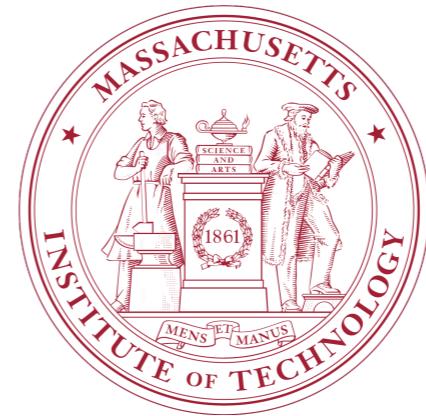
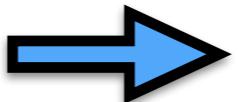


Testing Dark Matter with Galaxy Surveys

Mikhail (Misha) Ivanov
IAS —> MIT

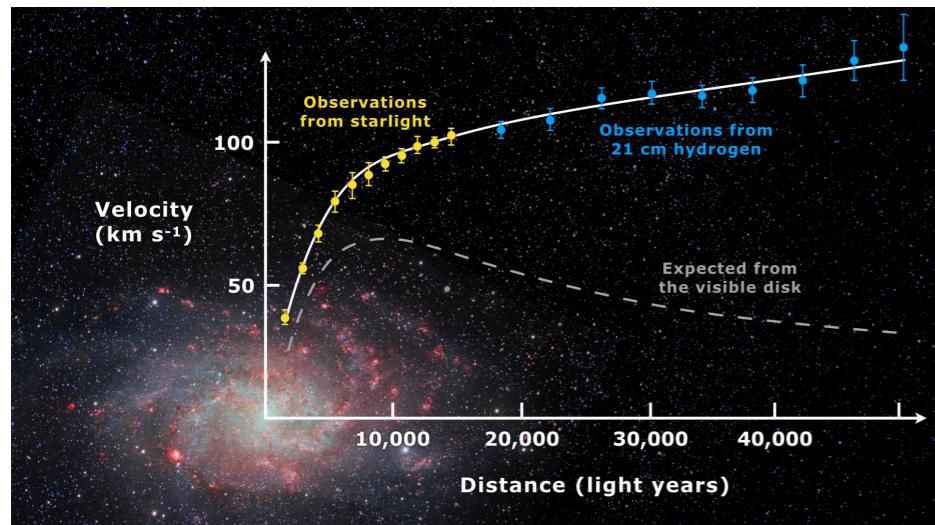


NHFP symposium, 18 Sep 2023

Thank you!

