Swanning Around with Chandra Star and Planet Formation in Cygnus OB2

A Chandra Very Large Project www.cygob2.org

J.J. Drake, N. Wright, M. Guarcello, V. Kashyap, T. Aldcroft, E. Flaccomio, Univ. Herts, Armagh Observatory Univ. Exeter, Imperial College, UCL, Palermo Observatory, Univ. Munich, Univ. Bordeaux, CEA, CES, Harvard-Smithsonian, Univ. Liege, NASA GSFC, Smith College, Southampton Univ., Univ. Toledo, Univ. Wyoming See N. Wright poster 133.08

Swa

Star and



nus OB2

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Star and

Нарру Birthday Mario!



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Acronym

Chandra Carina Complex Project – CCCP Chandra Orion Ultra-deep Project – COUP Chandra Cygnus OB2 Survey.....

CYCLES COCOA XXXXX ECCOO TOCCO CHESS CISCO

SUCCO COBS OBII Swan XanOpI CXCI XXXXXX XMuSIC

Cygnus X

Cygnus X

Cygnus OB2 d=1.5 kpc up to 50,000 Msun

Galactic Longitide

Scutumcen

Norma /

210

caurus Arm

240

270



Cy

Durter At

Parseus

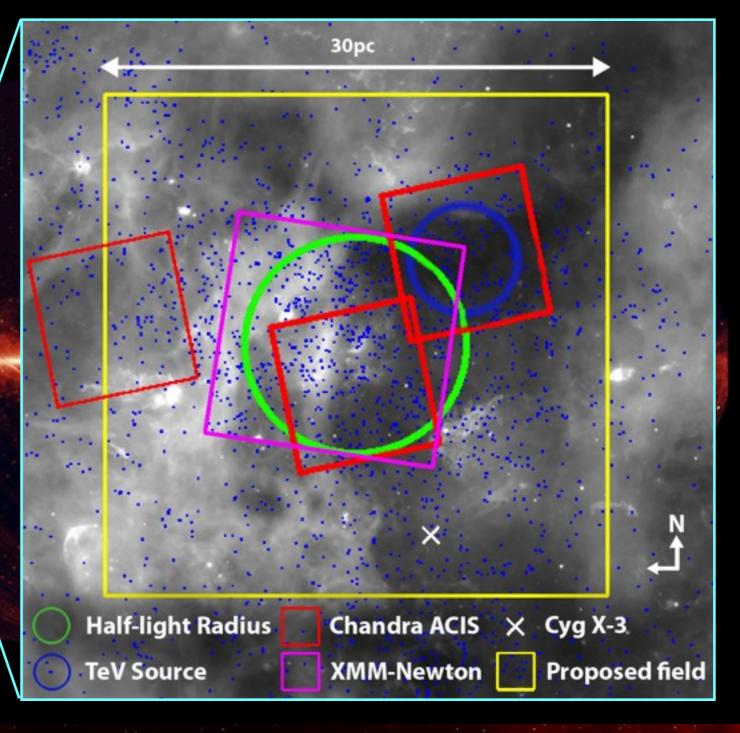
30,000 h

180



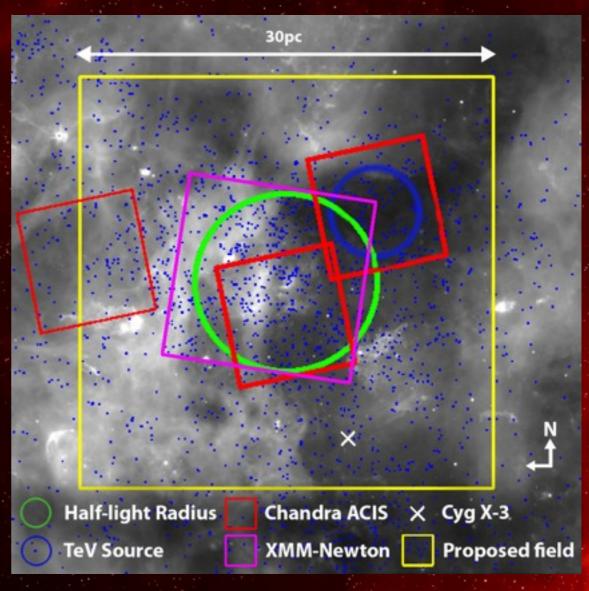
Cygnus X

Cygnus OB2 d=1.5 kpc up to 50,000 Msun Akari 9mm

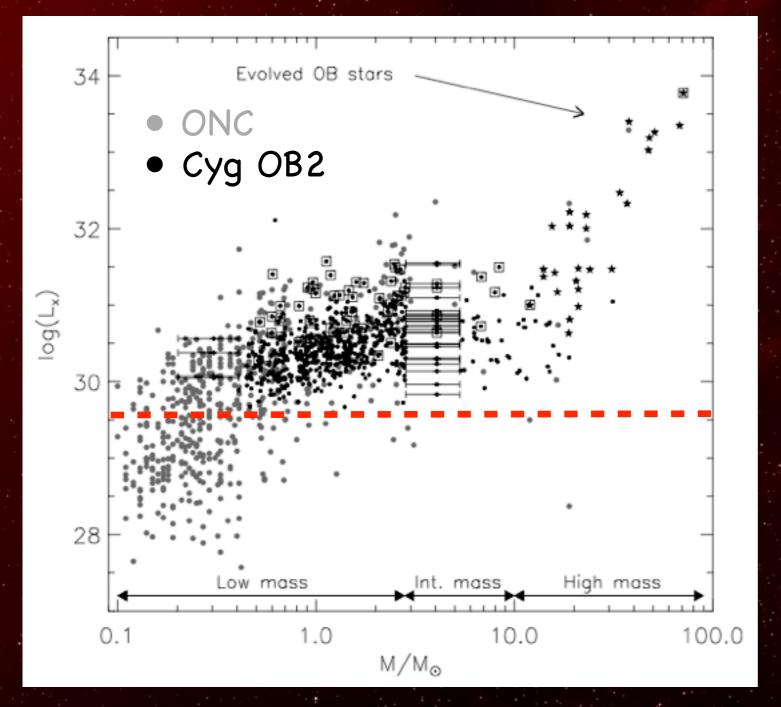


Motivation

- Nearest SF region (1.5 kpc) with pretentions to approaching "supercluster" dimensions
 - two O3 stars, up to 2000 OB stars
 - ø dynamically quite unevolved
- Extensive existing multi-wavelength coverage (inc. Spitzer Cygnus X legacy)
- X-rays isolate young population against back/foreground
 - Can reach < 1 Msun</p>
- MSF, protoplanetary disks, X-ray evolution



Survey Depth



Complete to 1 Msun ~50% complete for
0.5 Msun decent fraction of 0.1-0.5 Msun $OL_X < 10^{29.5} \text{ erg/s}$ 5000-10000 sources predicted

Survey Execution



Survey Execution

a

12,182 X-ray sources 90% of catalog will be Cyg OB2 members 300,000 UKIDDS NIR sources 100,000 optical sources

f: 0.90 86 10941 10942 N...: f: 0.90 f: 0.89

10962 93.2 Net

99.2

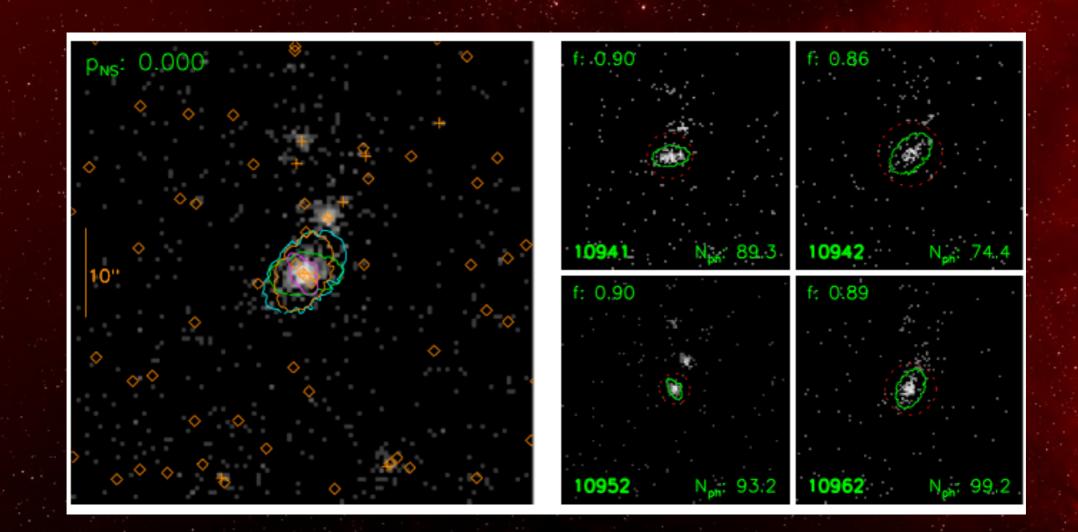
0.000 P_{NS}:

