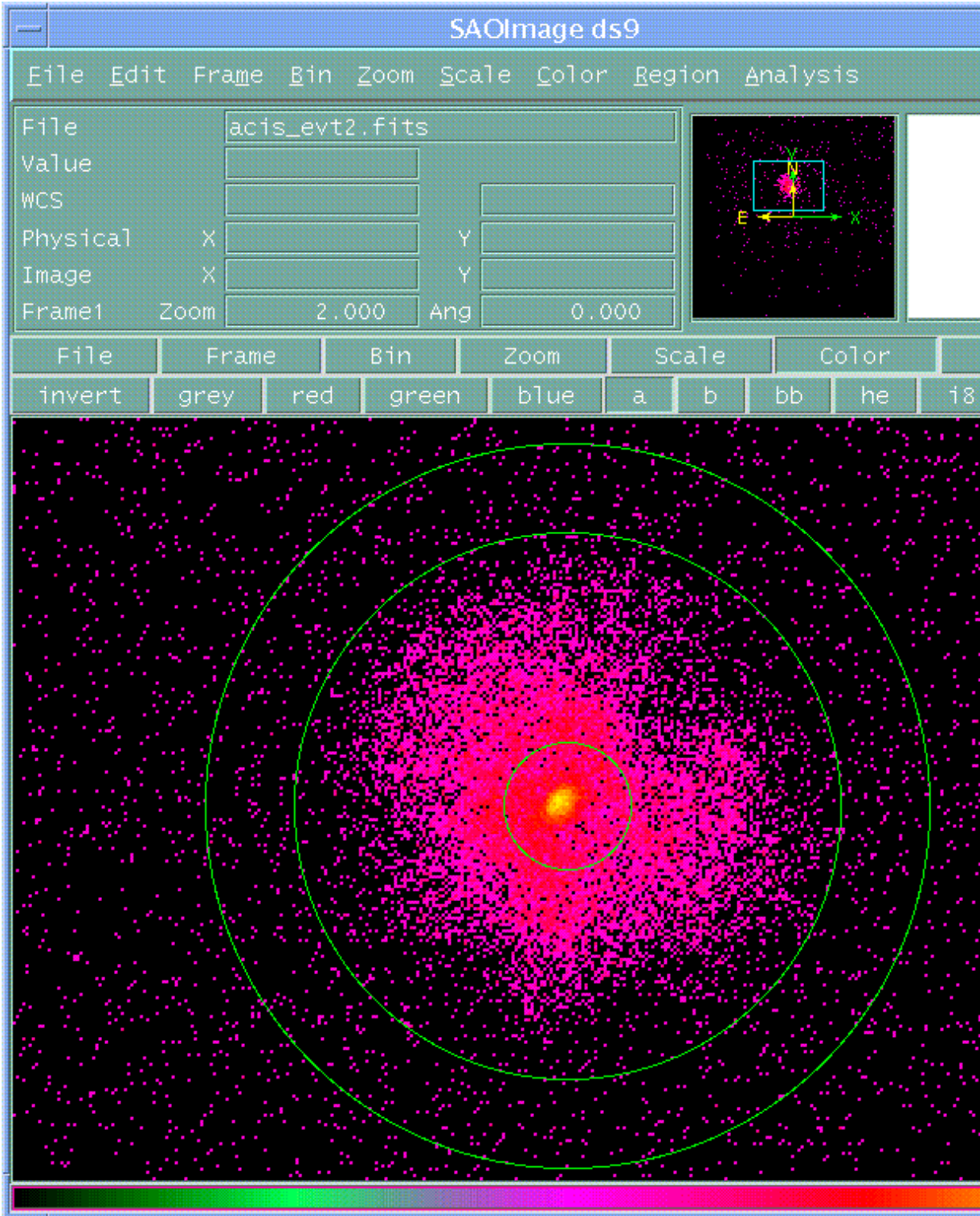
01 Dec 2006 updated for CIAO 3.4: CIAO version in warning

---

URL: [http://cxc.harvard.edu/ciao/threads/source\\_counts/](http://cxc.harvard.edu/ciao/threads/source_counts/)