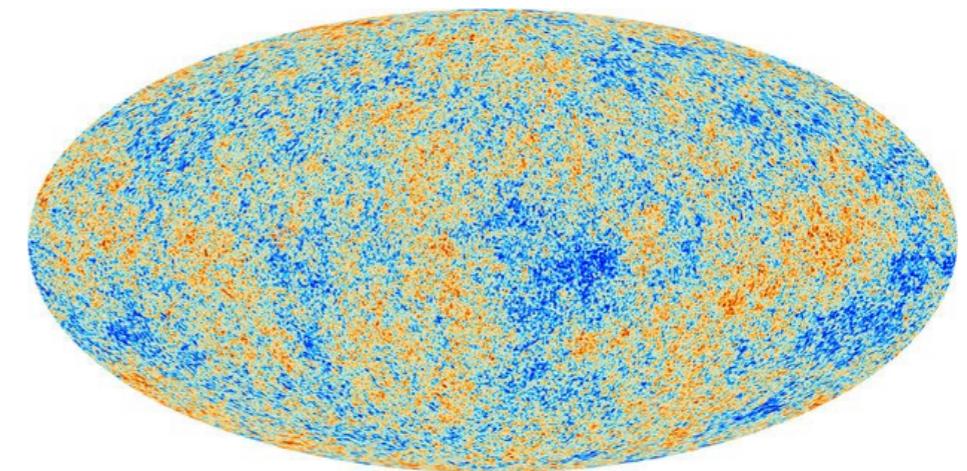


Dark Matter

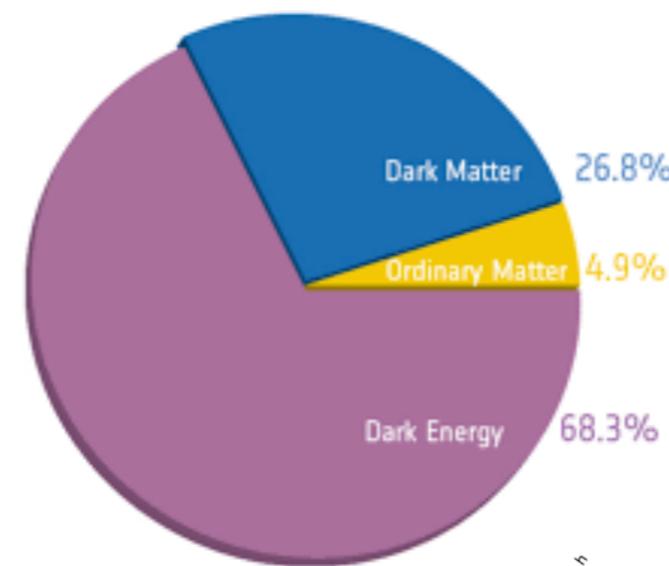


© Wikipedia

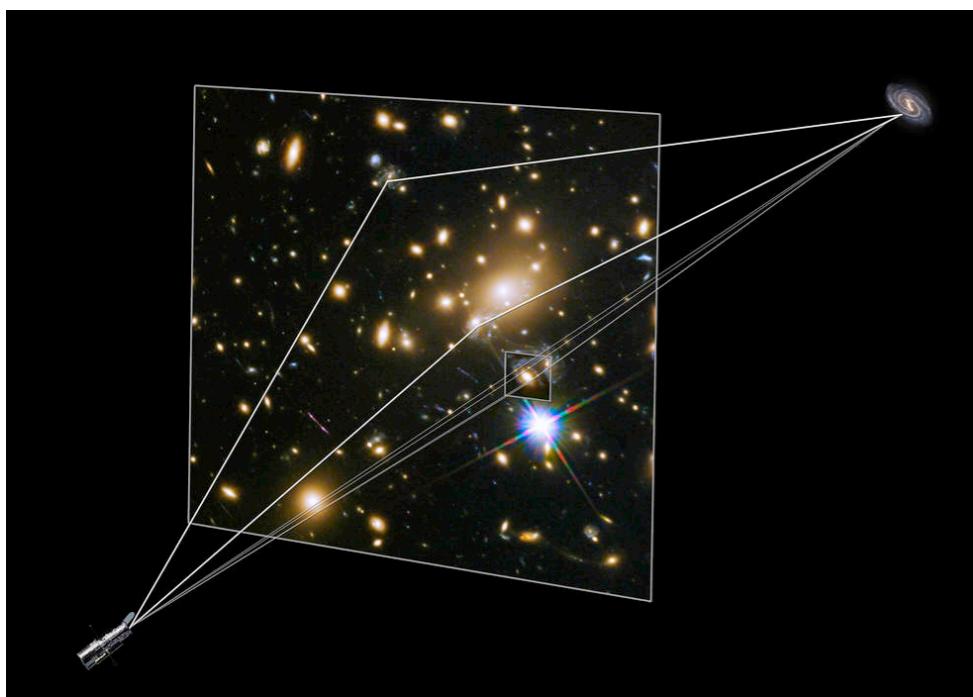
Rotation curves



Planck'18

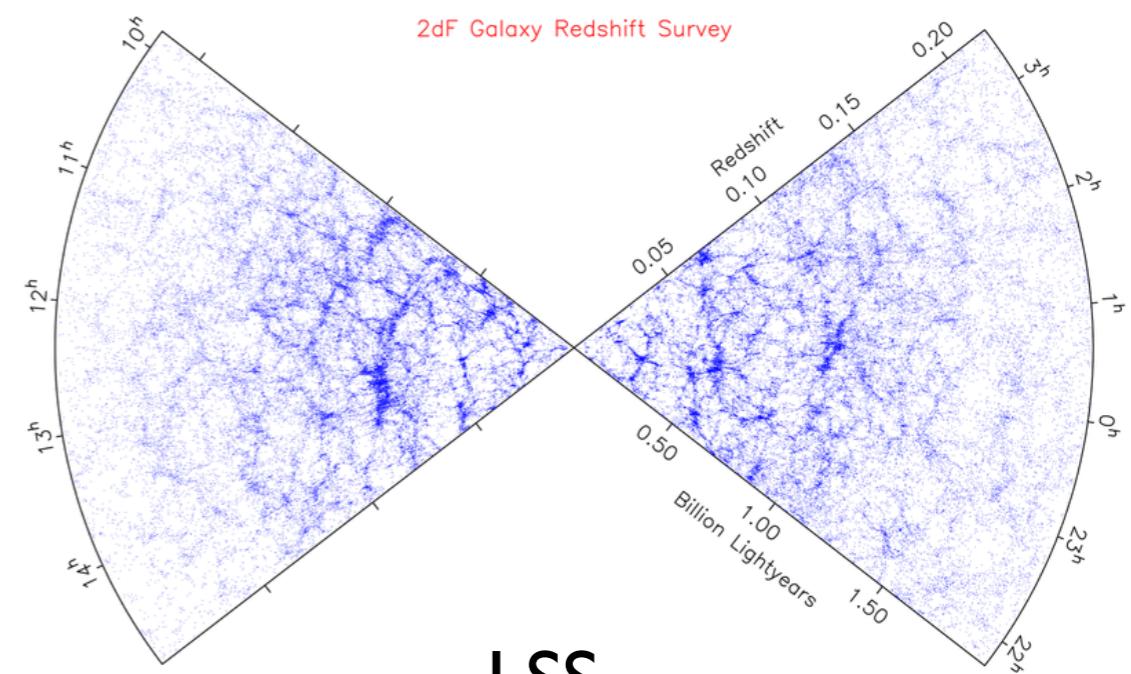