10952

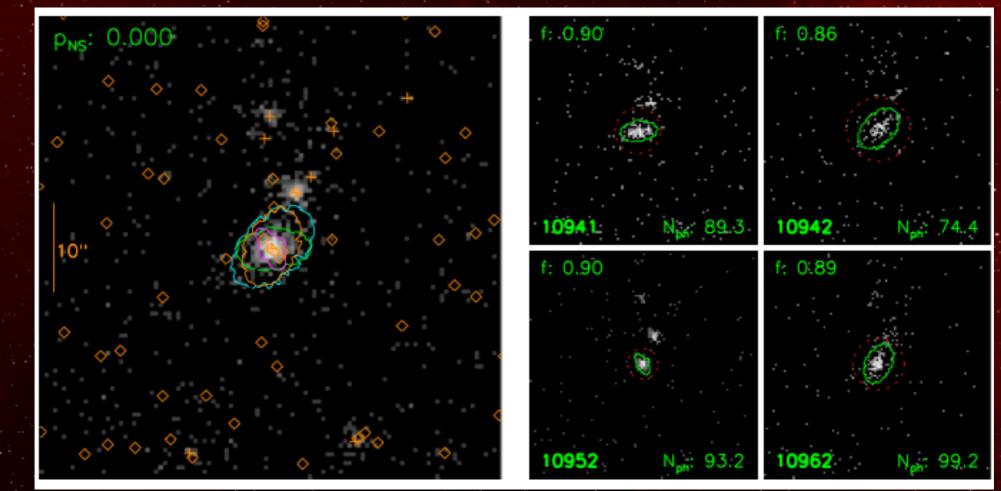
wavdetect, Palermo wave detect, "Vinay Detect"

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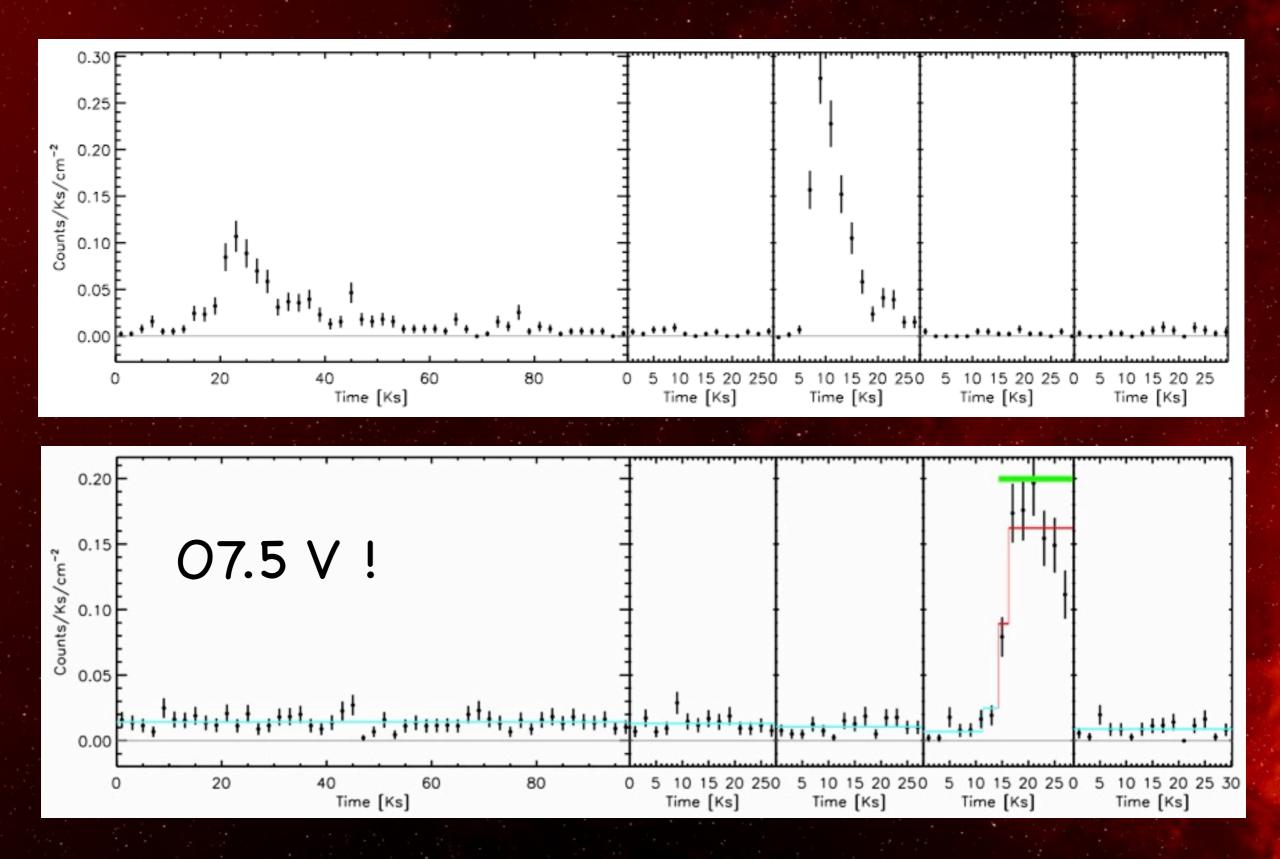
ACIS Extract (Broos et al)