Last modified: 1 Dec 2006

### Image 1: Source and Background regions



## Image 2: Viewing dmextract output in prism

prism : 1838\_simple.fits

File Edit Navigate Visualization Session Analysis

IMAGE	PRIMARY	NULL
TABLE HISTOGRAM		25 cols, 1 rows
TABLE GTI7		2 cols, 1 rows
TABLE GTI0		2 cols, 1 rows
TABLE GTI1		2 cols, 2 rows
TABLE GTI2		2 cols, 1 rows
TABLE GTI3		2 cols, 1 rows
TABLE GTI6		2 cols, 1 rows

```

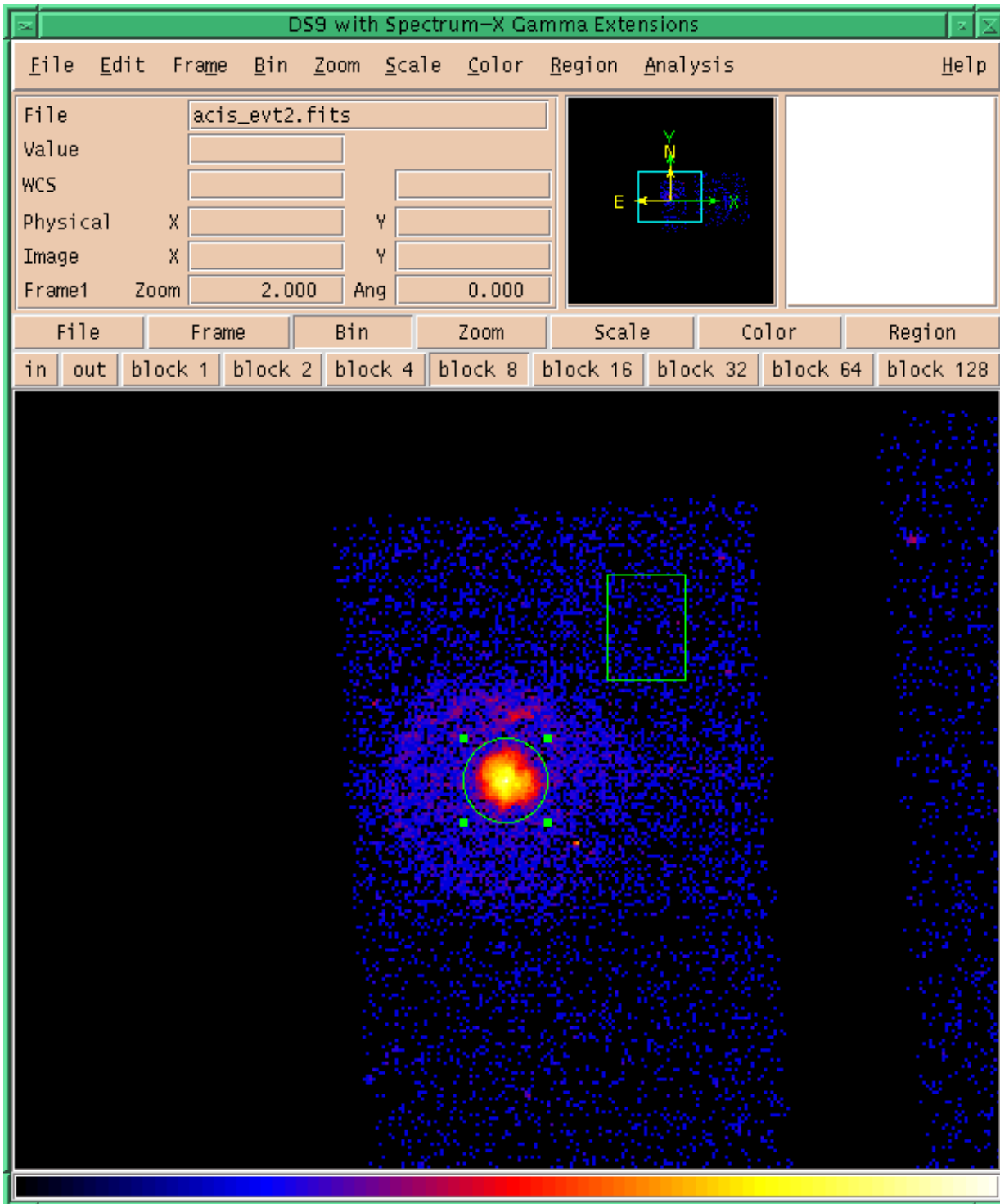
COMMENT This FITS file may contain long string keyword values
COMMENT continued over multiple keywords. The HEASARC convent
COMMENT character at the end of each substring which is then c
COMMENT on the next keyword which has the name CONTINUE.
ORIGIN ASC / Source of FITS file
CREATOR dmextract - Version CIAO3.3 / tool that created this o
REVISION 1 /
ASCDSVER CIAO 3.3 Wednesday, November 16, / String ASCDS
CHECKSUM UA5qV65nUA5nU55n / HDU checksum updated 2005-12-
DATASUM 2768198189 / data unit checksum updated 20
    
```

I	AREA	EXPOSURE	COUNTS	ERR_COUNTS	COUNT_RATE
Units	pixel**2	s	count	count	count/s
Types	float	double	double	double	double
1	1256.64	7854.4664748687	7523	86.7352292900641	0.957798982791604

View Mode: Read/Write      Displaying row 1 (1 total row)     

Mon 19-Dec 17:30:19 Loading file 1838\_simple.fits  
 Mon 19-Dec 17:30:19 Configuring Analysis Menu from file: /soft/ciao/bin/ciao.ans

### Image 3: Selecting regions in ds9



The circle is the currently-selected region (as indicated by the "handles" around the region).