CMB



Lensing

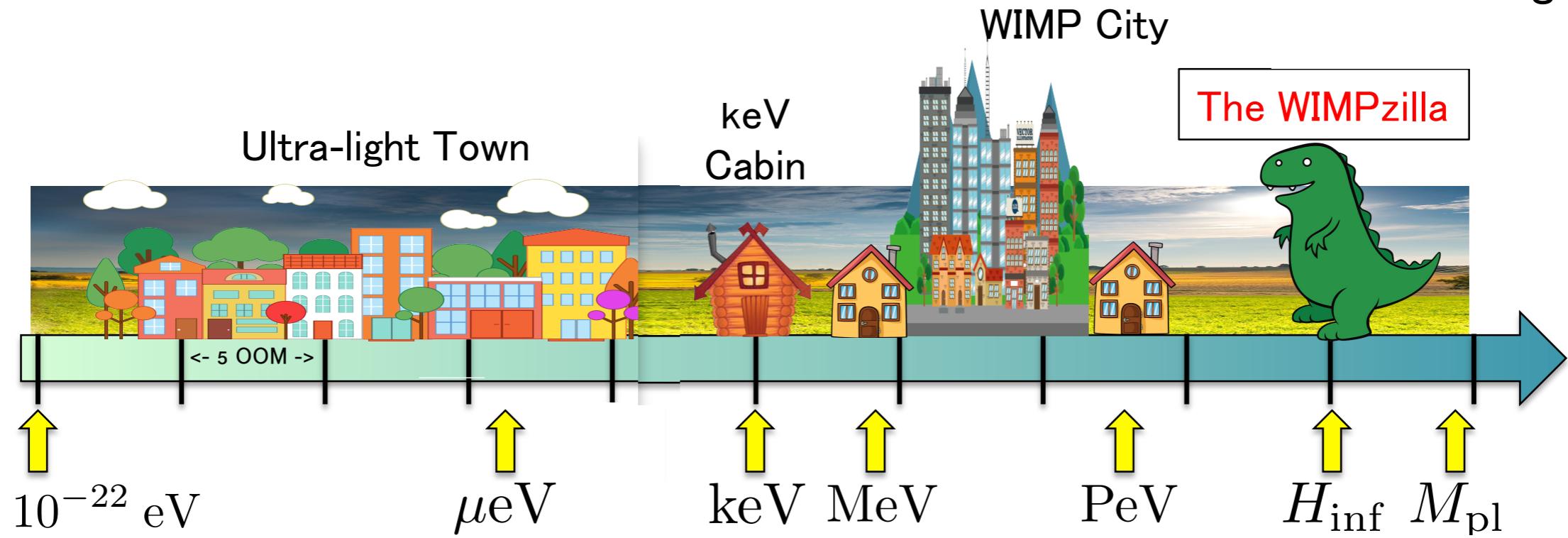
© NASA+ESA



LSS

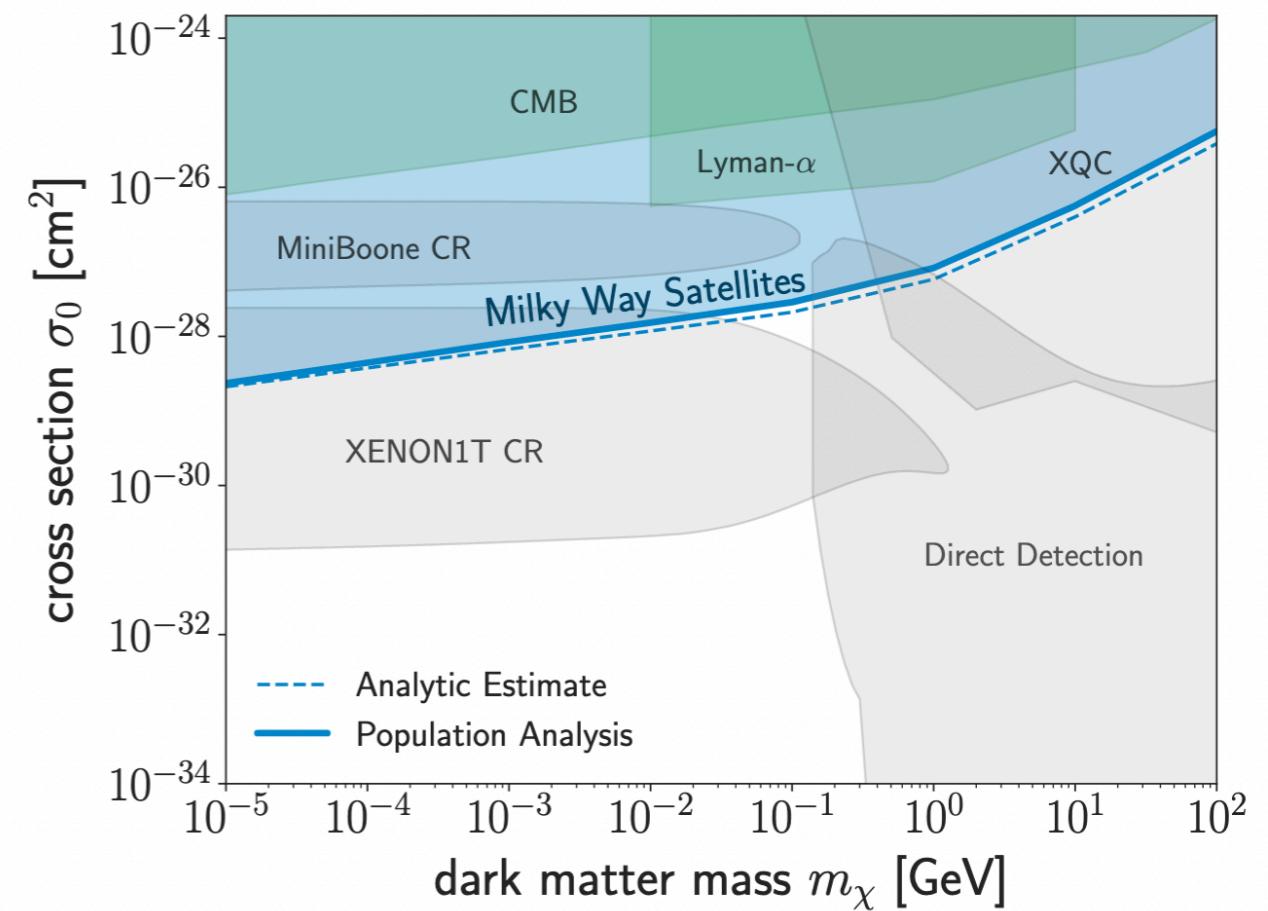
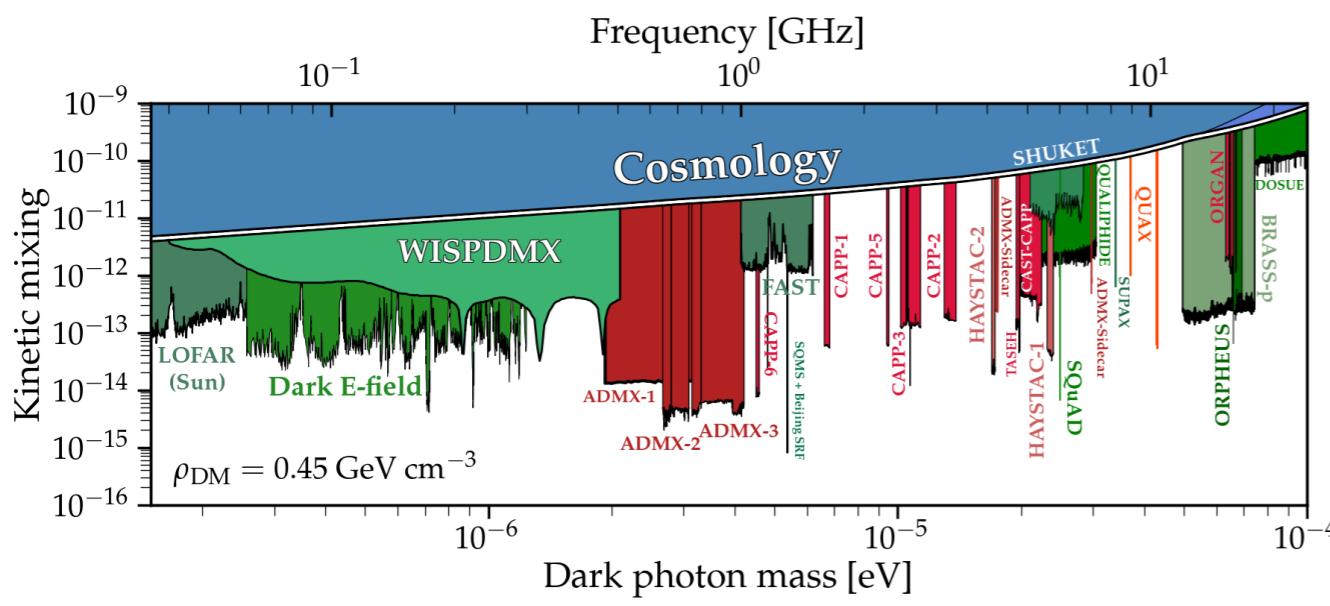
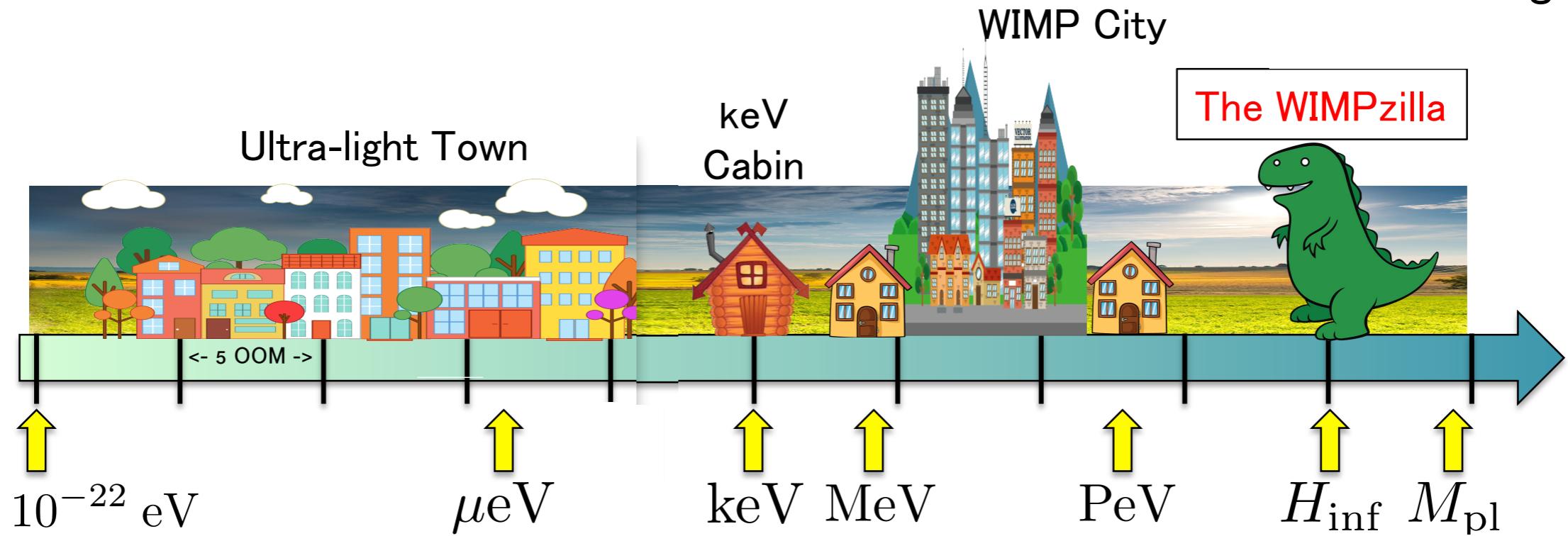
Dark Matter