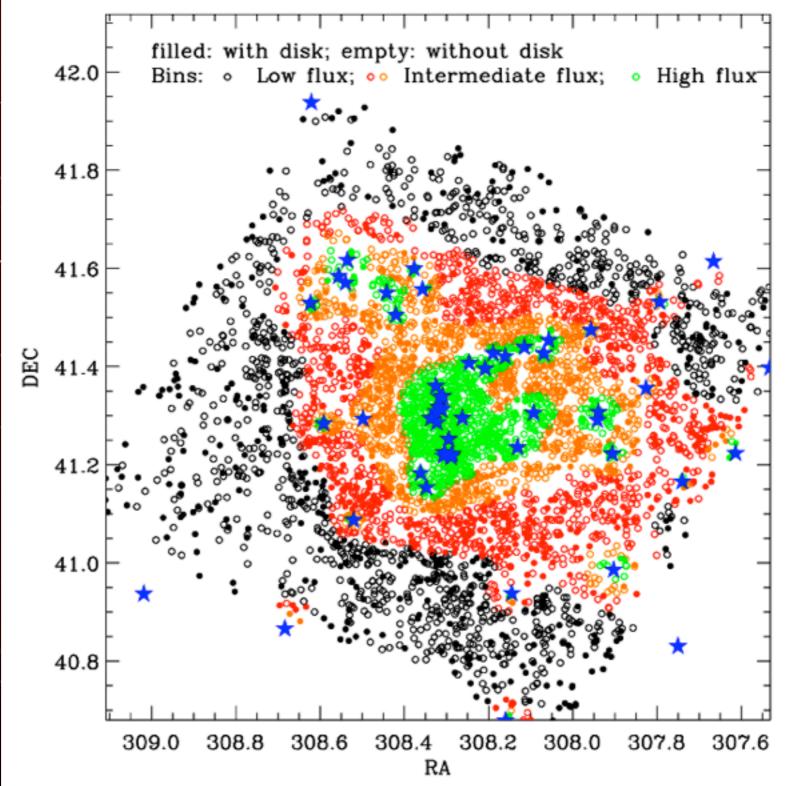
- wavdetect, Palermo wave detect, "Vinay Detect"
- ACIS Extract (Broos et al)
- Postdocs in windowless rooms