© A. Long @LWD



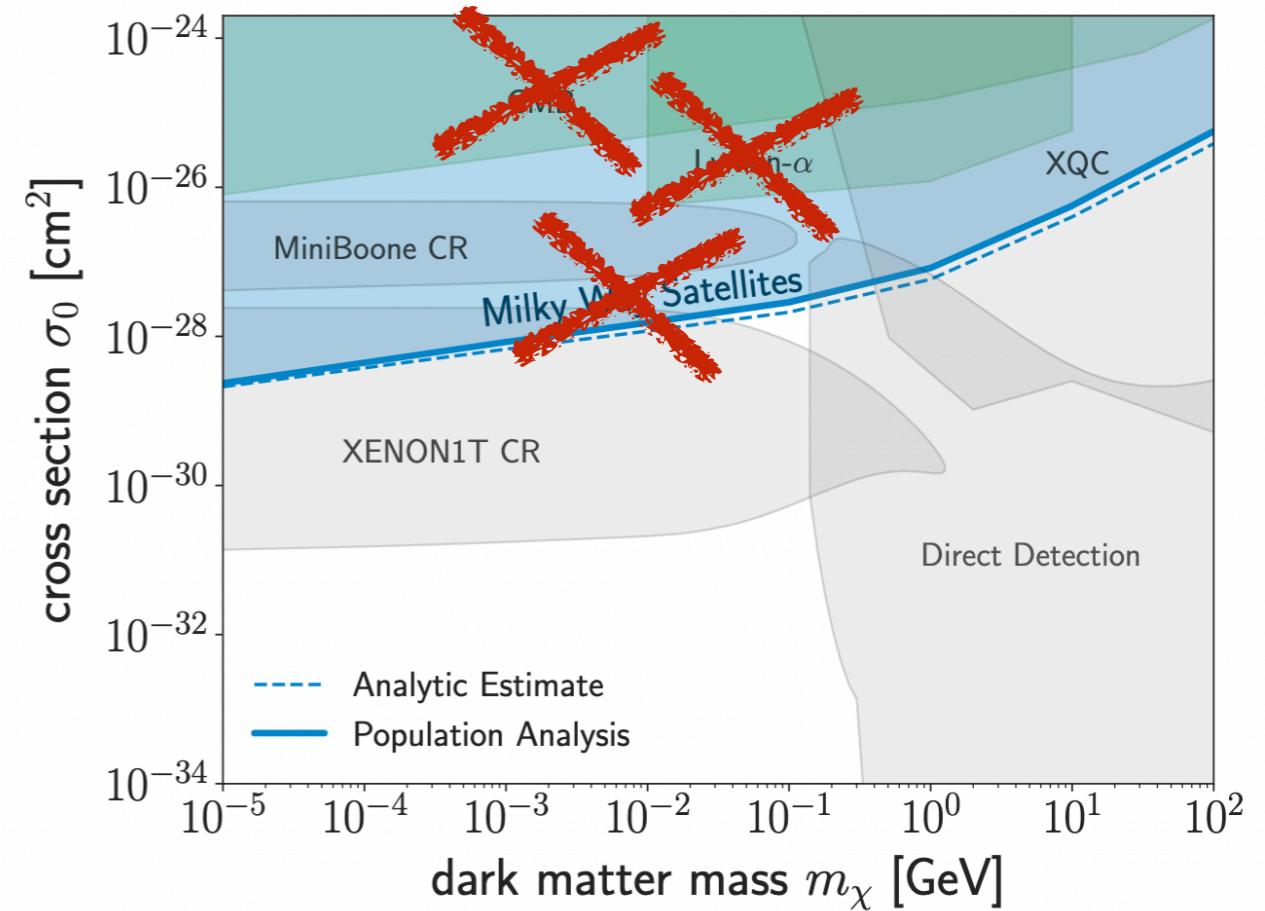
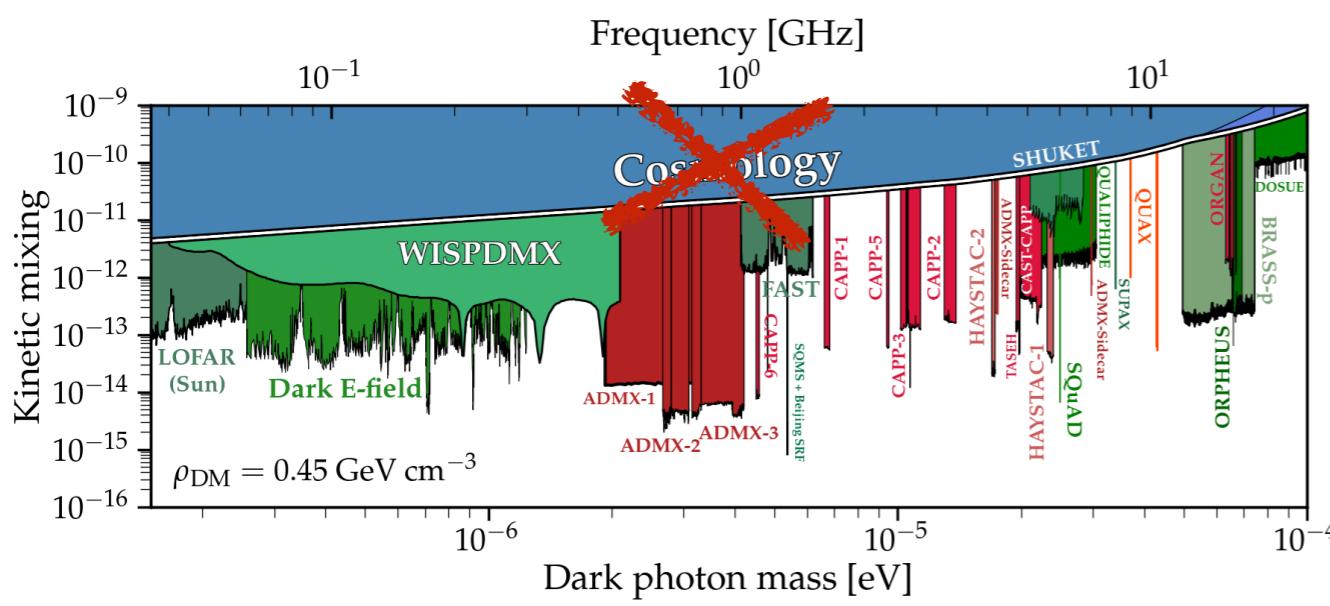
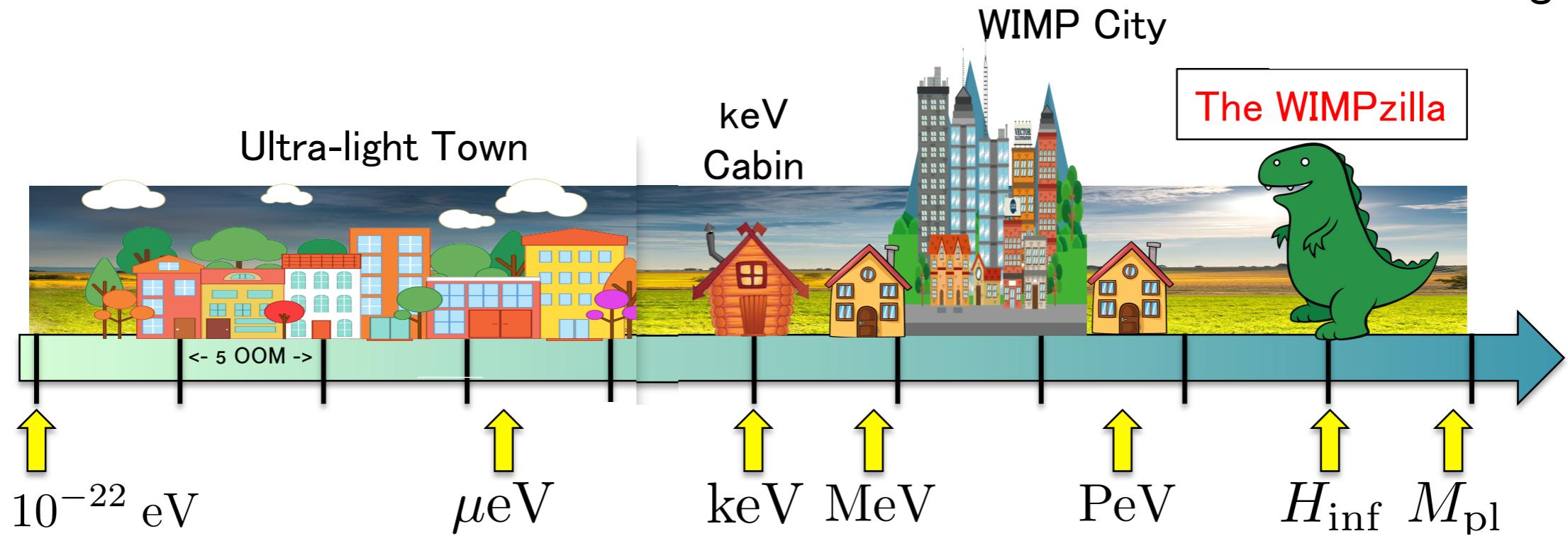
Dark Matter

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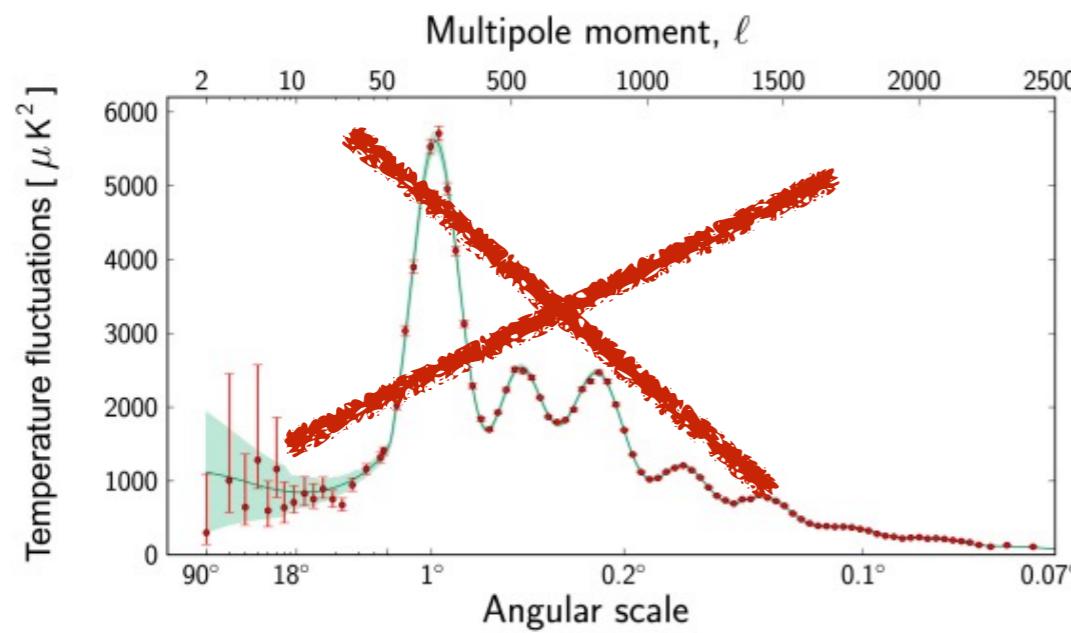
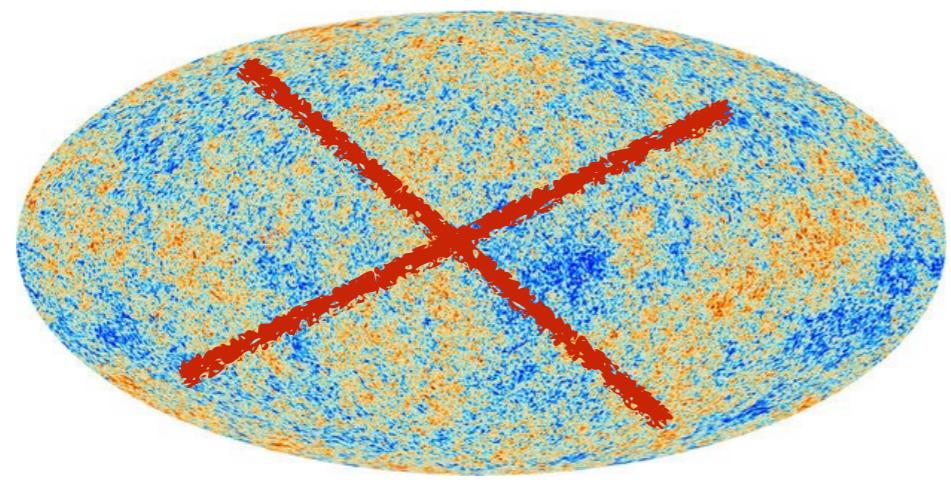
What if there are many DMs ?

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Why did we loose the sensitivity ?

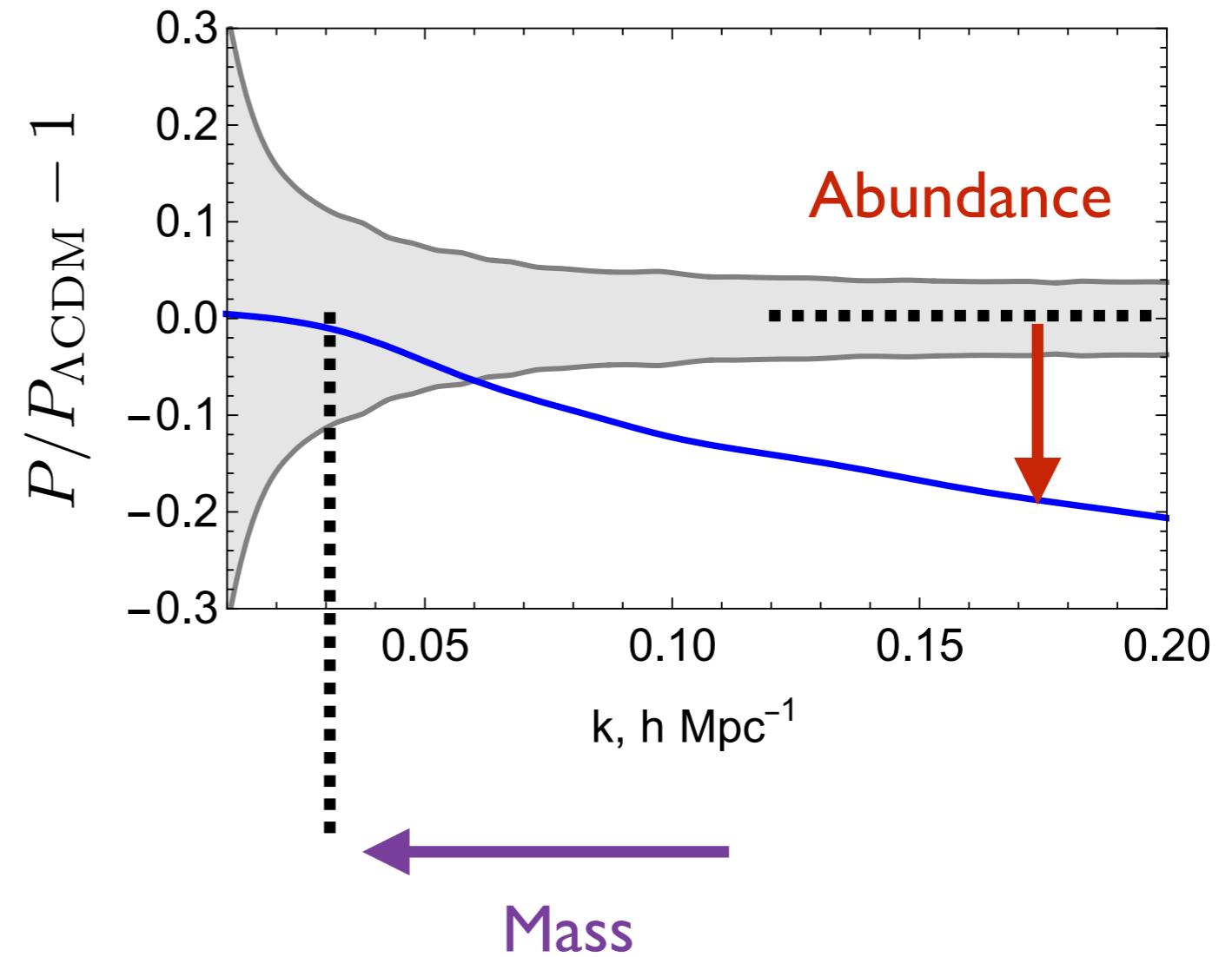
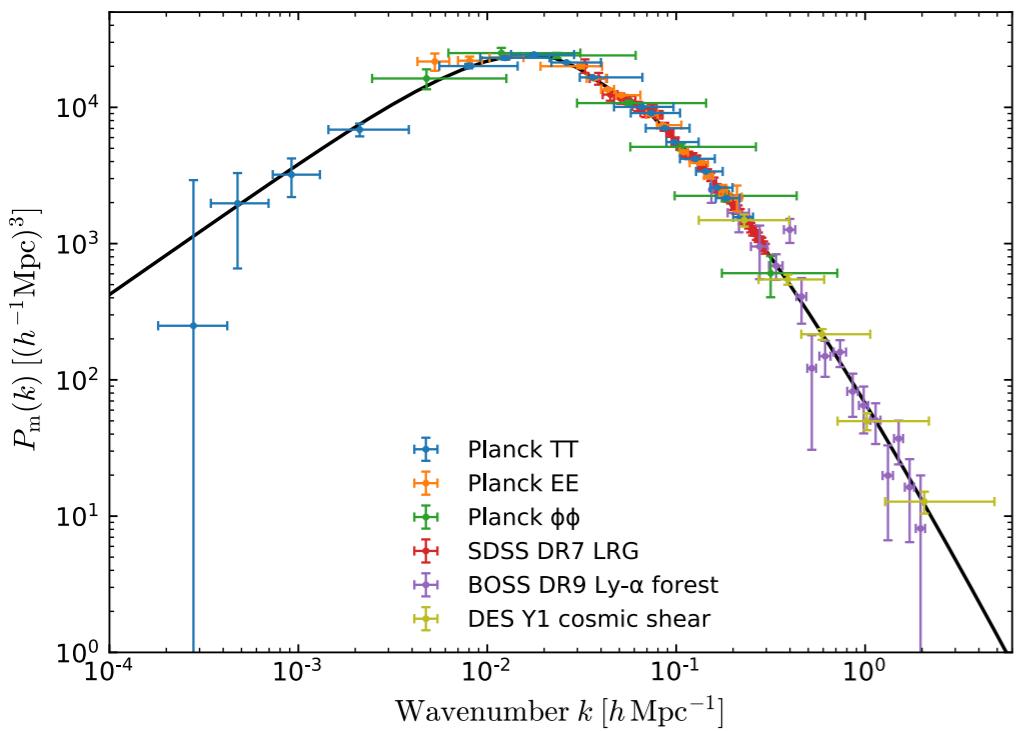
- Hard to sample parameter space with simulations
- CMB isn't very sensitive to DM fluctuations