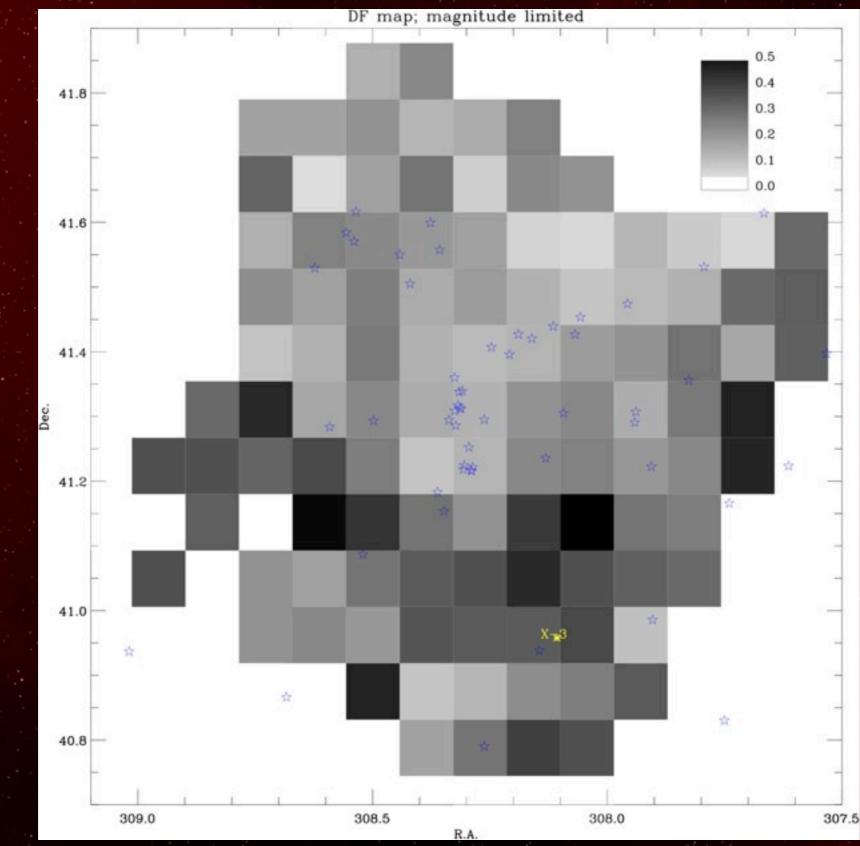
Flares, of course



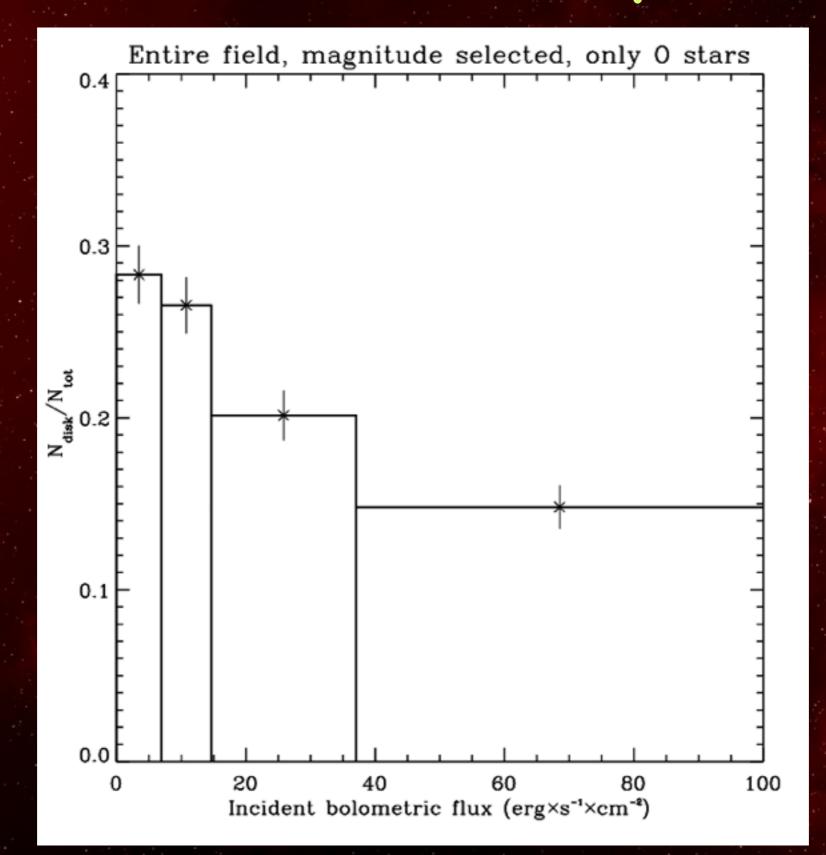
Disk UV Photoevaporation

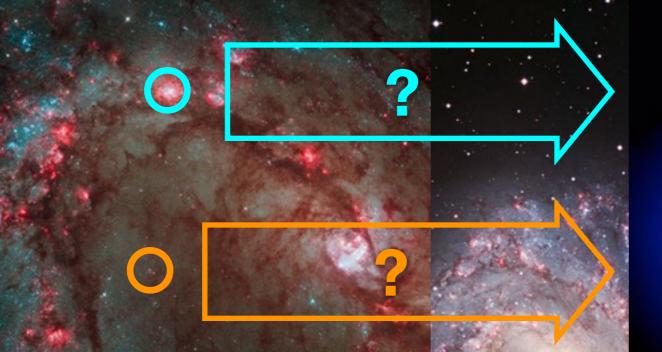


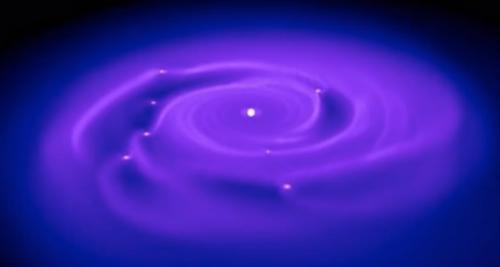
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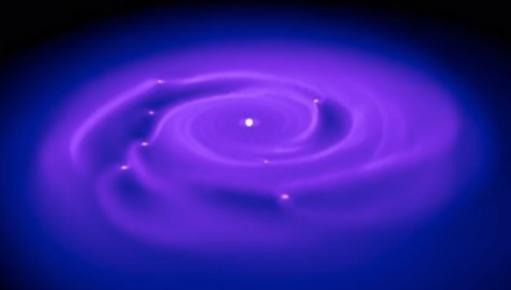
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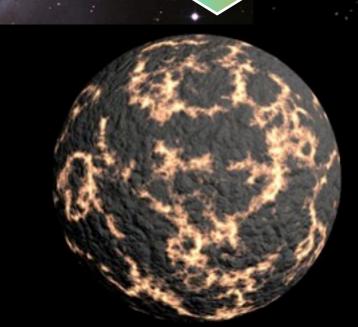




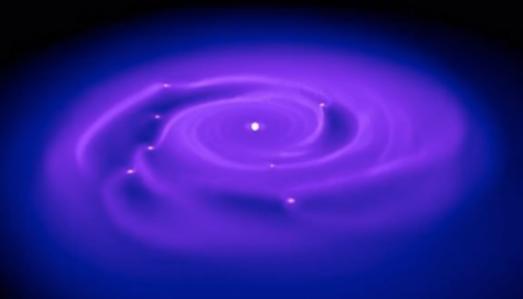






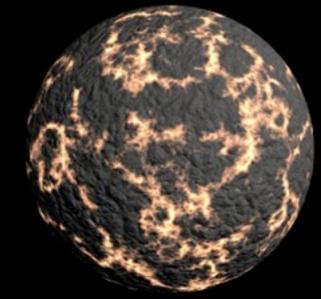














Conclusion

The Chandra Cygnus OB2 survey provides the best evidence to date that OB stars severely affect the evolution of the birthplaces of planets

Survey Science To Come

- Stellar census of Cyg OB2; high mass SF
 - IMF, variations
 - search for mass segregation
 - Oynamics, boundedness, energetics --> supercomputer simulation
 - Search for PWN
- Protoplanetary disk evolution
 - OV photoevaporation
 - comparison with younger Orion
- X-ray properties
 - large OB sample
 - v.large T Tauri sample 1+Myr older than Orion

Organisation

- Survey core at SAO: Drake (PI), Wright (Project Scientist), Aldcroft (technical lead), Guarcello (new postdoc in April), Fruscione
- Science consortium: IPHAS-like organisation
- Main participating groups: SAO, Palermo, various IPHAS
- Collaboration open and welcome
- jdrake@cfa.harvard.edu

Timeline/Logistics

- Observations Jan 29-April (with luck)
- Data processing at SAO: standard reproc, source detection, re-alignment, catalogue;
- central part of field fully processed by mid-March + preliminary source list
- web-based access; <u>www.cygob2.org</u> (see also Nick's wikipedia page)