Weinberg DM decoupling theorem

We can do better with Galaxies!

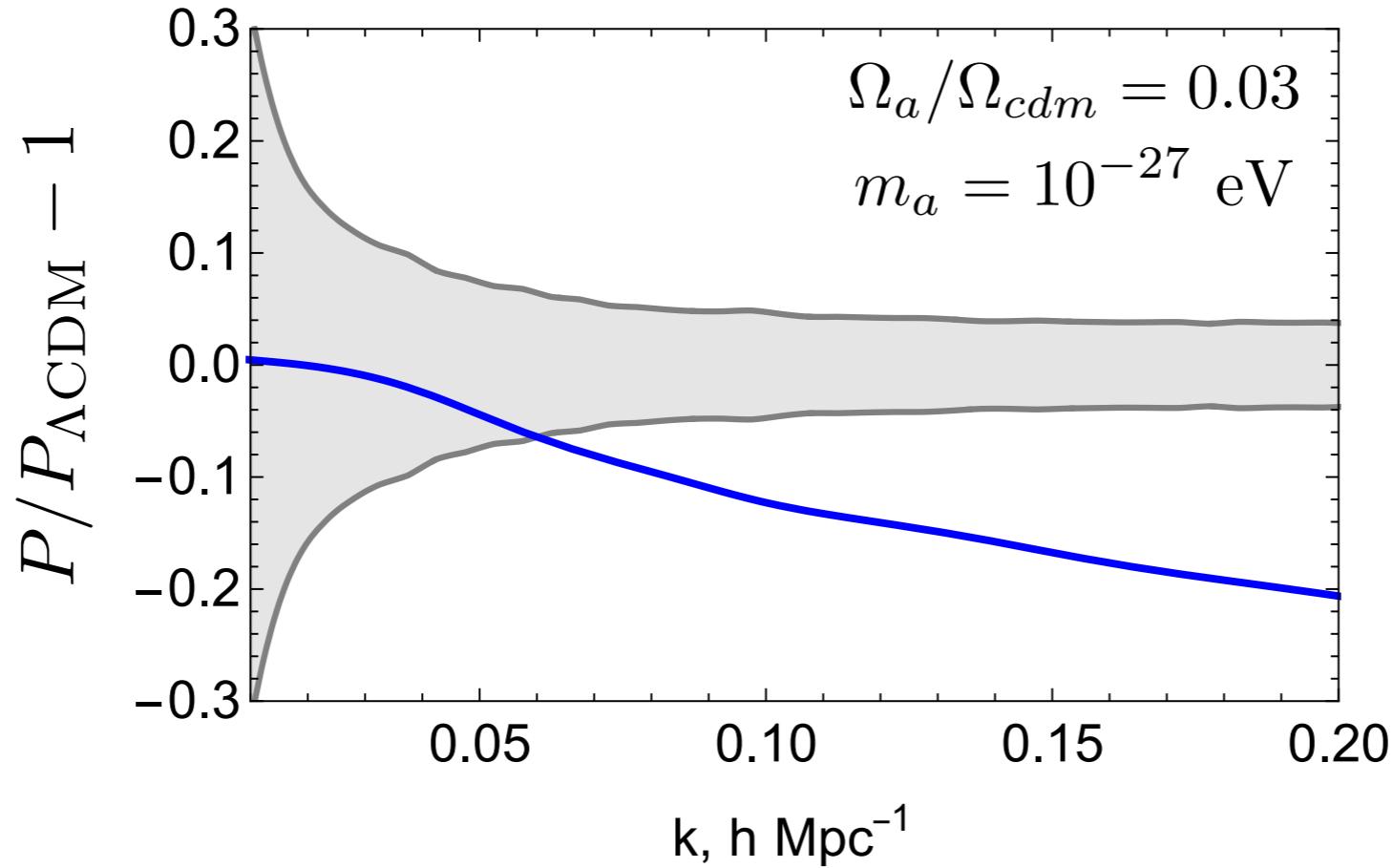
- Imagine two DM components, one is not exactly cold
- ~ there's a Jeans scale beyond which it won't cluster!



Application to axions

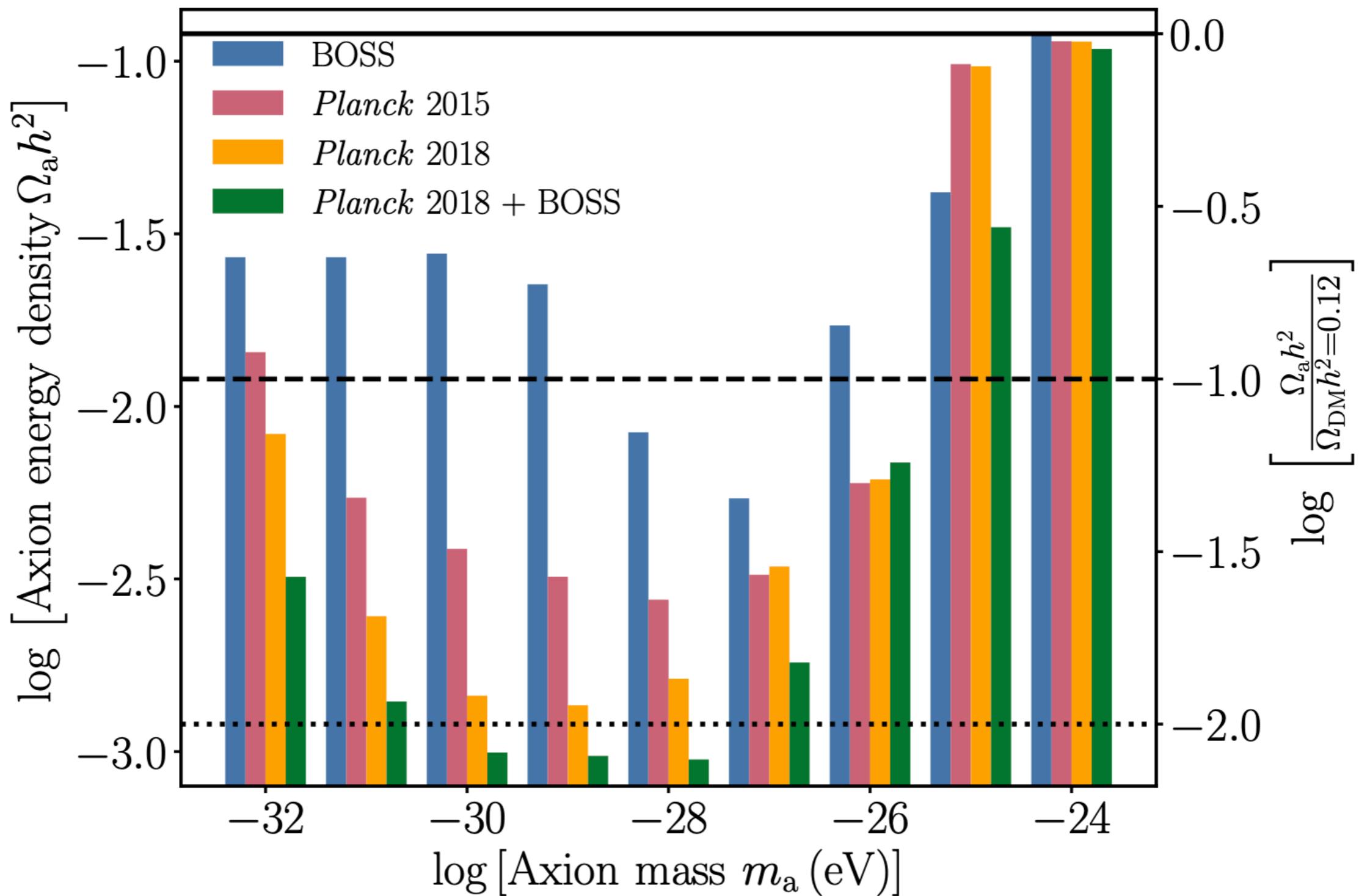
$$\dot{\mathbf{v}} + H\mathbf{v} + \frac{1}{a}(\mathbf{v} \cdot \nabla)\mathbf{v} = -\frac{1}{a}\nabla\Phi - \frac{\hbar^2}{2m^2a^3}\nabla p \quad \text{Hu, Barkana, Gruzinov (2000)}$$

$$p \equiv -\frac{\nabla^2\sqrt{\rho}}{\sqrt{\rho}}$$

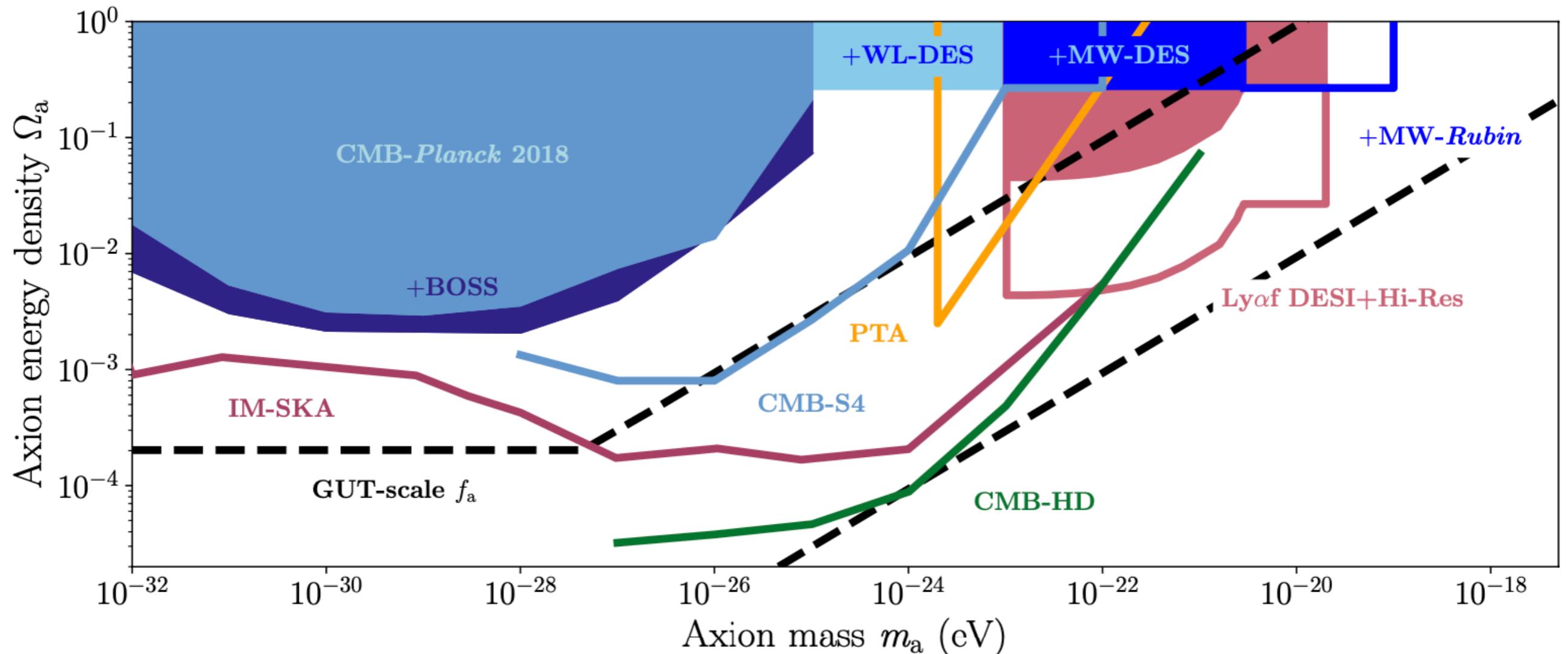


Lague, Rogers, ++(2021) ++ MI (2023)

Axion Dark Matter constraints



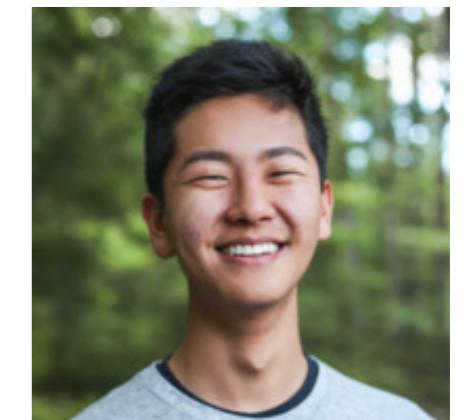
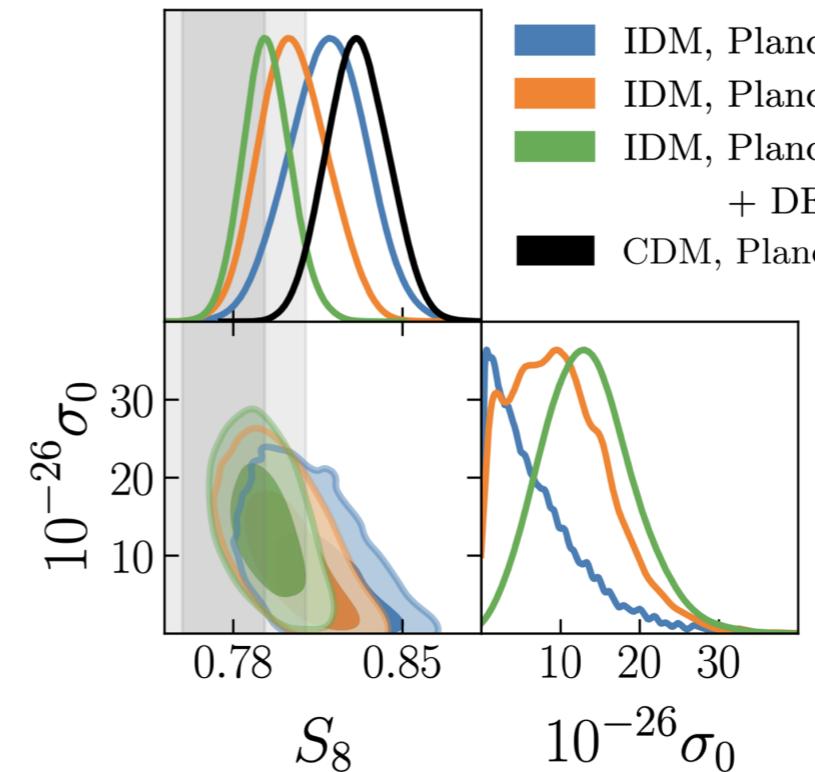
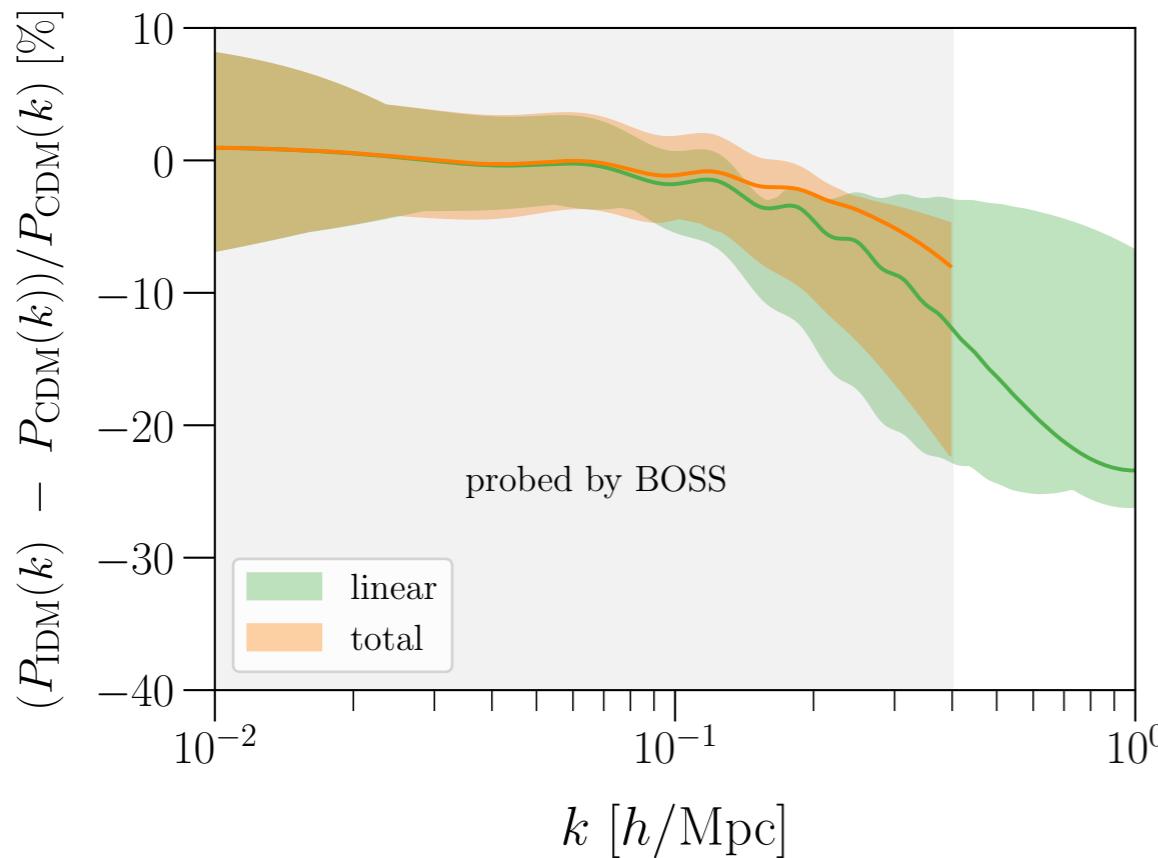
Axion Dark Matter constraints



DM - baryon interactions: apparent evidence ?



motivated by direct detections



[Adam He, MI, Rui, Gluscevic \(2023\)](#)

$\sim 10\%$ of DM $\sim m_\chi \sim 1$ MeV interacts w/ baryons

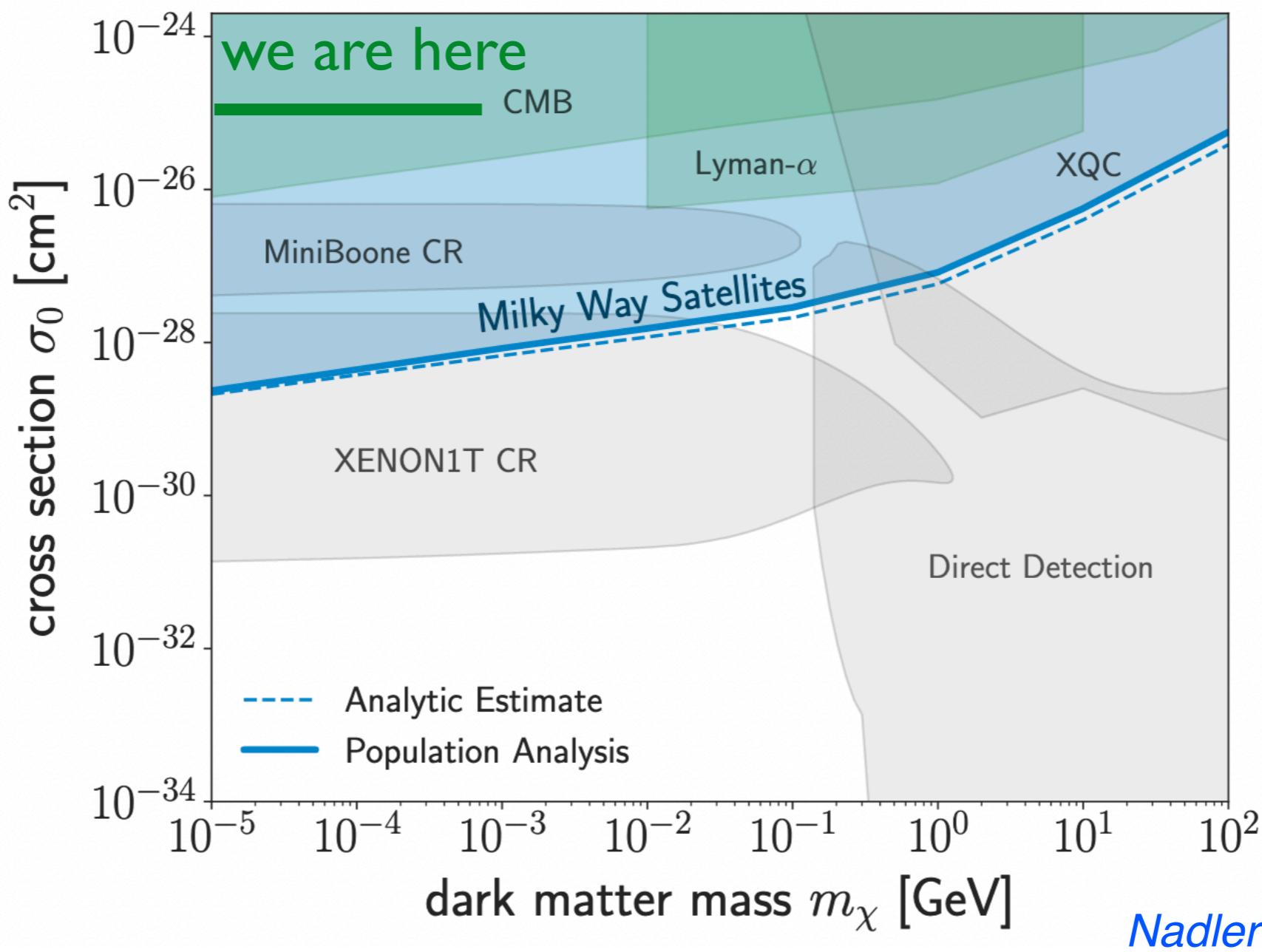
$$\sigma_0 = 1.34^{+0.51}_{-0.67} \times 10^{-25} \text{ cm}^2$$

[Dvorkin, Blum, Kamionkowski ++ \(2014\)](#)

[Gluscevic, Boddy \(2018\)](#)

[Slatyer, Wu \(2018\)](#)

DM - baryon interactions



Assumptions! e.g. 100% DM interacts w/ baryons

Summary



Galaxy Power Spectrum - unique probe of DM



Intriguing results with BOSS



More DM science with future galaxy surveys

Thank you!

